

Hi!

Simplifying life with simplified models

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based on:

arXiv: 1312.4175

S. Kraml, U. Laa, A. Lessa, W. Magerl, D. Proschofsky, W. Waltenberger
and

arXiv: 1308.3735

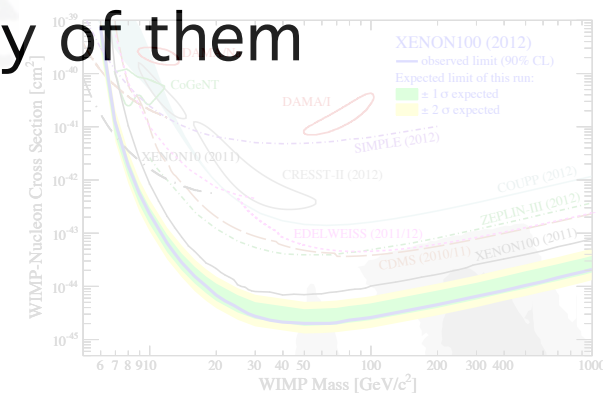
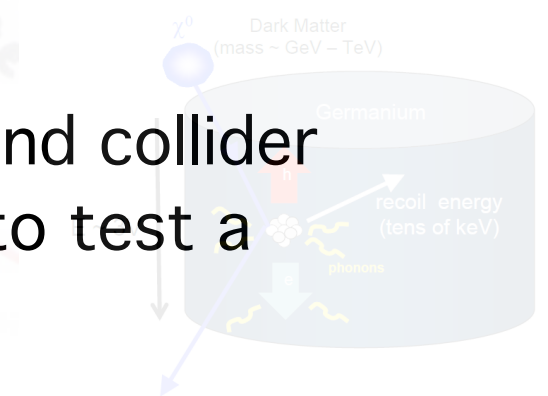
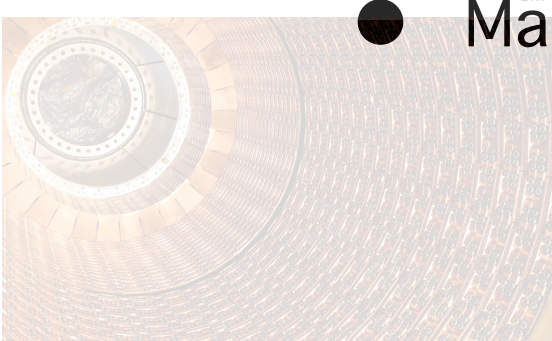
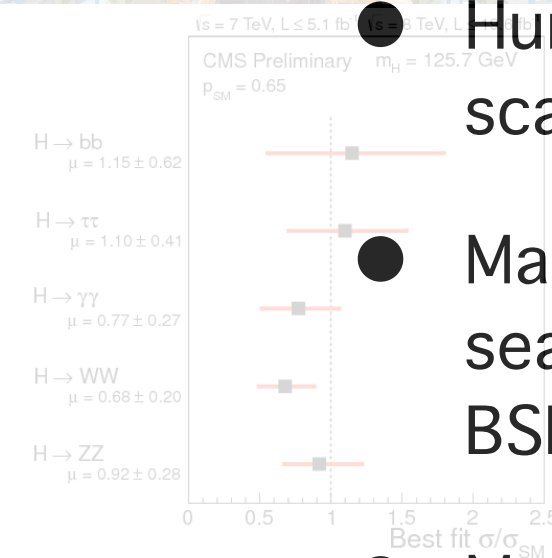
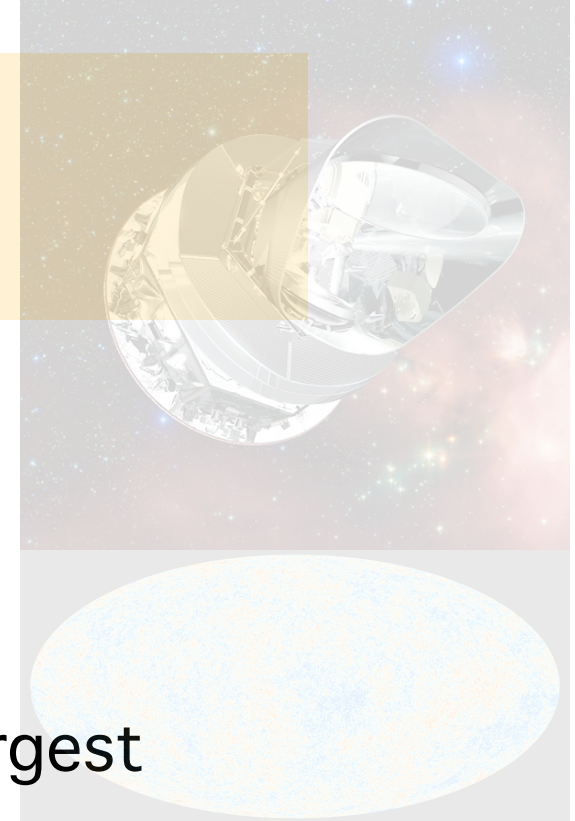
G. Belanger, G. Drieu La Rochelle, B. Dumont, R. Godbole, S. Kraml

The scene

Hunt for BSM physics is strong from the smallest to largest scales

Many new and interesting results from astrophysics and collider searches exist and they must be taken into account to test a BSM theory

Many BSM theories and no conclusive evidence for any of them



The LHC BSM frontier

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... is it?

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: SUSY 2013

ATLAS Preliminary

$$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \quad \sqrt{s} = 7, 8 \text{ TeV}$$

	Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	Reference
Inclusive Searches	MSUGRA/CMSSM	0	2-6 jets	Yes	20.3	\tilde{q}, \tilde{g} 1.7 TeV	$m(\tilde{q})=m(\tilde{g})$ ATLAS-CONF-2013-047
	MSUGRA/CMSSM	1 e, μ	3-6 jets	Yes	20.3	\tilde{g} 1.2 TeV	any $m(\tilde{q})$ ATLAS-CONF-2013-062
	MSUGRA/CMSSM	0	7-10 jets	Yes	20.3	\tilde{g} 1.1 TeV	any $m(\tilde{q})$ 1308.1841
	$\tilde{q}\tilde{q}, \tilde{q}\tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	\tilde{q} 740 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	\tilde{g} 1.3 TeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-047
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow q\tilde{\chi}_1^\pm \rightarrow q\tilde{q}W^\pm \tilde{\chi}_1^0$	1 e, μ	3-6 jets	Yes	20.3	\tilde{g} 1.18 TeV	$m(\tilde{\chi}_1^0)<200 \text{ GeV}, m(\tilde{\chi}^\pm)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$ ATLAS-CONF-2013-062
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow q\tilde{q}(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$	2 e, μ	0-3 jets	-	20.3	\tilde{g} 1.12 TeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-089
	GMSB ($\tilde{\ell}$ NLSP)	2 e, μ	2-4 jets	Yes	4.7	\tilde{g} 1.24 TeV	$\tan\beta<15$ 1208.4688
	GMSB ($\tilde{\ell}$ NLSP)	1-2 τ	0-2 jets	Yes	20.7	\tilde{g} 1.4 TeV	$\tan\beta>18$ ATLAS-CONF-2013-026
	GGM (bino NLSP)	2 γ	-	Yes	4.8	\tilde{g} 1.07 TeV	$m(\tilde{\chi}_1^0)>50 \text{ GeV}$ 1209.0753
	GGM (wino NLSP)	1 $e, \mu + \gamma$	-	Yes	4.8	\tilde{g} 619 GeV	$m(\tilde{\chi}_1^0)>50 \text{ GeV}$ ATLAS-CONF-2012-144
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	4.8	\tilde{g} 900 GeV	$m(\tilde{\chi}_1^0)>220 \text{ GeV}$ 1211.1167
	GGM (higgsino NLSP)	2 $e, \mu (Z)$	0-3 jets	Yes	5.8	\tilde{g} 690 GeV	$m(\tilde{H})>200 \text{ GeV}$ ATLAS-CONF-2012-152
	Gravitino LSP	0	mono-jet	Yes	10.5	$F^{1/2}$ scale 645 GeV	$m(\tilde{g})>10^{-4} \text{ eV}$ ATLAS-CONF-2012-147
3 rd gen. \tilde{g} med.	$\tilde{g}\tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	20.1	\tilde{g} 1.2 TeV	$m(\tilde{\chi}_1^0)<600 \text{ GeV}$ ATLAS-CONF-2013-061
	$\tilde{g}\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0	7-10 jets	Yes	20.3	\tilde{g} 1.1 TeV	$m(\tilde{\chi}_1^0)<350 \text{ GeV}$ 1308.1841
	$\tilde{g}\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^\pm$	0-1 e, μ	3 b	Yes	20.1	\tilde{g} 1.34 TeV	$m(\tilde{\chi}_1^0)<400 \text{ GeV}$ ATLAS-CONF-2013-061
	$\tilde{g}\tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^\pm$	0-1 e, μ	3 b	Yes	20.1	\tilde{g} 1.3 TeV	$m(\tilde{\chi}_1^0)<300 \text{ GeV}$ ATLAS-CONF-2013-061
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	20.1	\tilde{b}_1 100-620 GeV	$m(\tilde{\chi}_1^0)<90 \text{ GeV}$ 1308.2631
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1\tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$	2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{b}_1 275-430 GeV	$m(\tilde{\chi}_1^\pm)=2 m(\tilde{\chi}_1^0)$ ATLAS-CONF-2013-007
	$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	1-2 e, μ	1-2 b	Yes	4.7	\tilde{t}_1 110-167 GeV	$m(\tilde{\chi}_1^0)=55 \text{ GeV}$ 1208.4305, 1209.2102
	$\tilde{t}_1\tilde{t}_1$ (light), $\tilde{t}_1\tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$	2 e, μ	0-2 jets	Yes	20.3	\tilde{t}_1 130-220 GeV	$m(\tilde{\chi}_1^0)=m(\tilde{t}_1)-m(W)-50 \text{ GeV}, m(\tilde{t}_1)<m(\tilde{\chi}_1^\pm)$ ATLAS-CONF-2013-048
	$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$	2 e, μ	2 jets	Yes	20.3	\tilde{t}_1 225-525 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-065
	$\tilde{t}_1\tilde{t}_1$ (medium), $\tilde{t}_1\tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	0	2 b	Yes	20.1	\tilde{t}_1 150-580 GeV	$m(\tilde{\chi}_1^0)<200 \text{ GeV}, m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)=5 \text{ GeV}$ 1308.2631
	$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0$	1 e, μ	1 b	Yes	20.7	\tilde{t}_1 200-610 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-037
	$\tilde{t}_1\tilde{t}_1$ (heavy), $\tilde{t}_1\tilde{t}_1 \rightarrow t\tilde{\chi}_1^\pm$	0	2 b	Yes	20.5	\tilde{t}_1 320-660 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-024
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1\tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet/c-tag	Yes	20.3	\tilde{t}_1 90-200 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)<85 \text{ GeV}$ ATLAS-CONF-2013-068
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 $e, \mu (Z)$	1 b	Yes	20.7	\tilde{t}_1 500 GeV	$m(\tilde{\chi}_1^0)>150 \text{ GeV}$ ATLAS-CONF-2013-025
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2\tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 $e, \mu (Z)$	1 b	Yes	20.7	\tilde{t}_2 271-520 GeV	$m(\tilde{t}_1)=m(\tilde{\chi}_1^0)+180 \text{ GeV}$ ATLAS-CONF-2013-025
	$\tilde{t}_1\tilde{t}_2, \tilde{t}_1\tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 $e, \mu (Z)$	1 b	Yes	20.7	\tilde{t}_2 271-520 GeV	$m(\tilde{t}_1)=m(\tilde{\chi}_1^0)+180 \text{ GeV}$ ATLAS-CONF-2013-025
EW direct	$\tilde{\ell}_L\tilde{\ell}_L, \tilde{\ell}_L\tilde{\ell}_L \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ	0	Yes	20.3	$\tilde{\ell}$ 85-315 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ ATLAS-CONF-2013-049
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm\tilde{\chi}_1^\pm \rightarrow \tilde{\ell}\nu(\ell\tilde{\nu})$	2 e, μ	0	Yes	20.3	$\tilde{\chi}_1^\pm$ 125-450 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$ ATLAS-CONF-2013-049
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm\tilde{\chi}_1^\pm \rightarrow \tilde{\tau}\nu(\tau\tilde{\nu})$	2 τ	-	Yes	20.7	$\tilde{\chi}_1^\pm$ 180-330 GeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}, m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$ ATLAS-CONF-2013-028
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_L\nu\tilde{\ell}_L(\ell\tilde{\nu}), \ell\tilde{\nu}\tilde{\ell}_L(\ell\tilde{\nu})$	3 e, μ	0	Yes	20.7	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ 600 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_1^\pm)+m(\tilde{\chi}_1^0))$ ATLAS-CONF-2013-035
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$	3 e, μ	0	Yes	20.7	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ 315 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0$, sleptons decoupled ATLAS-CONF-2013-035
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0$	1 e, μ	2 b	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$ 285 GeV	$m(\tilde{\chi}_1^\pm)=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0$, sleptons decoupled ATLAS-CONF-2013-093
Long-lived particles	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	20.3	$\tilde{\chi}_1^\pm$ 270 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0)=160 \text{ MeV}, \tau(\tilde{\chi}_1^\pm)=0.2 \text{ ns}$ ATLAS-CONF-2013-069
	Stable, stopped \tilde{g} R-hadron	0	1-5 jets	Yes	22.9	\tilde{g} 832 GeV	$m(\tilde{\chi}_1^0)=100 \text{ GeV}, 10 \mu\text{s}<\tau(\tilde{g})<1000 \text{ s}$ ATLAS-CONF-2013-057
	GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$	1-2 μ	-	-	15.9	$\tilde{\chi}_1^0$ 475 GeV	$10<\tan\beta<50$ ATLAS-CONF-2013-058
	GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$	2 γ	-	Yes	4.7	$\tilde{\chi}_1^0$ 230 GeV	$0.4<\tau(\tilde{\chi}_1^0)<2 \text{ ns}$ 1304.6310
RPV	$\tilde{q}\tilde{q}, \tilde{\chi}_1^0 \rightarrow q\tilde{q}\mu$ (RPV)	1 μ , displ. vtx	-	-	20.3	\tilde{q} 1.0 TeV	$1.5<c\tau<156 \text{ mm}, \text{BR}(\mu)=1, m(\tilde{\chi}_1^0)=108 \text{ GeV}$ ATLAS-CONF-2013-092
	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e + \mu$	2 e, μ	-	-	4.6	$\tilde{\nu}_\tau$ 1.61 TeV	$\lambda'_{311}=0.10, \lambda_{132}=0.05$ 1212.1272
	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e(\mu) + \tau$	1 $e, \mu + \tau$	-	-	4.6	$\tilde{\nu}_\tau$ 1.1 TeV	$\lambda'_{311}=0.10, \lambda_{1(2)33}=0.05$ 1212.1272
	Bilinear RPV CMSSM	1 e, μ	7 jets	Yes	4.7	\tilde{q}, \tilde{g} 1.2 TeV	$m(\tilde{q})=m(\tilde{g}), c\tau_{\text{LSP}}<1 \text{ mm}$ ATLAS-CONF-2012-140
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm\tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^\pm\tilde{\chi}_1^\pm \rightarrow ee\tilde{\nu}_\mu, e\mu\tilde{\nu}_e$	4 e, μ	-	Yes	20.7	$\tilde{\chi}_1^\pm$ 760 GeV	$m(\tilde{\chi}_1^0)>300 \text{ GeV}, \lambda_{121}>0$ ATLAS-CONF-2013-036
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm\tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^\pm\tilde{\chi}_1^\pm \rightarrow \tau\tau\tilde{\nu}_e, e\tau\tilde{\nu}_\tau$	3 $e, \mu + \tau$	-	Yes	20.7	$\tilde{\chi}_1^\pm$ 350 GeV	$m(\tilde{\chi}_1^0)>80 \text{ GeV}, \lambda_{133}>0$ ATLAS-CONF-2013-036
	$\tilde{g}\tilde{g} \rightarrow q\tilde{q}q$	0	6-7 jets	-	20.3	\tilde{g} 916 GeV	$\text{BR}(t)=\text{BR}(b)=\text{BR}(c)=0\%$ ATLAS-CONF-2013-091
Other	$\tilde{g}\tilde{g} \rightarrow \tilde{t}_1\tilde{t}_1, \tilde{t}_1\tilde{t}_1 \rightarrow b\tilde{s}$	2 e, μ (SS)	0-3 b	Yes	20.7	\tilde{g} 880 GeV	ATLAS-CONF-2013-007
	Scalar gluon pair, sgluon $\rightarrow q\tilde{q}$	0	4 jets	-	4.6	sgluon 100-287 GeV	incl. limit from 1110.2693 1210.4826
	Scalar gluon pair, sgluon $\rightarrow t\tilde{t}$	2 e, μ (SS)	1 b	Yes	14.3	sgluon 800 GeV	ATLAS-CONF-2013-051
Other	WIMP interaction (D5, Dirac χ)	0	mono-jet	Yes	10.5	M^* scale 704 GeV	$m(\chi)<80 \text{ GeV}$, limit of $<687 \text{ GeV}$ for D8 ATLAS-CONF-2012-147

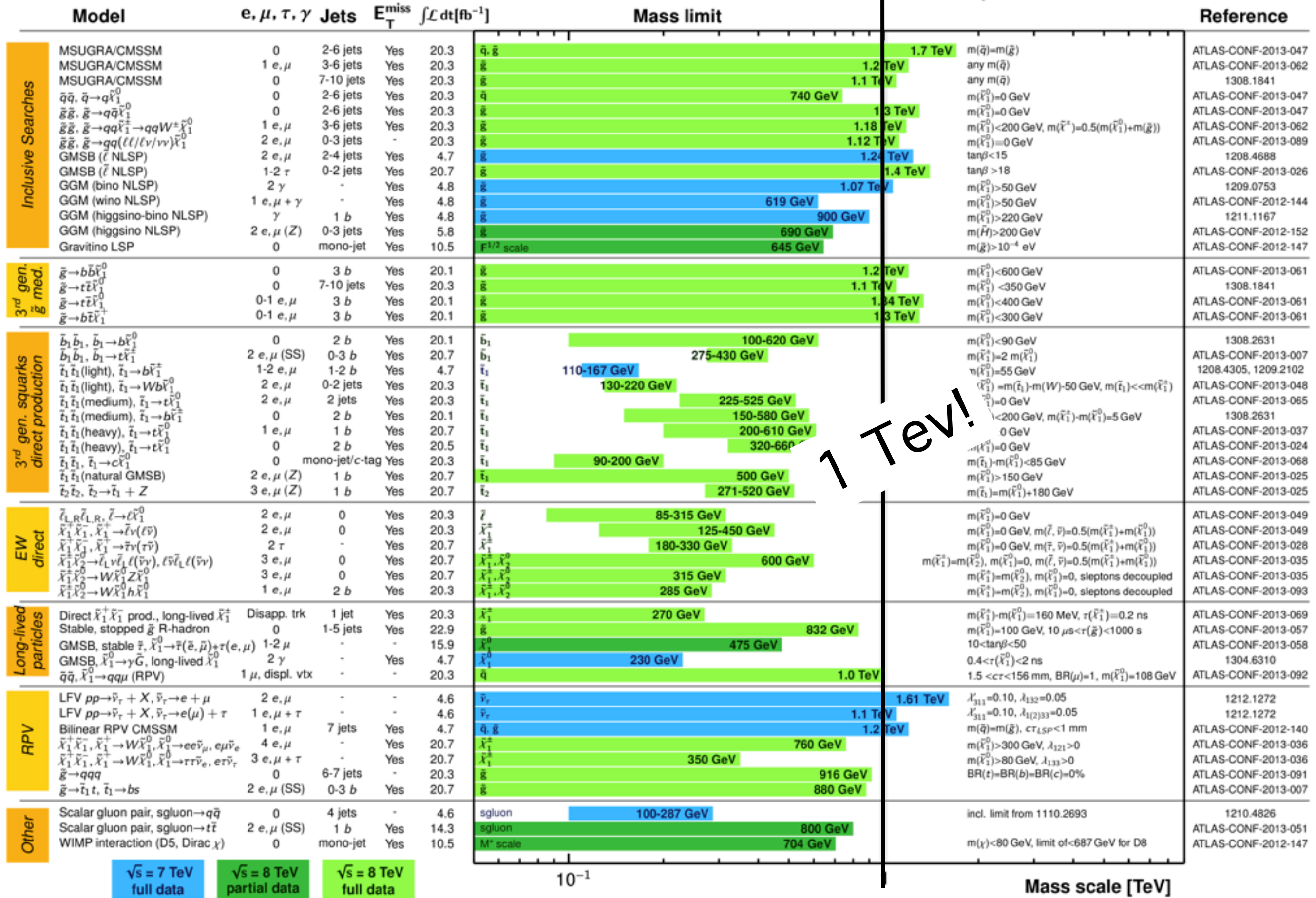
*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1σ theoretical signal cross section uncertainty.

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: SUSY 2013

ATLAS Preliminary

$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$ $\sqrt{s} = 7, 8 \text{ TeV}$



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The LHC frontier

- Theoretical model development is influenced by conclusions at 8 TeV, e.g. GUT scale SUSY models like cMSSM are pushed to higher scales and we are thinking of more non-minimal models
- The strategies for 13 TeV results depend on the conclusions at 8 TeV
- It is necessary to interpret the results in the most generic fashion and test as many models as possible

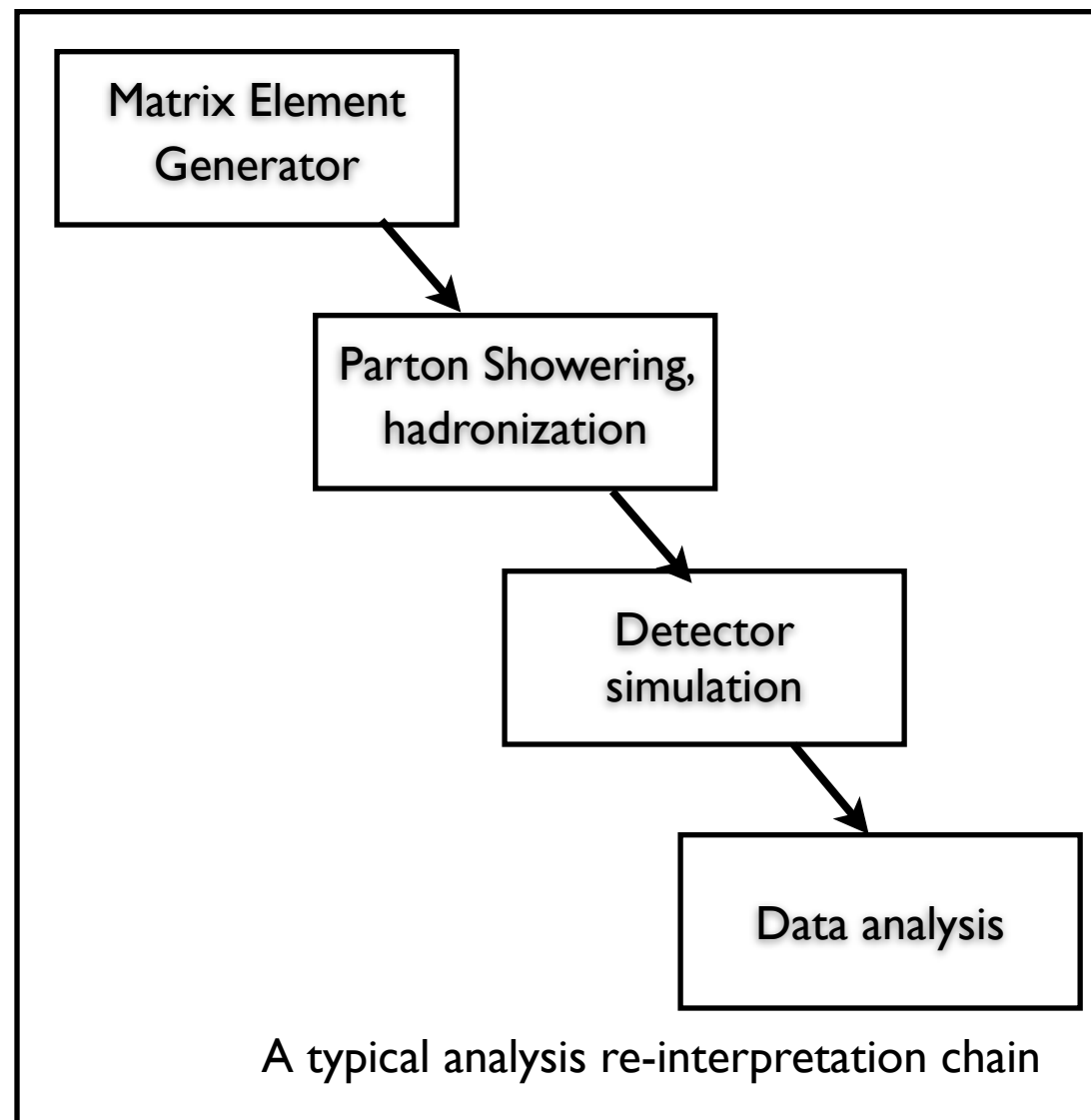
Analyses statistics (SUSY searches only)

	8 TeV (Any luminosity)	8 TeV (20 fb ⁻¹)
ATLAS	39	23
CMS	28	20

Huge number of searches
Easier said than done!

- A way to test our favorite BSM model against LHC results should exist

Traditional approach

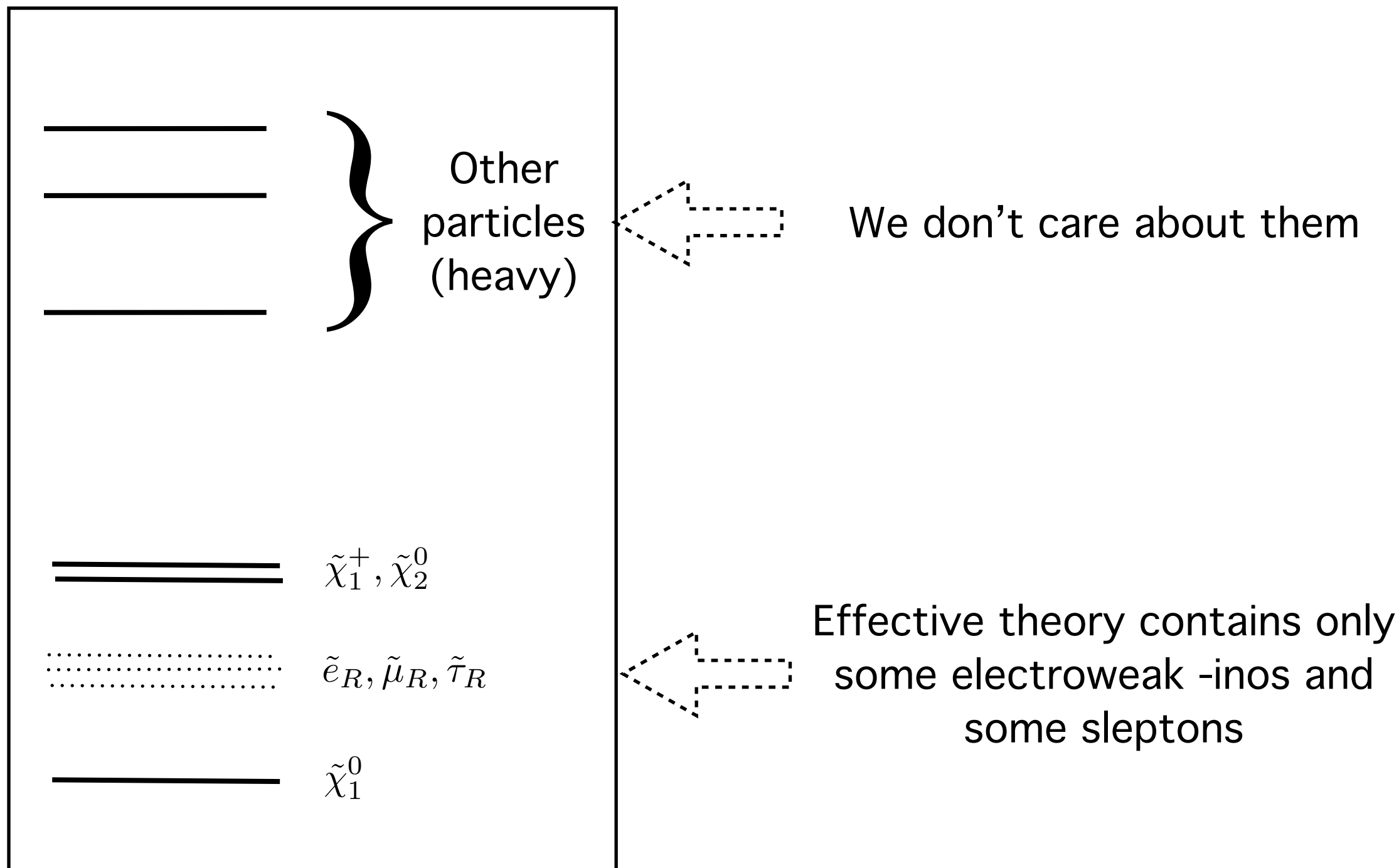


- Interpretation of LHC searches are model dependent
 - Model dependence comes while converting the number of events observed to a limit on particle masses
 - For a more generic case:
 1. Re-interpret the results yourself
 2. Use simplified model spectra results
-
- Re-interpreting the results yourself involves re-implementing the analysis, requires expertise, large computing power, time consuming
 - We stick to simplified models results

What is an SMS result?

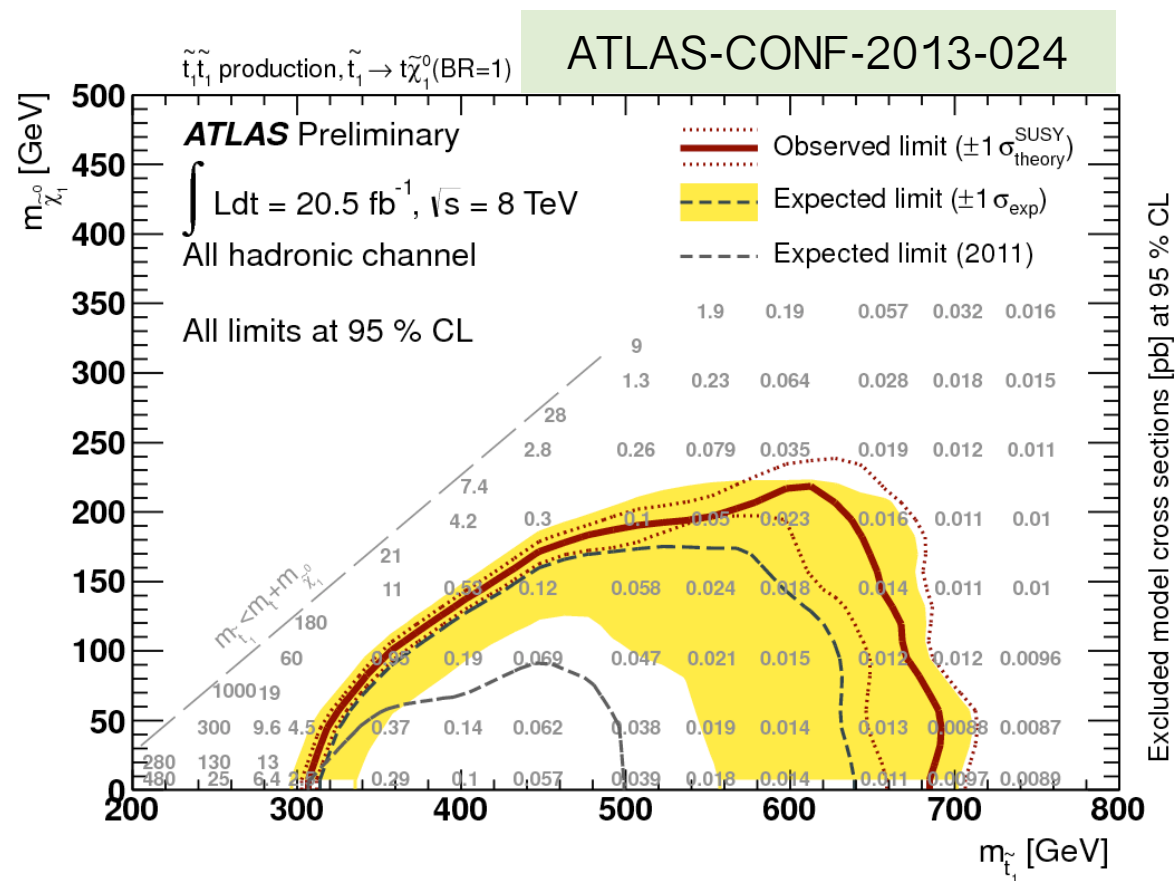
- SMS are an effective-Lagrangian description of BSM involving a limited set of new particles.

What is an SMS result?



What is an SMS result?

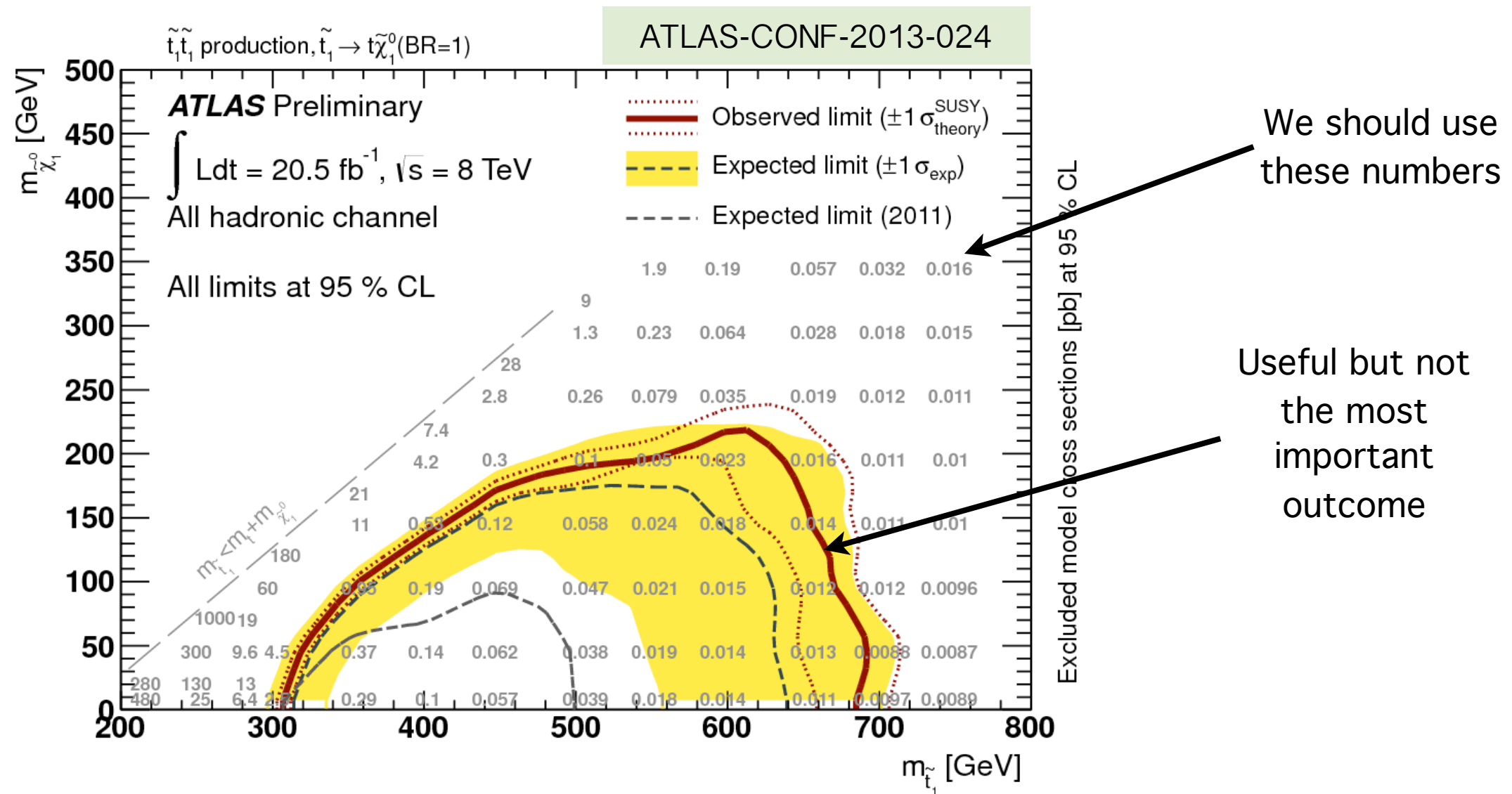
Note: the grid numbers on the plot are more important than the exclusion lines



- Every SMS interpretation is based on a set of assumptions and is applicable for specific topologies e.g. $t\bar{t}$ + MET

- A generic point in e.g. SUSY parameter space contains many topologies and is sensitive to more than one SMS interpretation e.g. $t\bar{t}$ + MET, $b\bar{b}$ + MET

How to read an SMS result



- 95% CL UL is the unfolded maximum amount of cross-section allowed for a specific decay chain and a mass combination

Is $\sigma_{\text{XBR}}(\text{ttbar} + \text{MET}, \text{Mother mass}, \text{LSP mass})$ of your model $>$ the number on the plot? -- Yes, point excluded; No, point allowed

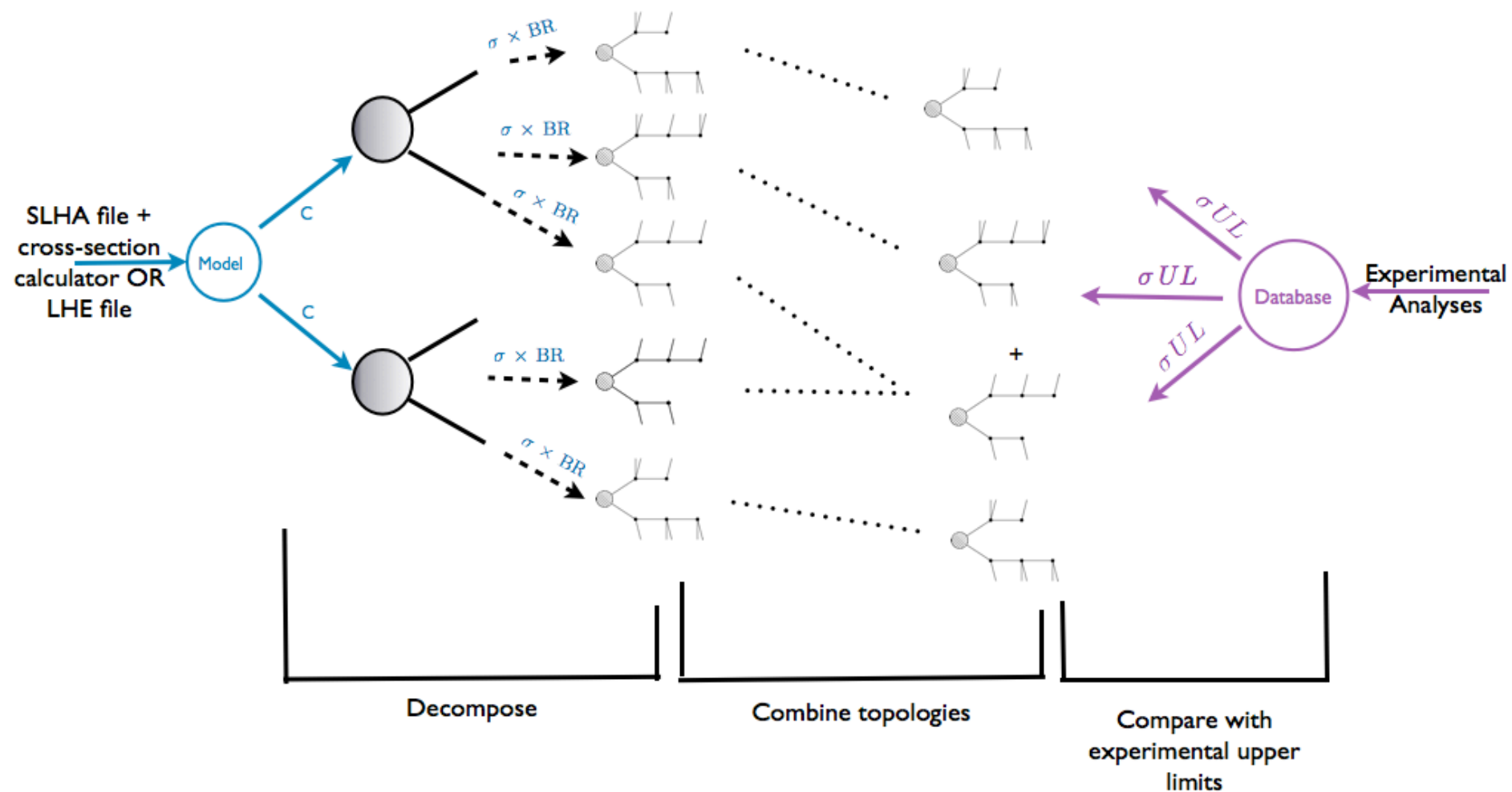
Can we have a centralized database of all the SMS results to check a given SUSY point in parameter space by decomposing it into SMS topologies?

Central concept of



SModelS framework

- It assumes, for most experimental searches, the BSM model can be approximated by a sum over effective simplified models



- Current implementation assumes R-parity is conserved

Given
Spectra

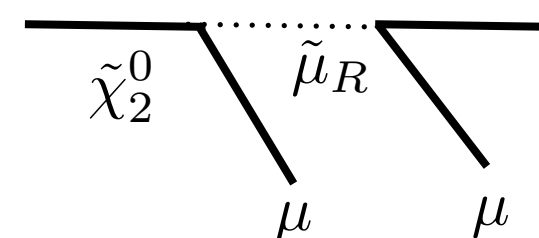
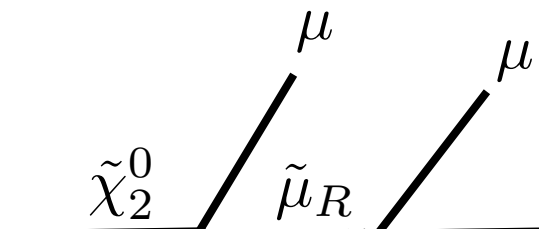
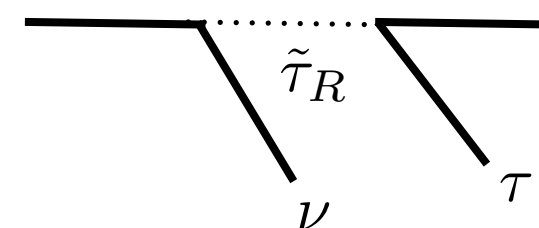
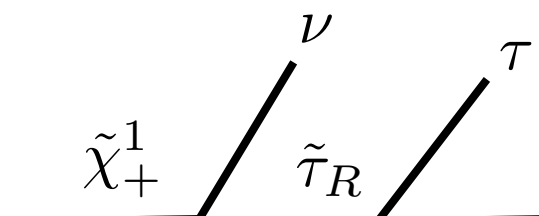
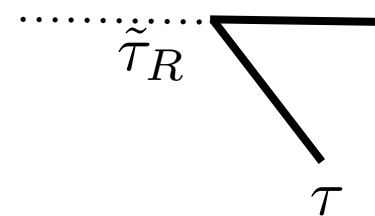
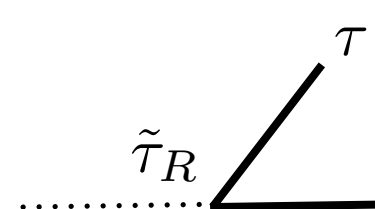
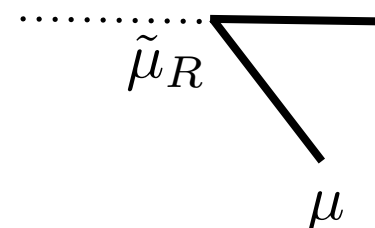
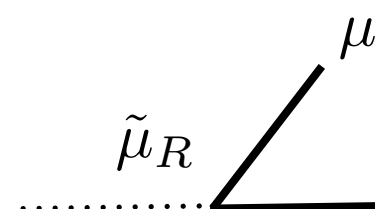
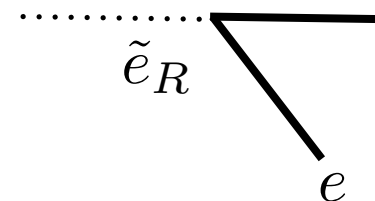
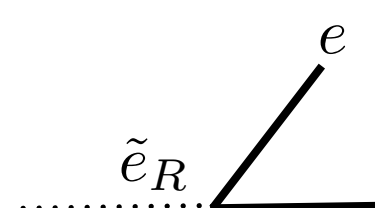
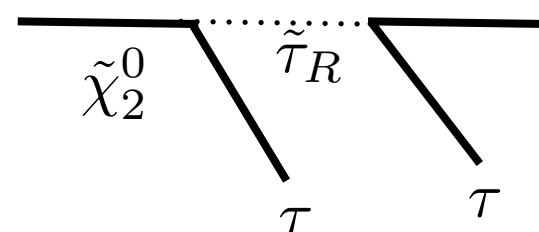
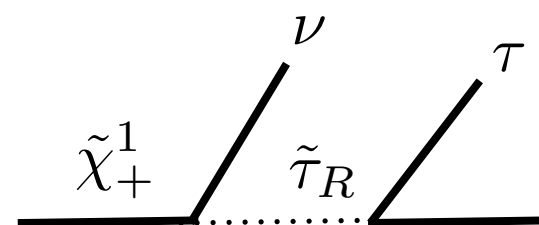
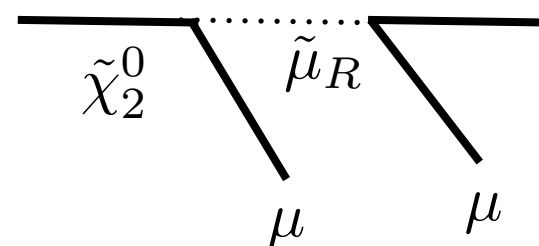
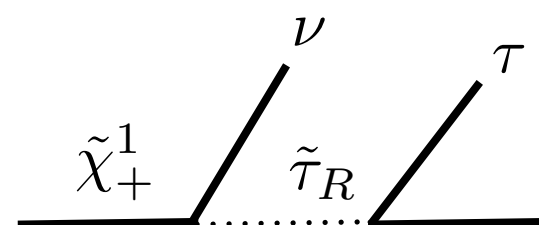
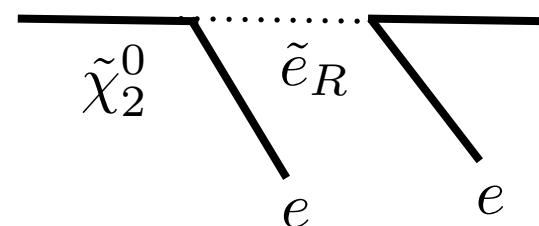
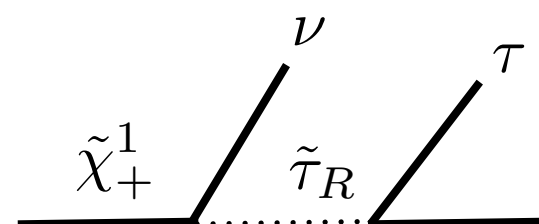
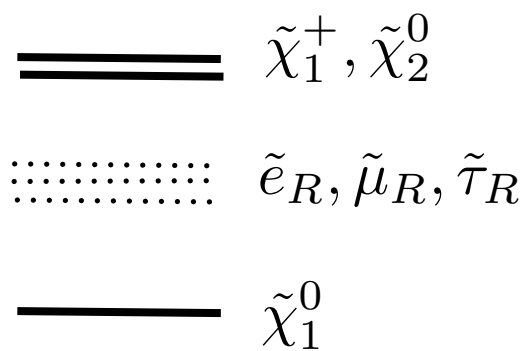
===== $\tilde{\chi}_1^+, \tilde{\chi}_2^0$

..... $\tilde{e}_R, \tilde{\mu}_R, \tilde{\tau}_R$

———— $\tilde{\chi}_1^0$

Decomposition

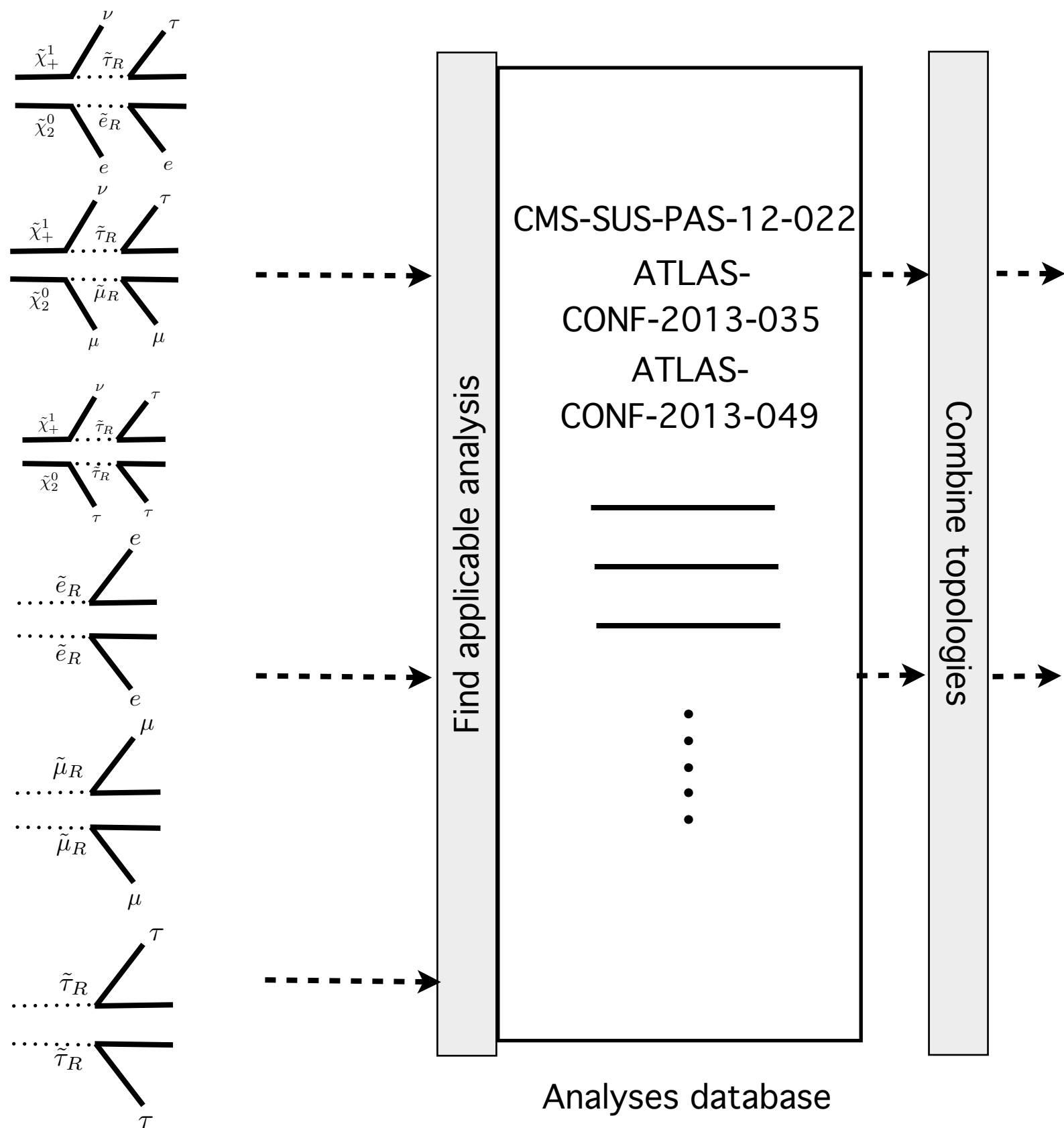
Given
Spectra

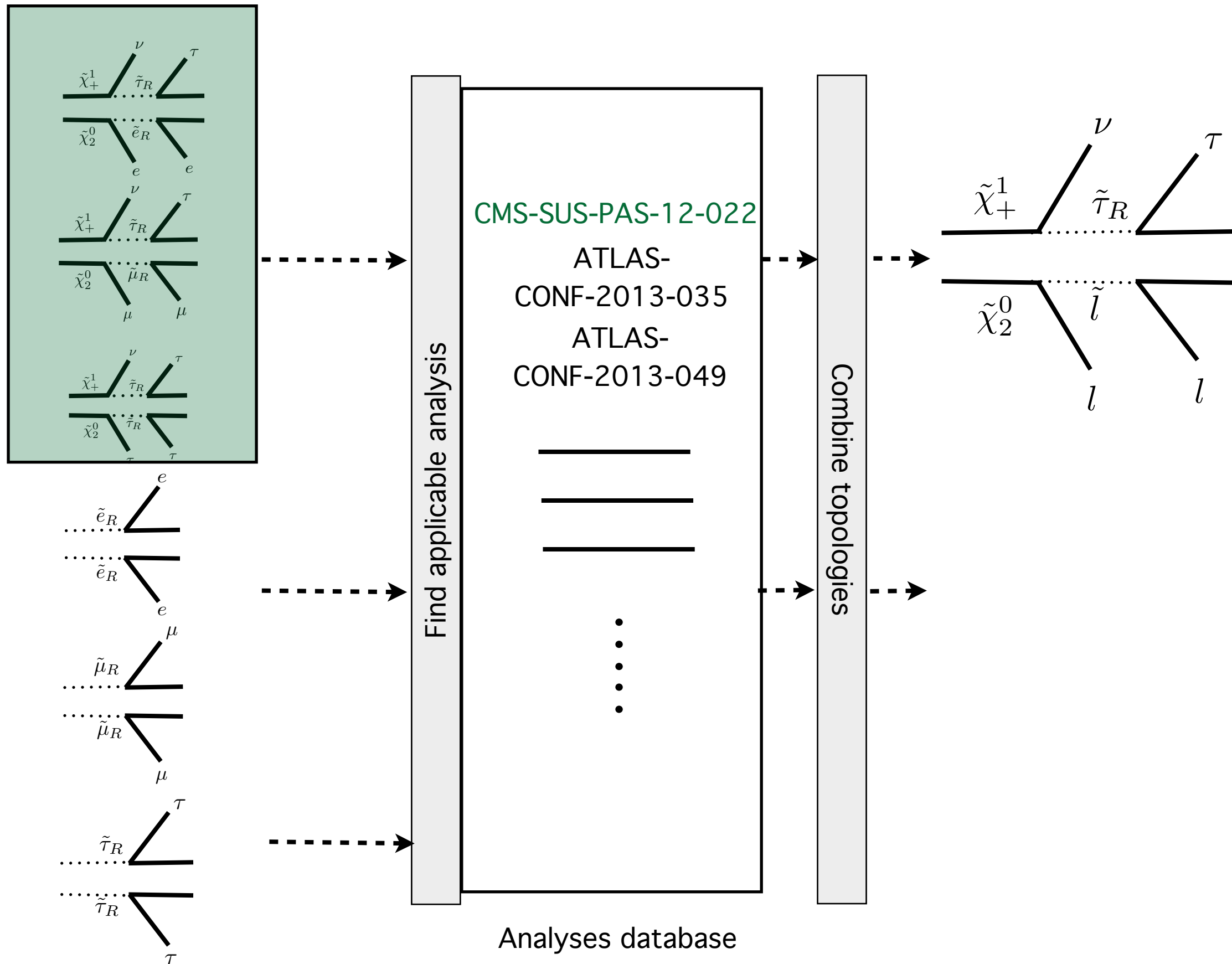


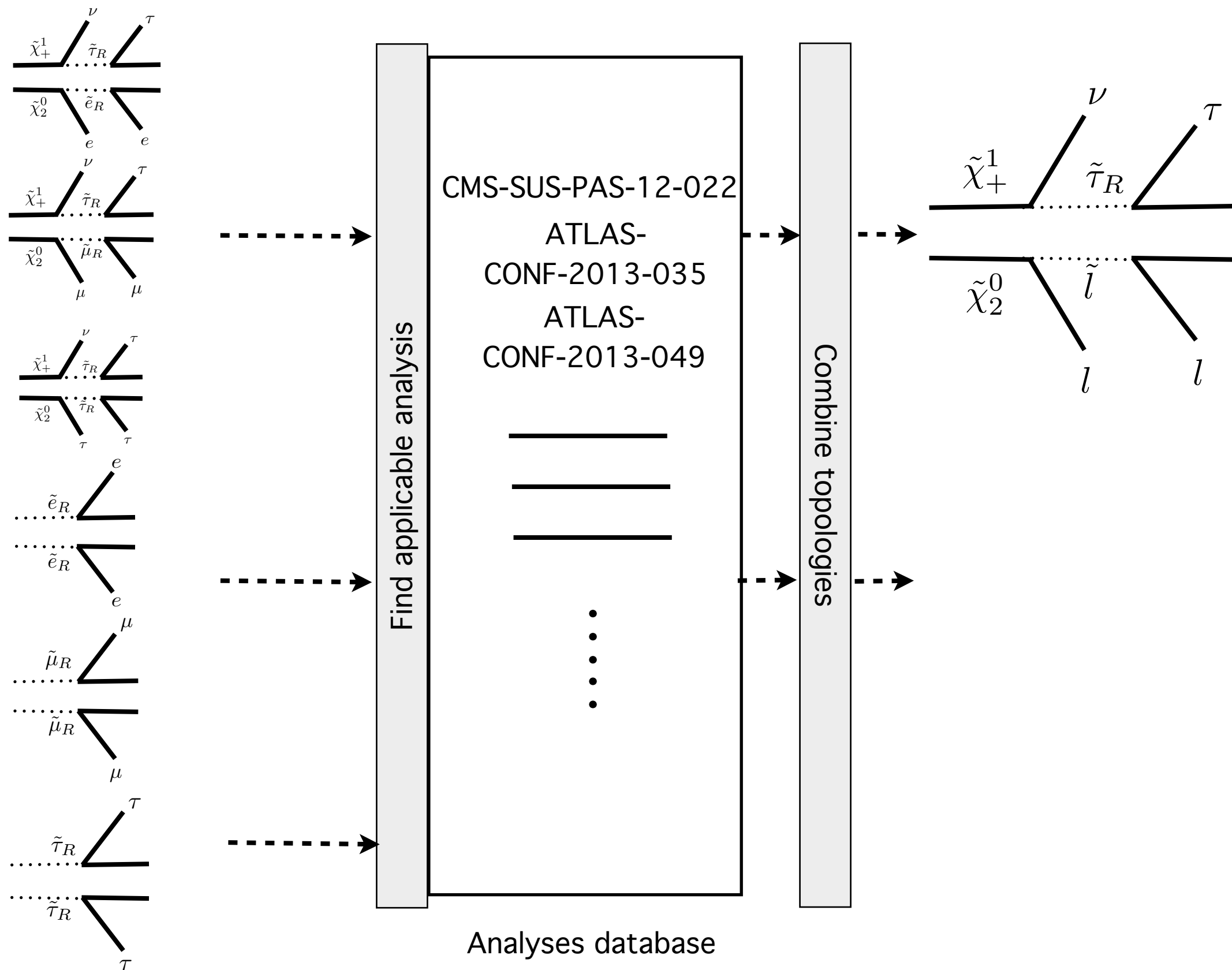
■

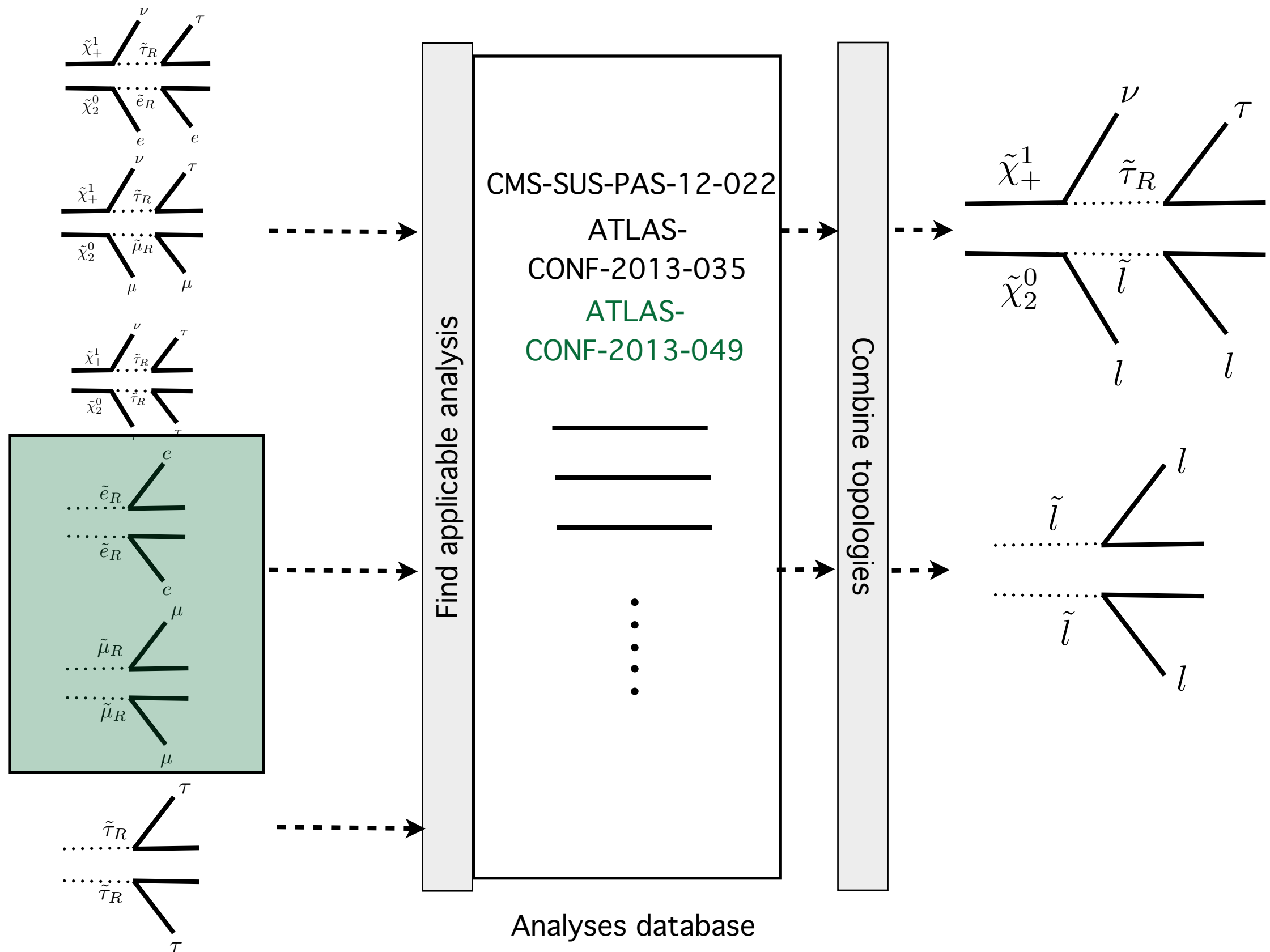
■

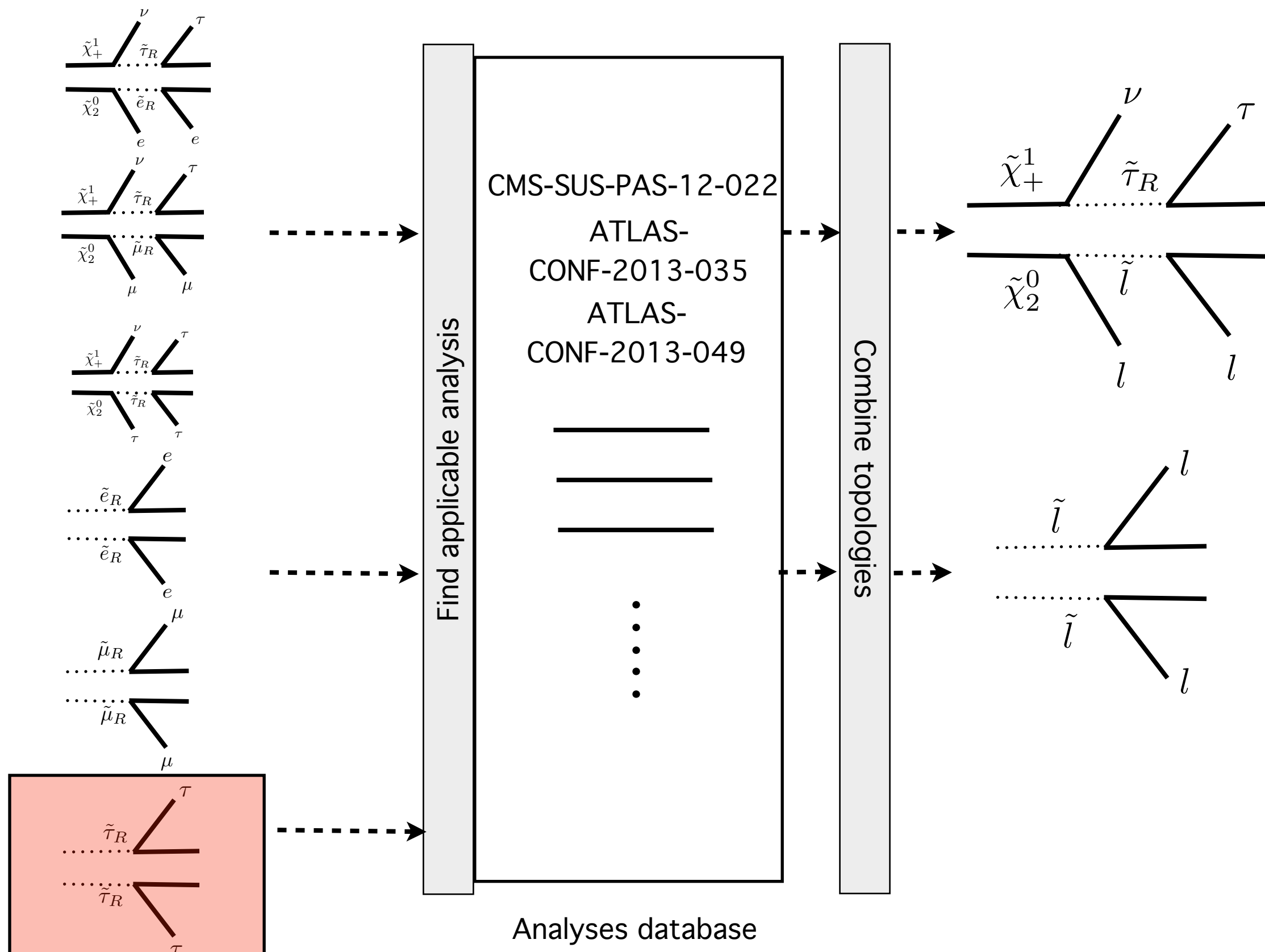
■

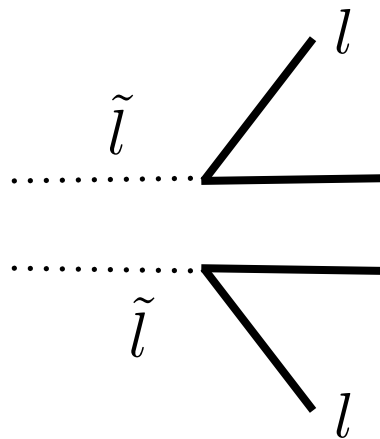
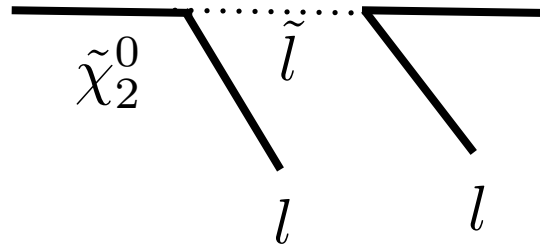
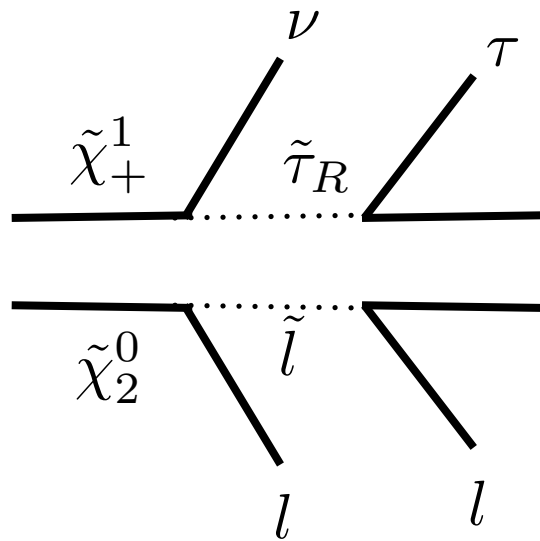






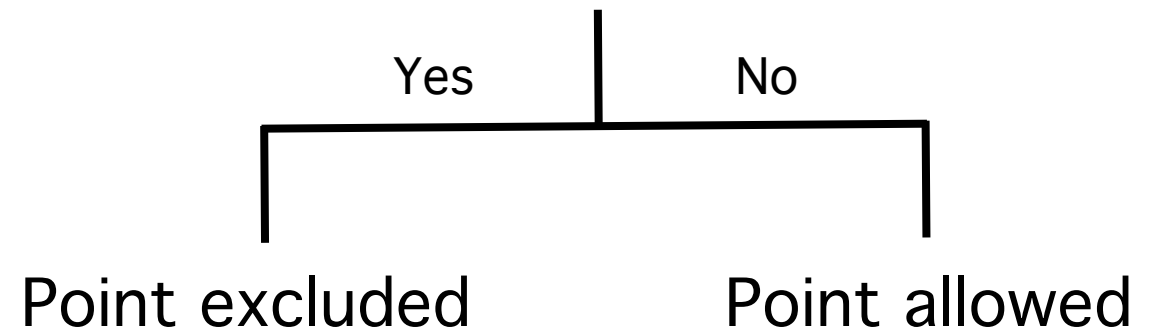






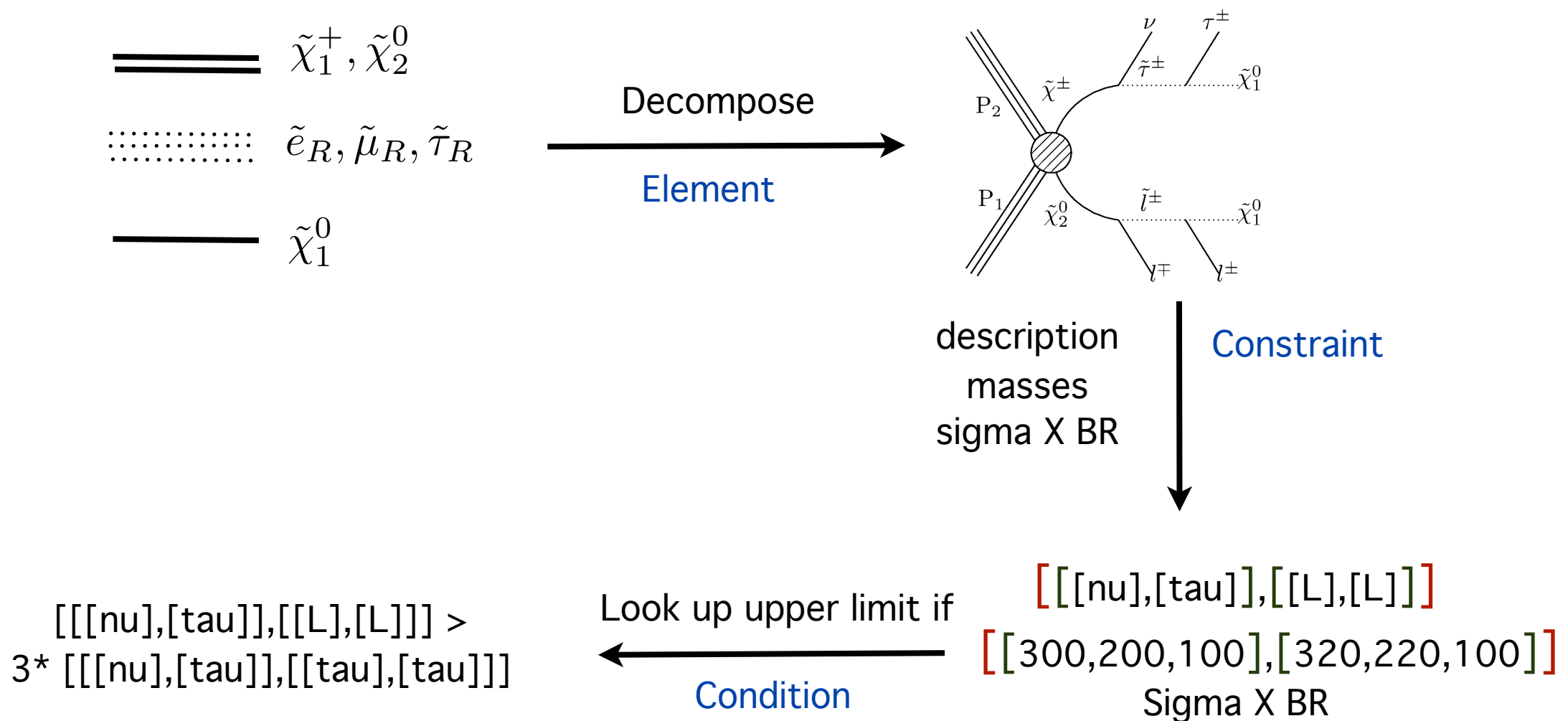
Look-up experimental limits

Is theory prediction > experimental limit?



SModelS framework

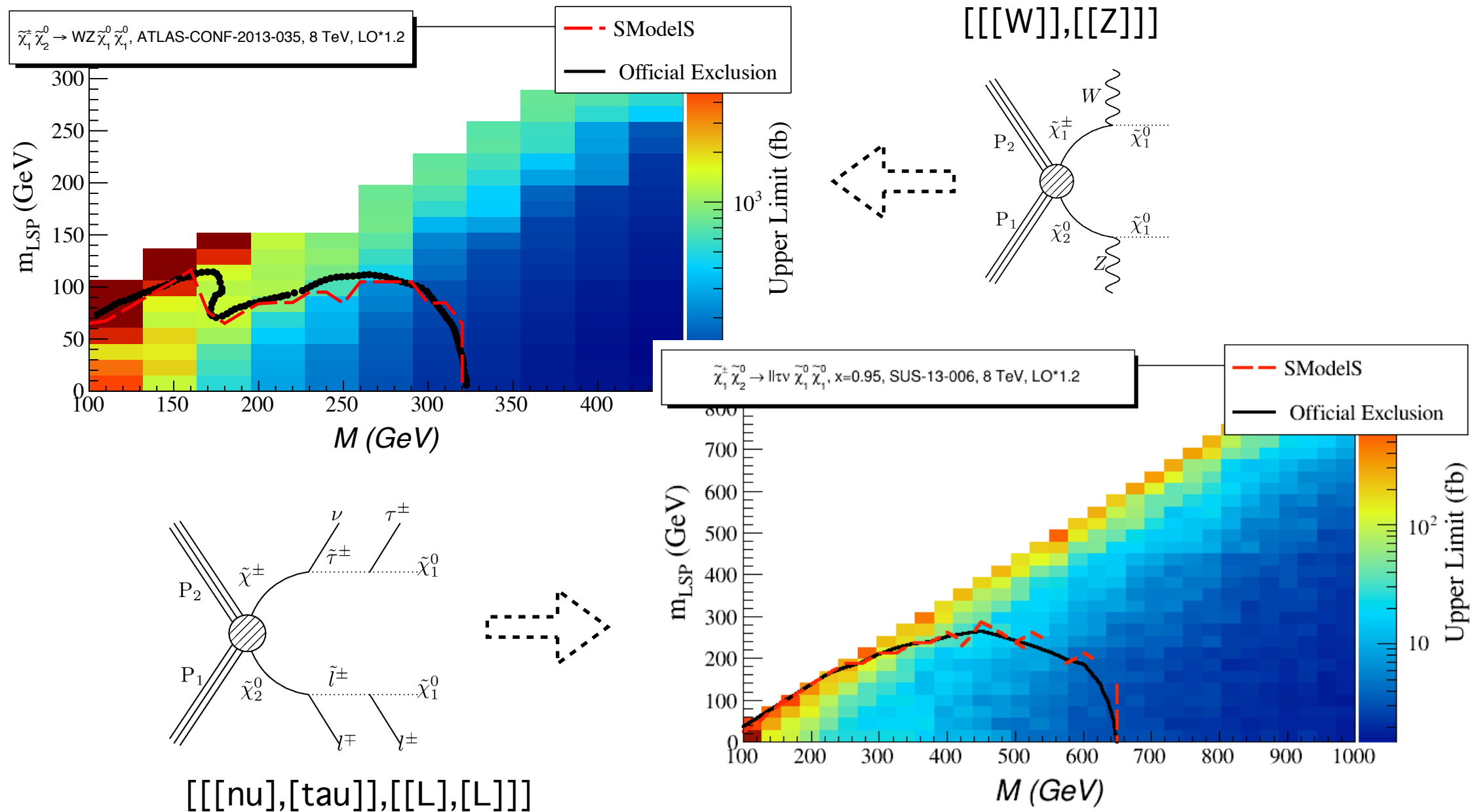
- Consider:



- The framework does not depend on characteristics of SUSY particles, can also be applied to decompose any BSM spectra of arbitrary complexity

How do we know it works?

- The code has been validated through the reproduction of various SMS exclusion curves



Typical examples of validation plot

Salient features

- Code is equipped to decompose any BSM model with a Z_2 symmetry
- It can handle compressed topologies
- It can take care of invisible decays
- It has the most comprehensive database of simplified model results, 22 CMS, 24 ATLAS (7 + 8 TeV)
- Now a web SLHA interface is available to check your point

Good, so what do you learn out of it?

based on:
arXiv: 1312.4175 [hep-ph]

SUSY scan - weak sector

- pMSSM scan over 6 parameters

- | | | | | | |
|-------|-------|--------------|-----------------|-----------------|----------|
| M_2 | μ | $\tan \beta$ | $M_{\tilde{L}}$ | $M_{\tilde{E}}$ | A_τ |
| 0.1–1 | 0.1–1 | 3–60 | 0.1–1 | 0.1–1 | ± 1 |

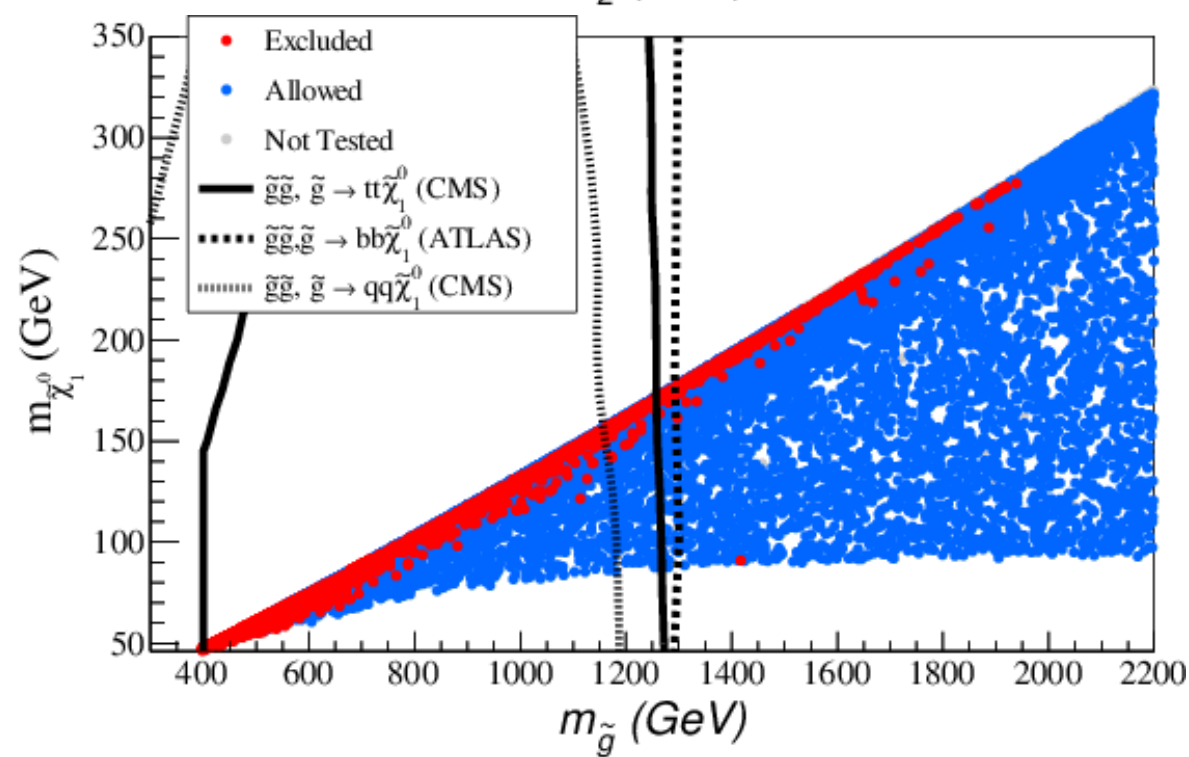
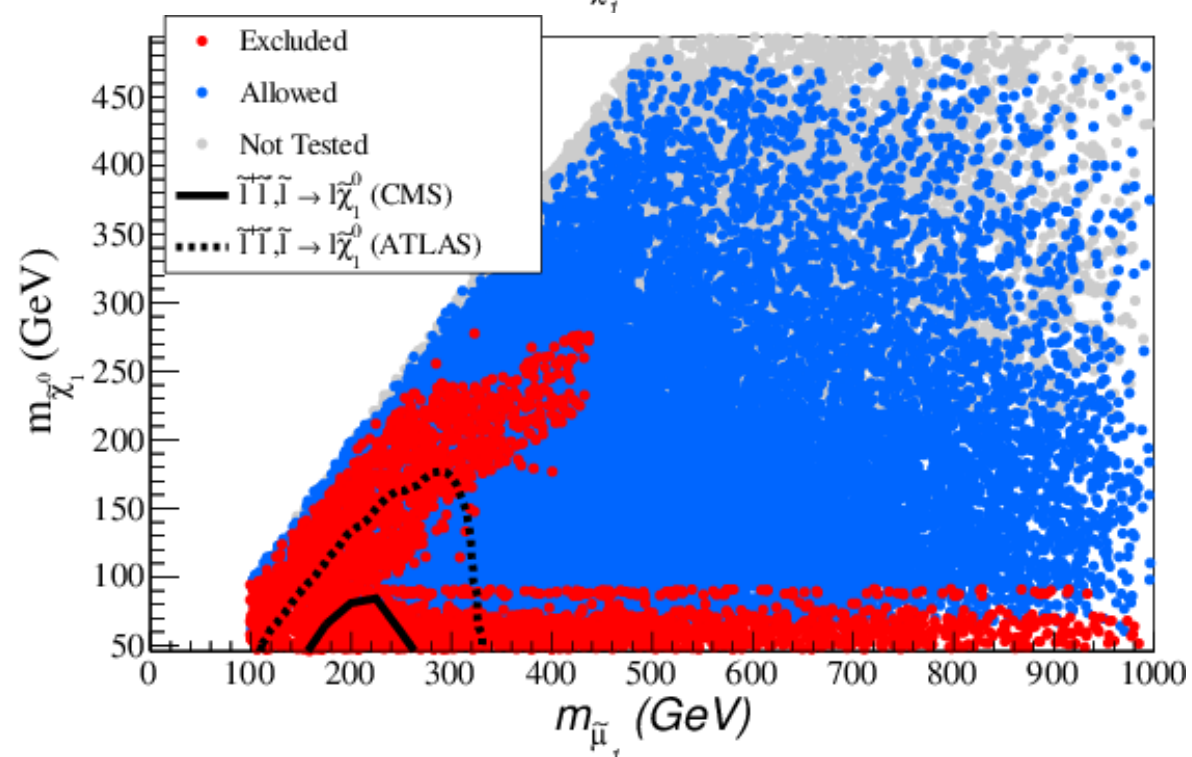
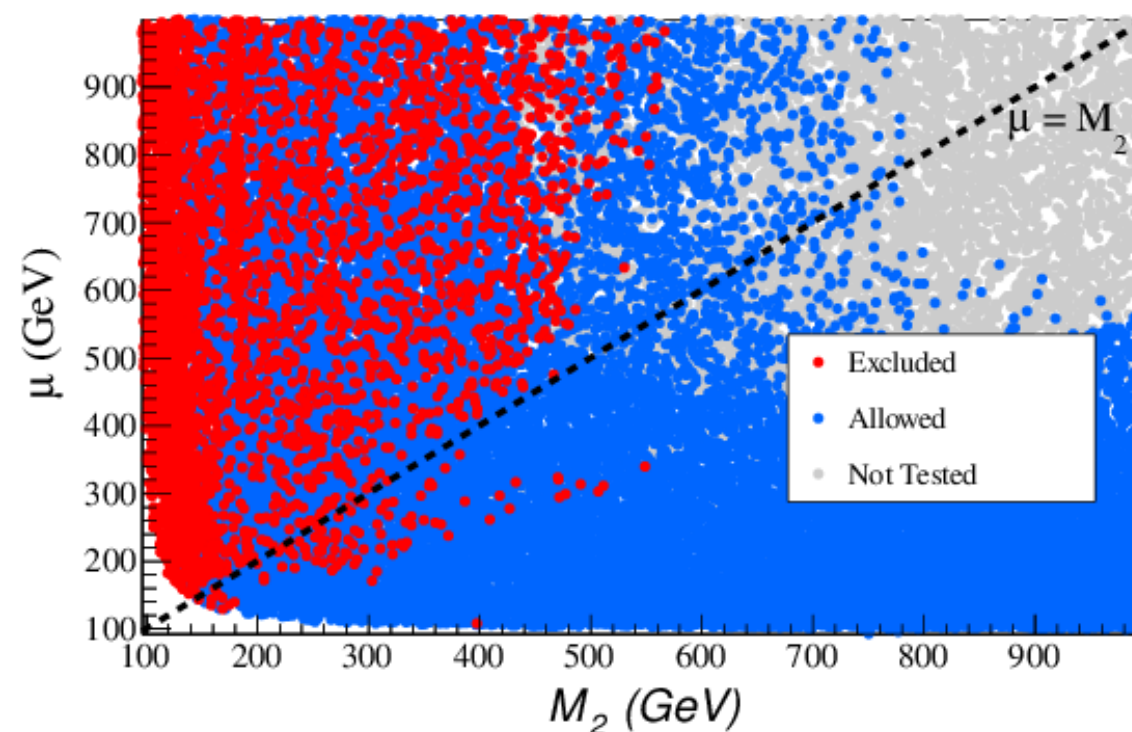
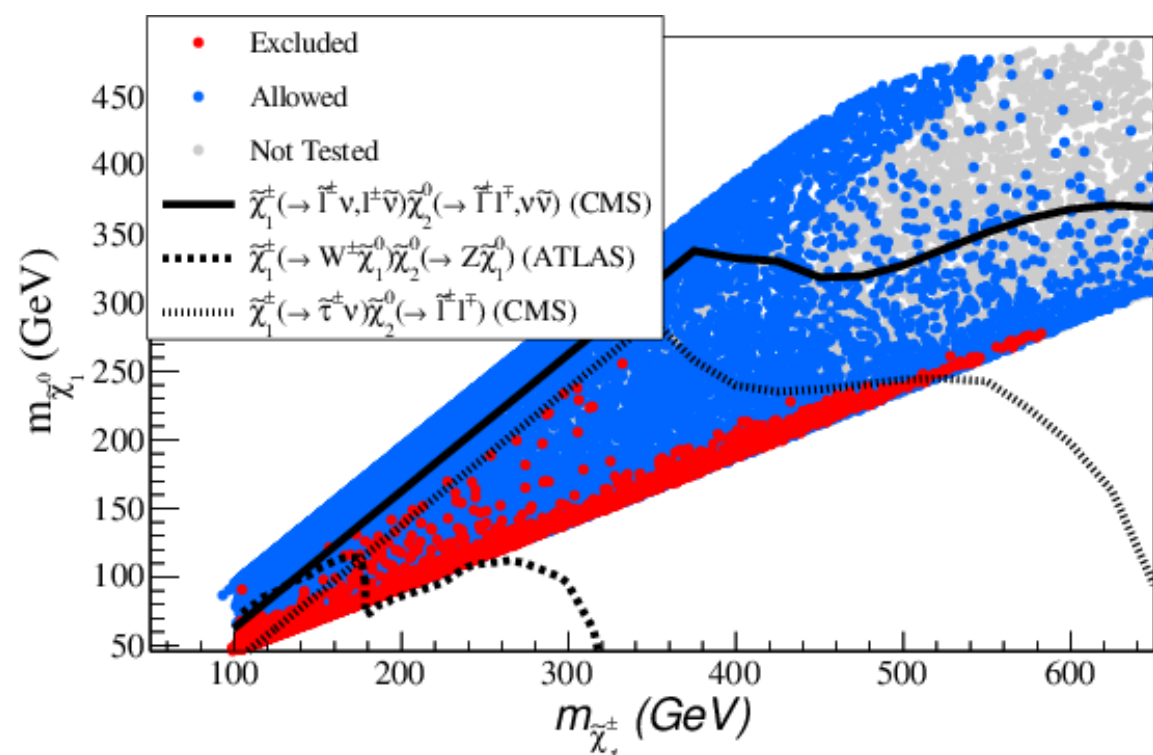
- Gaugino masses obey GUT relation
- Flavor constraints, invisible Z width, Higgs mass, LEP limits imposed

- Limits obtained will always be conservative

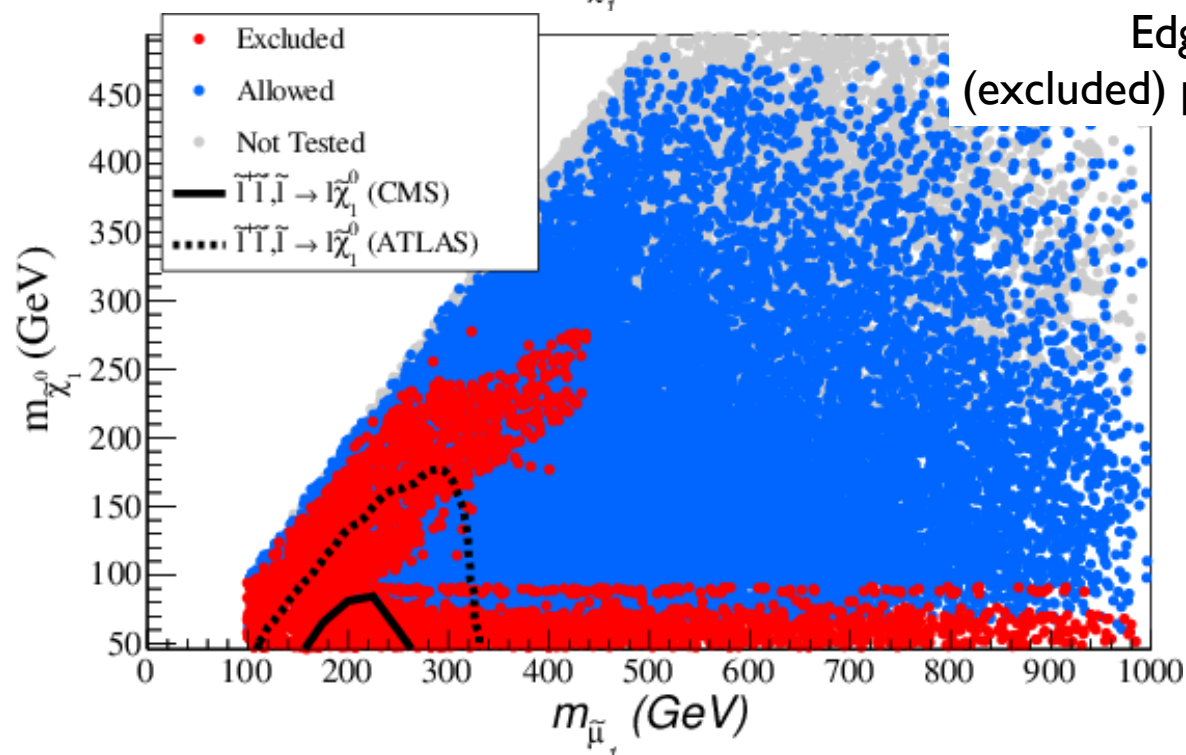
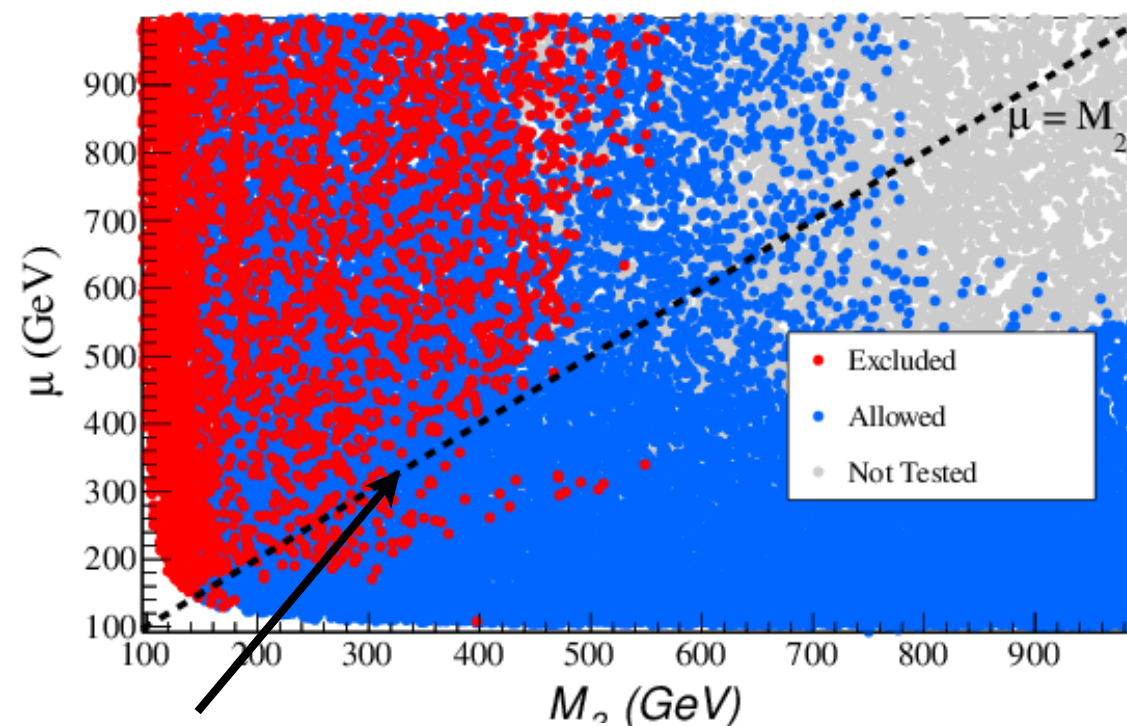
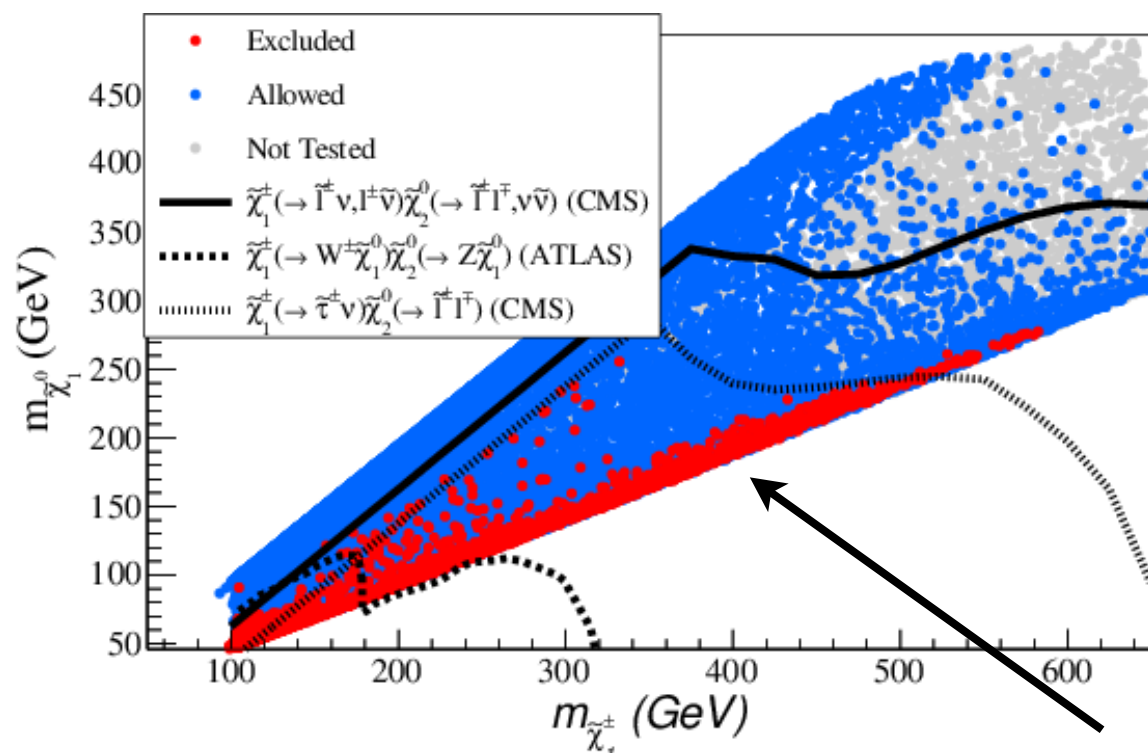
- We probe electroweak -ino decays via WZ and sleptons, direct slepton production and gluino decays

NB: A similar scan was also performed for strong sector particles

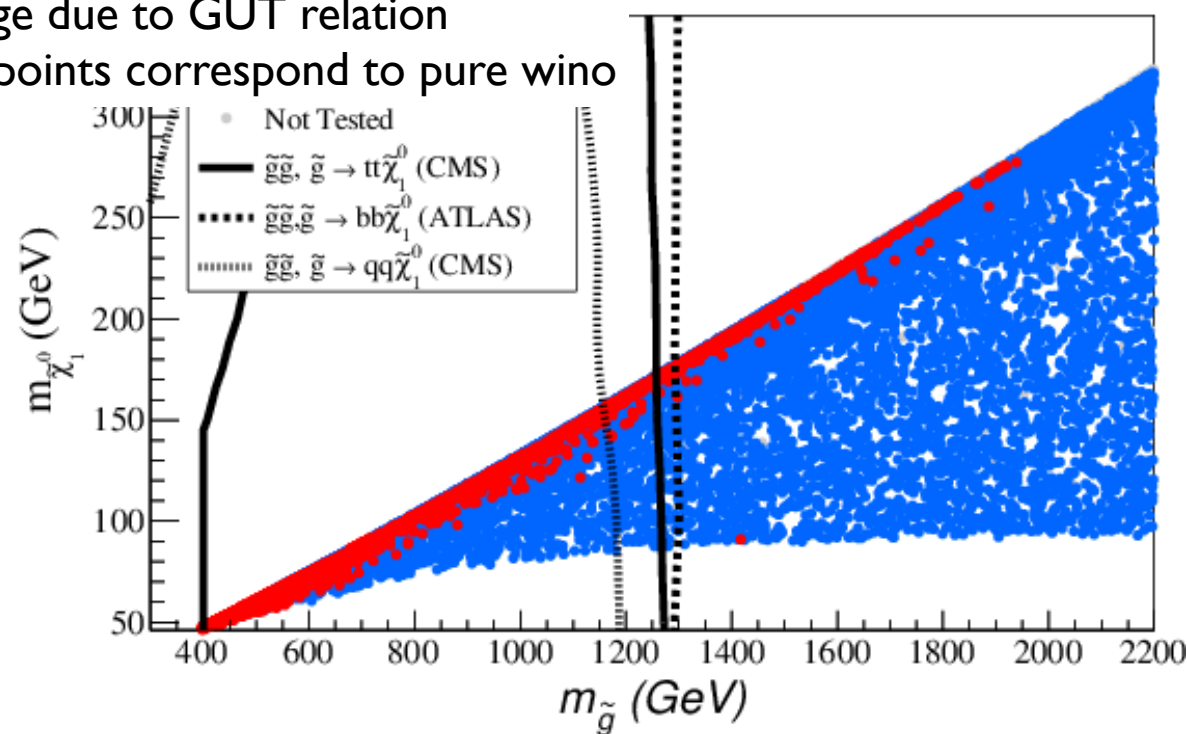
SUSY scan - weak sector



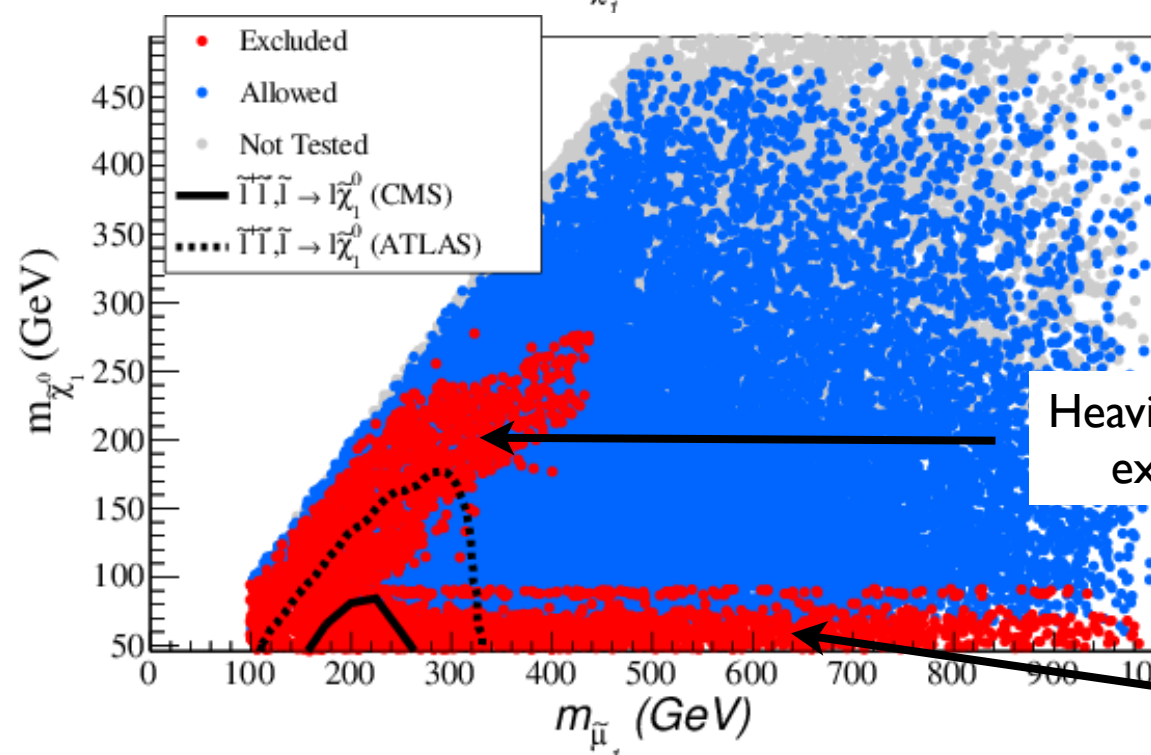
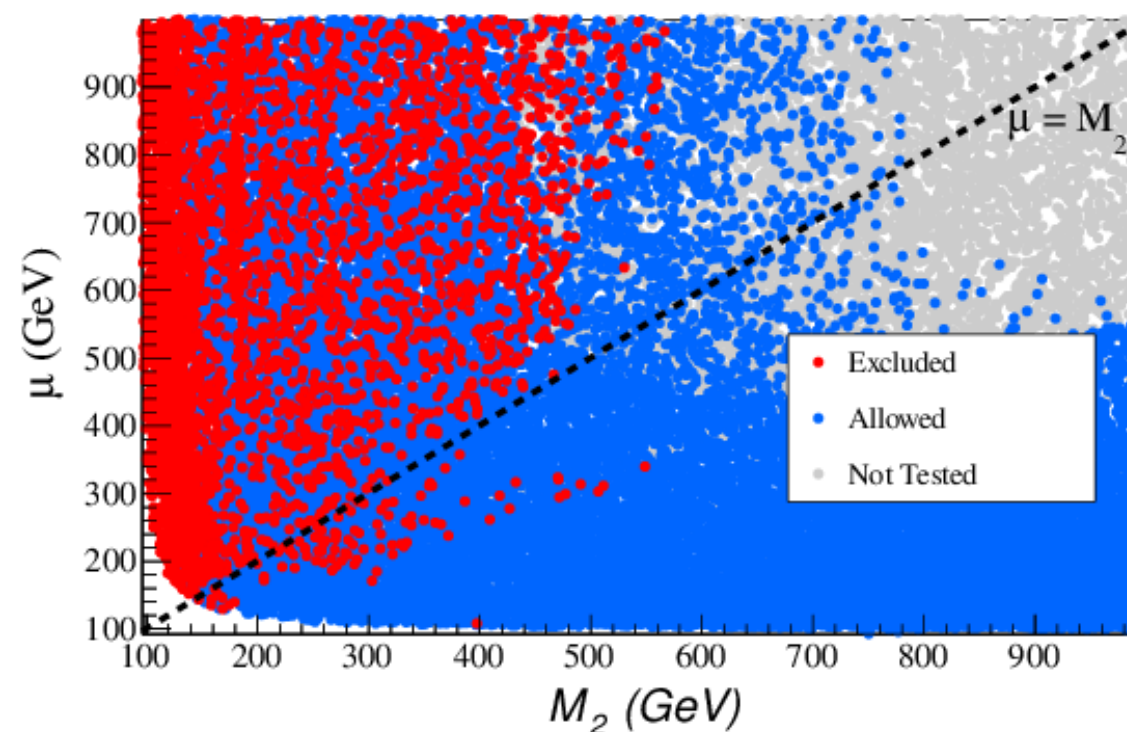
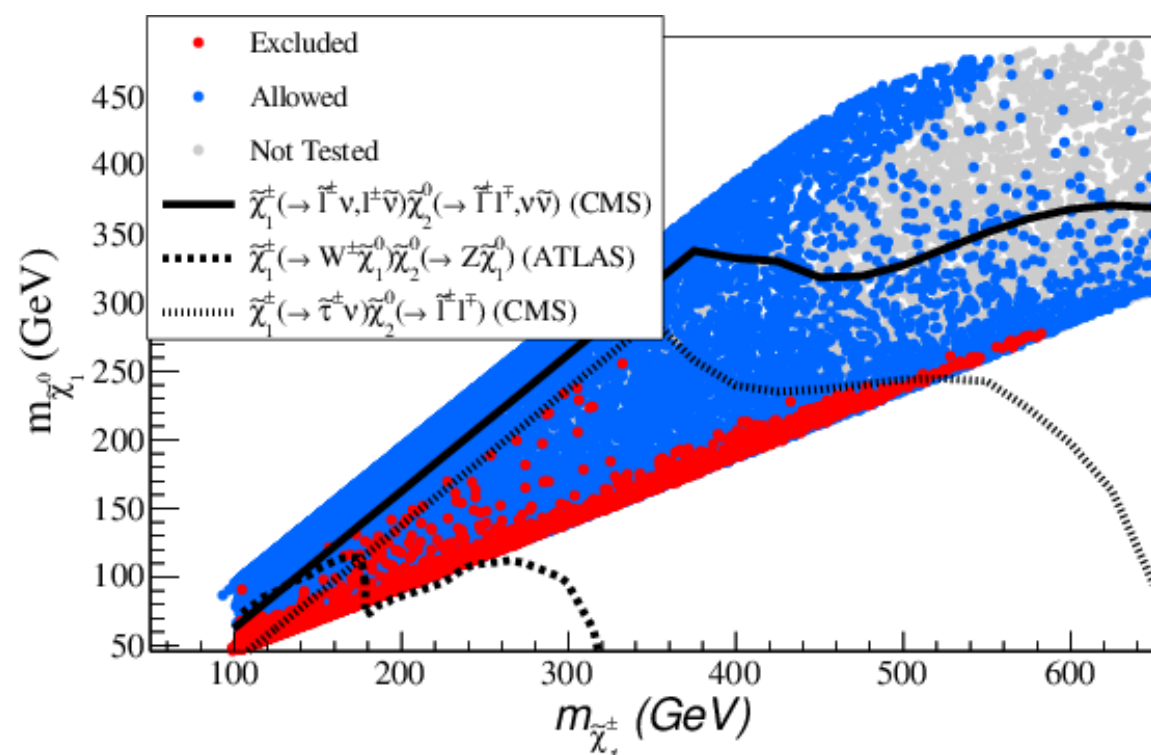
SUSY scan - weak sector



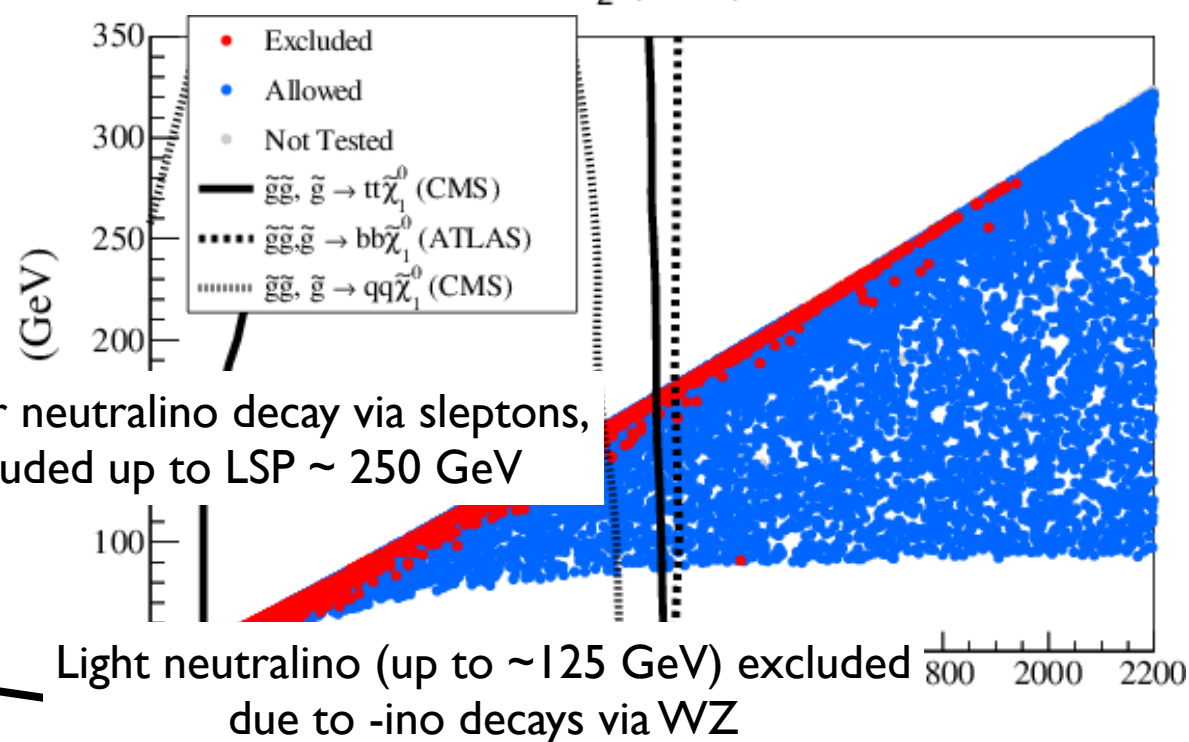
Edge due to GUT relation
(excluded) points correspond to pure wino



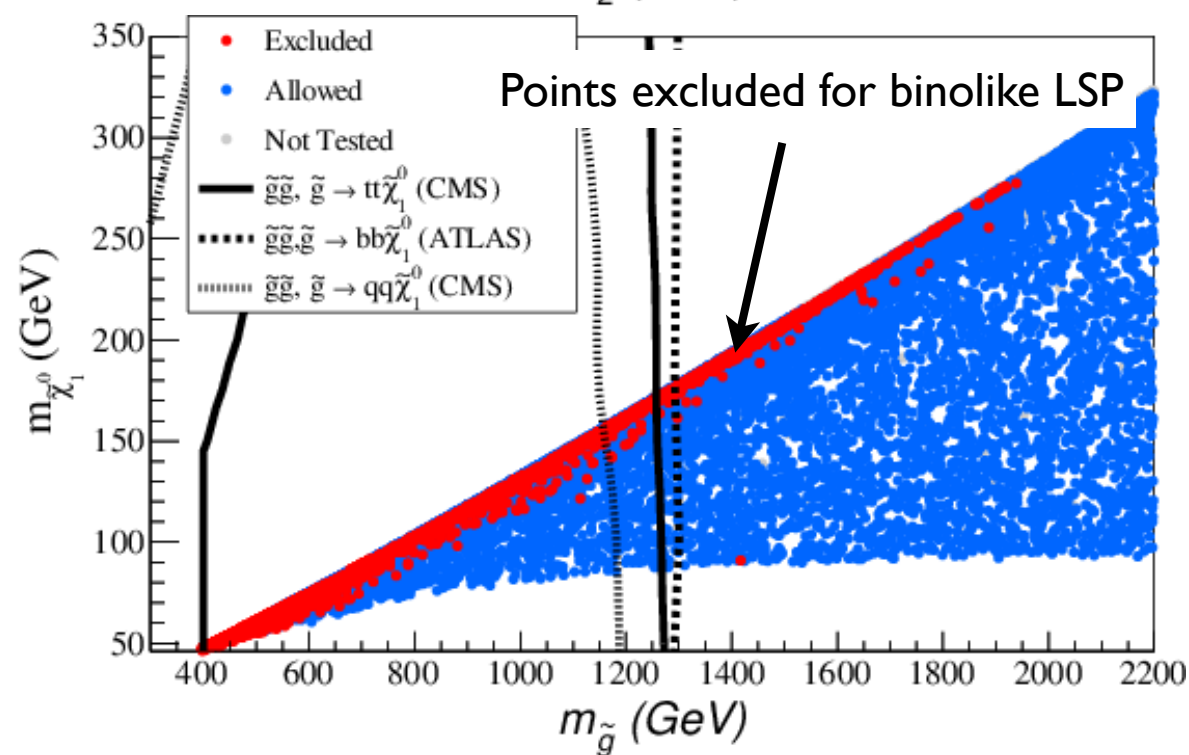
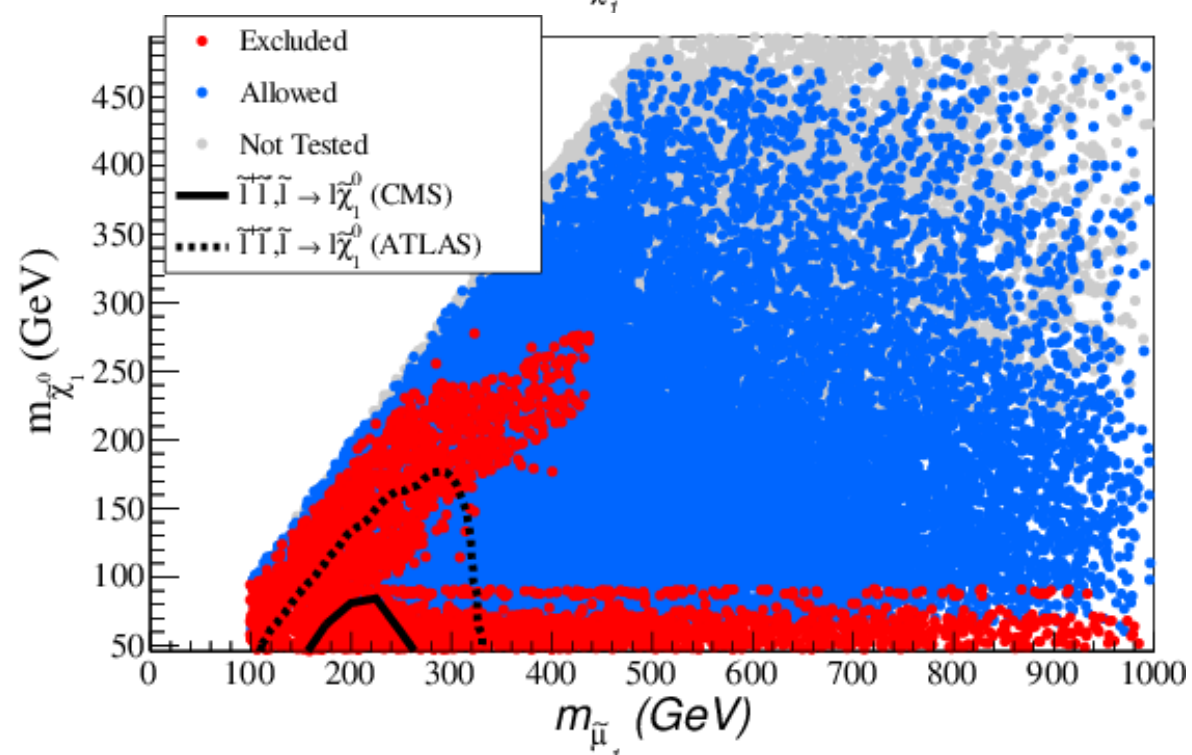
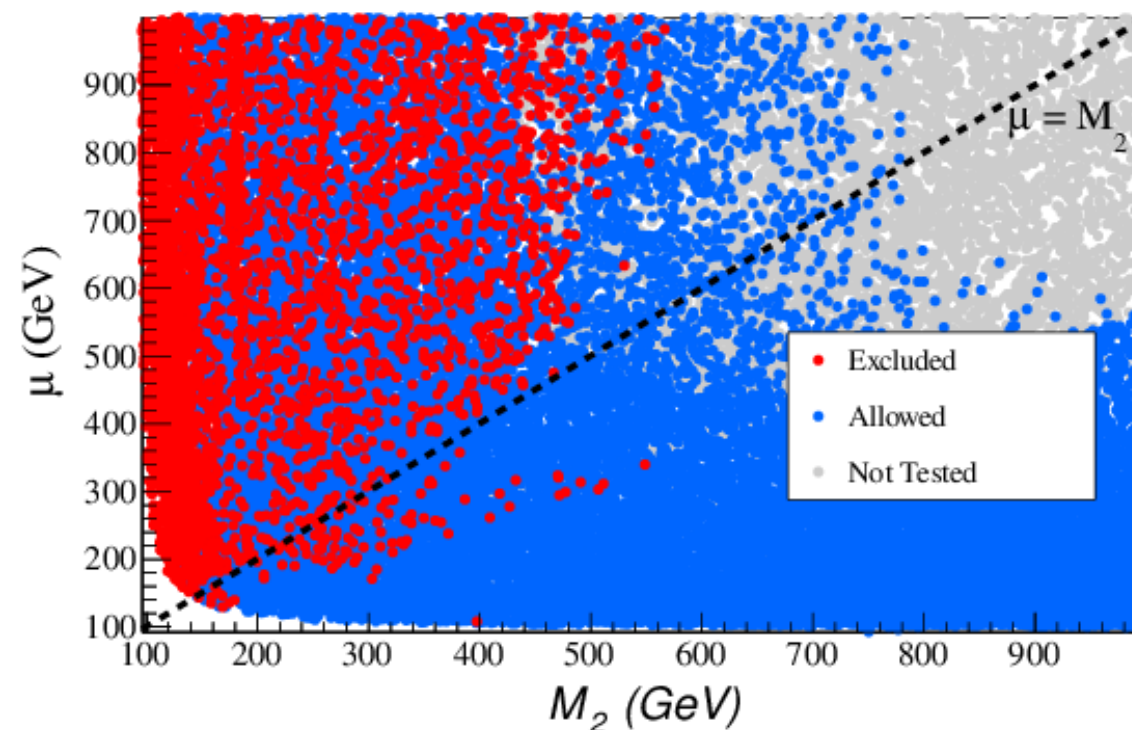
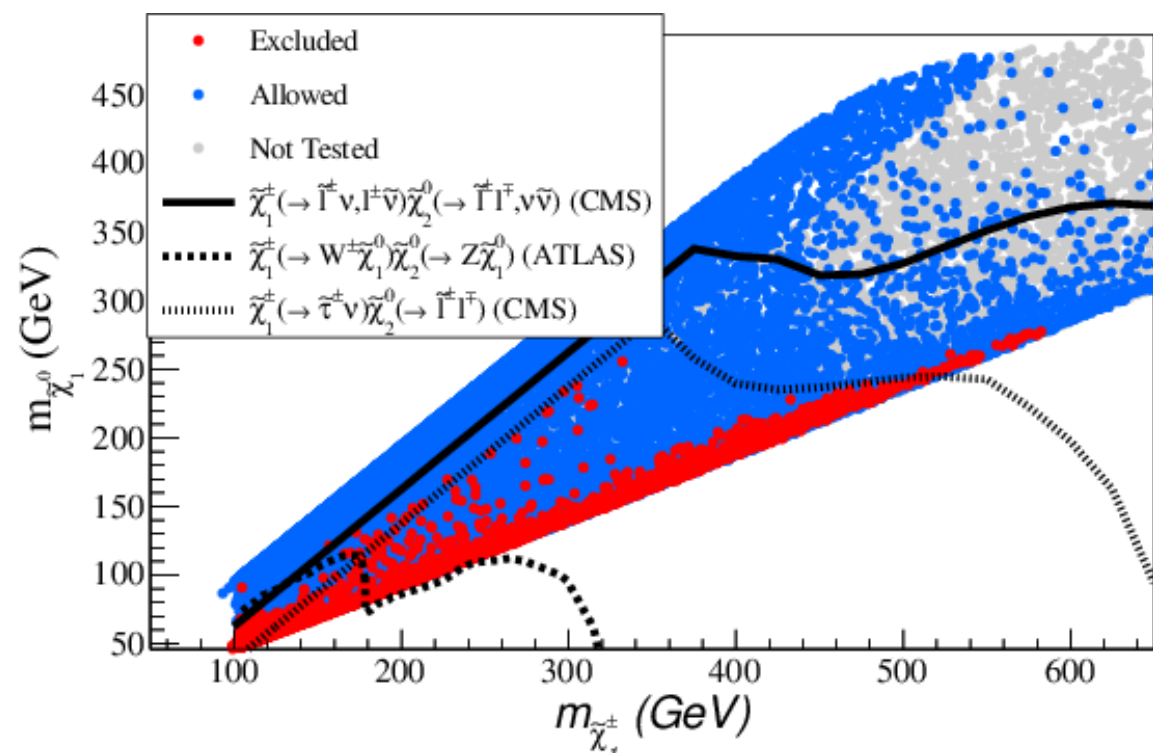
SUSY scan - weak sector



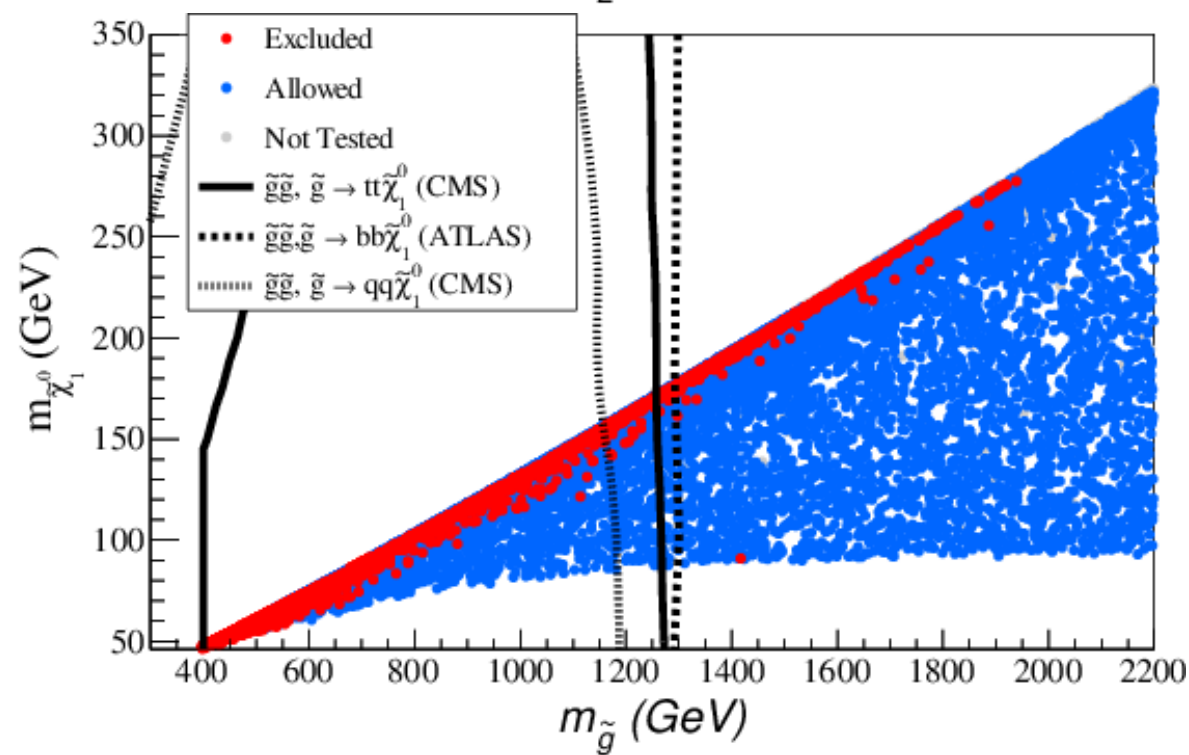
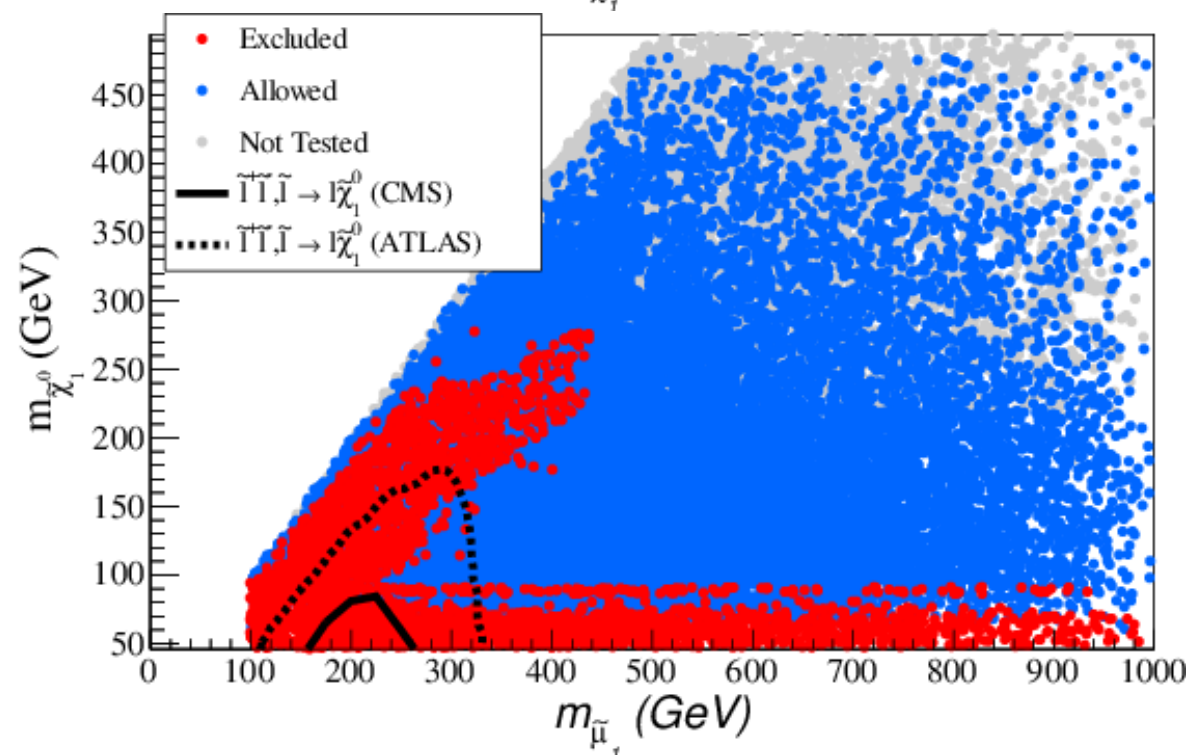
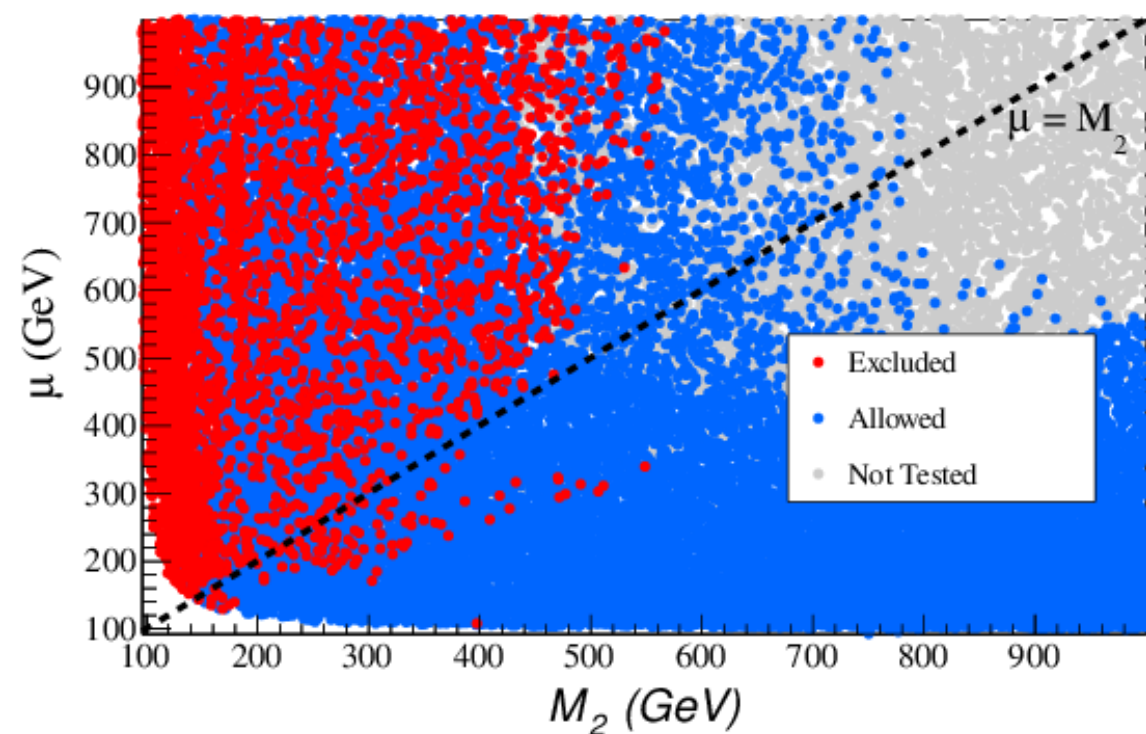
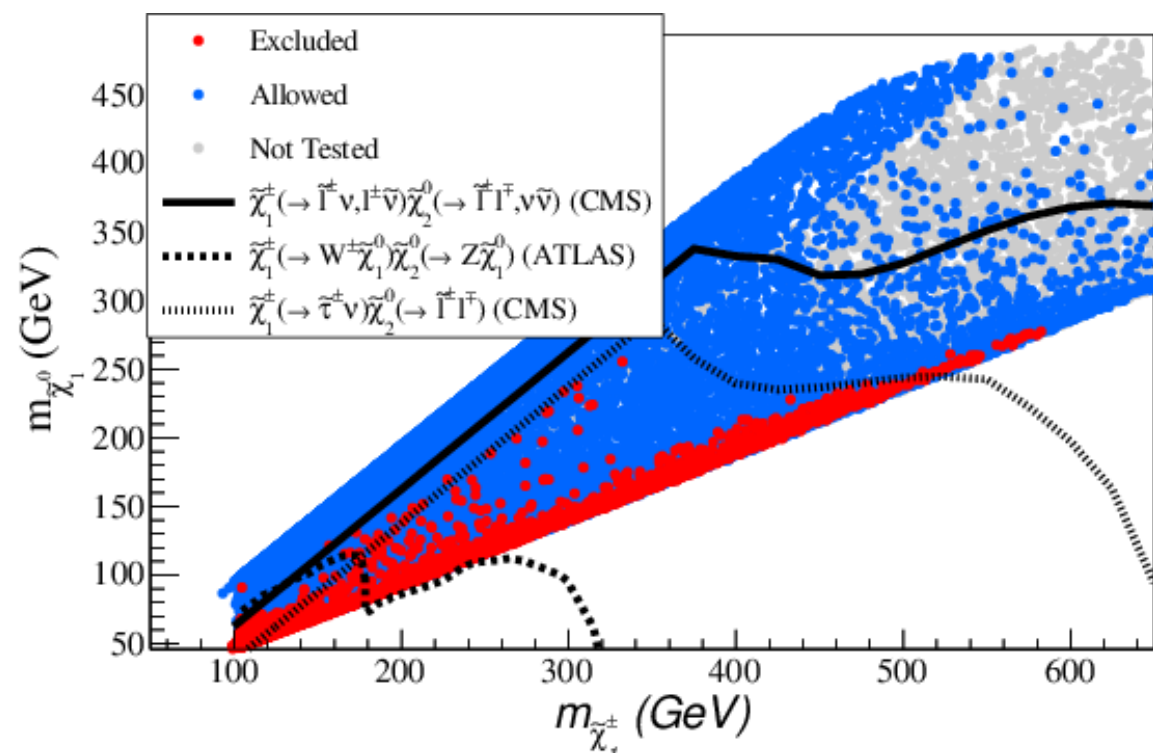
Heavier neutralino decay via sleptons, excluded up to LSP ~ 250 GeV



SUSY scan - weak sector

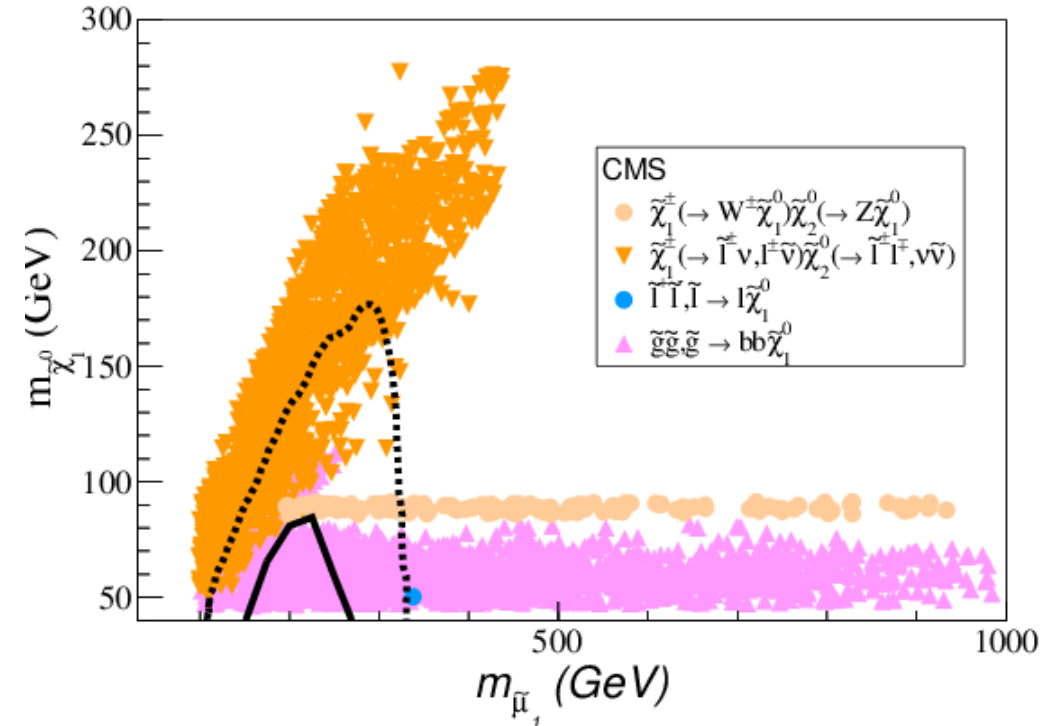
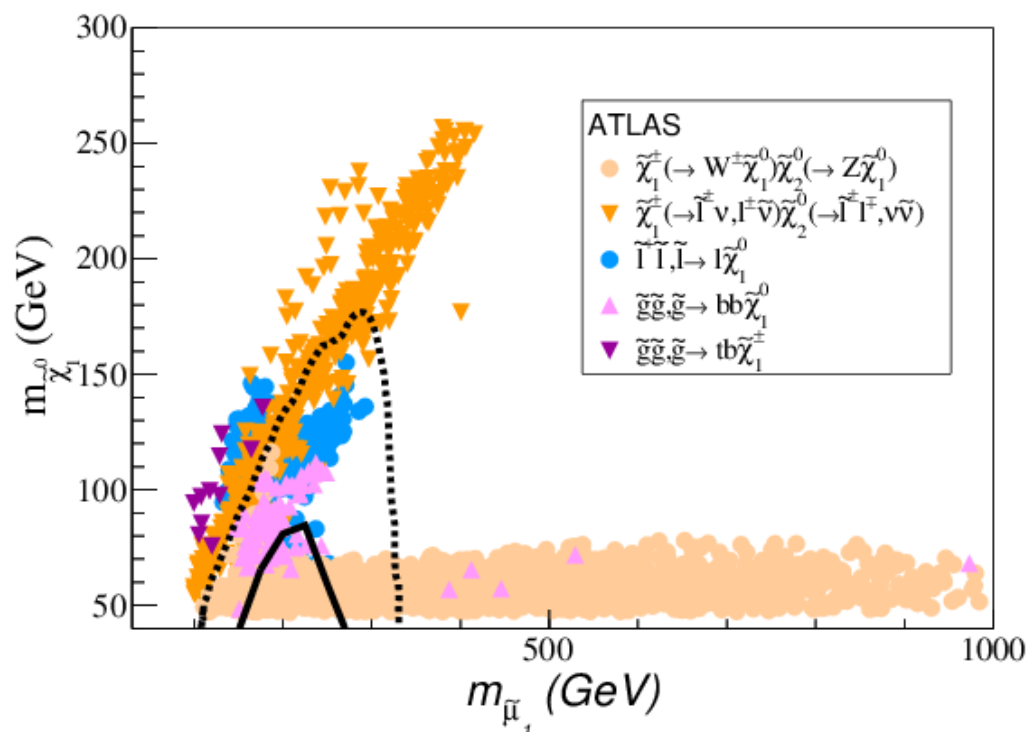
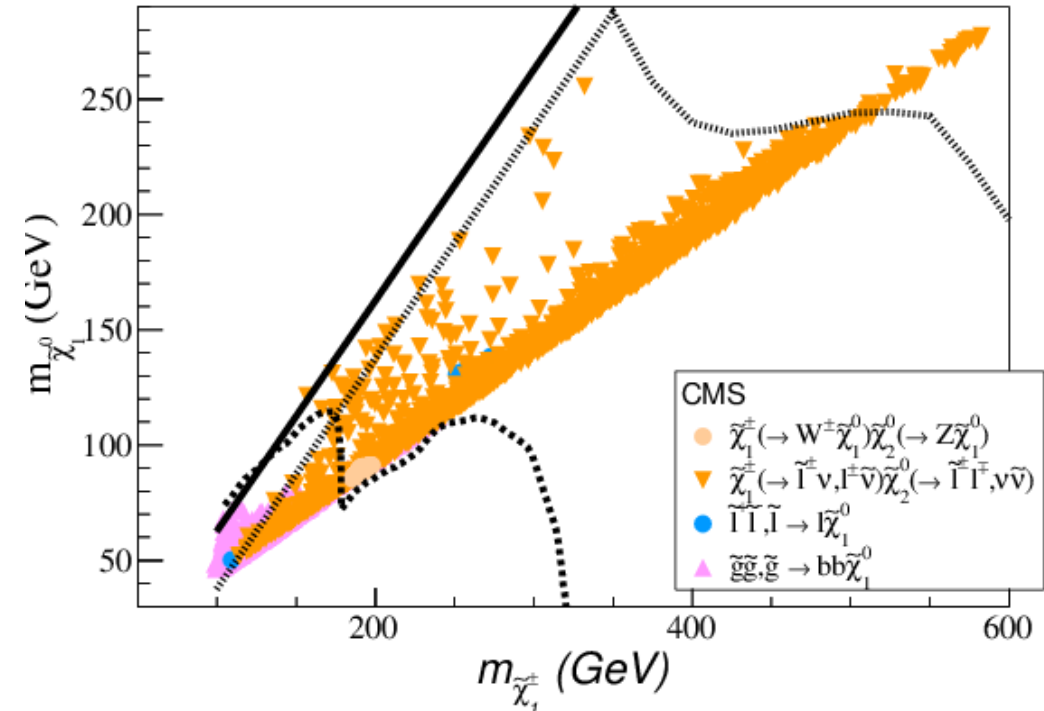
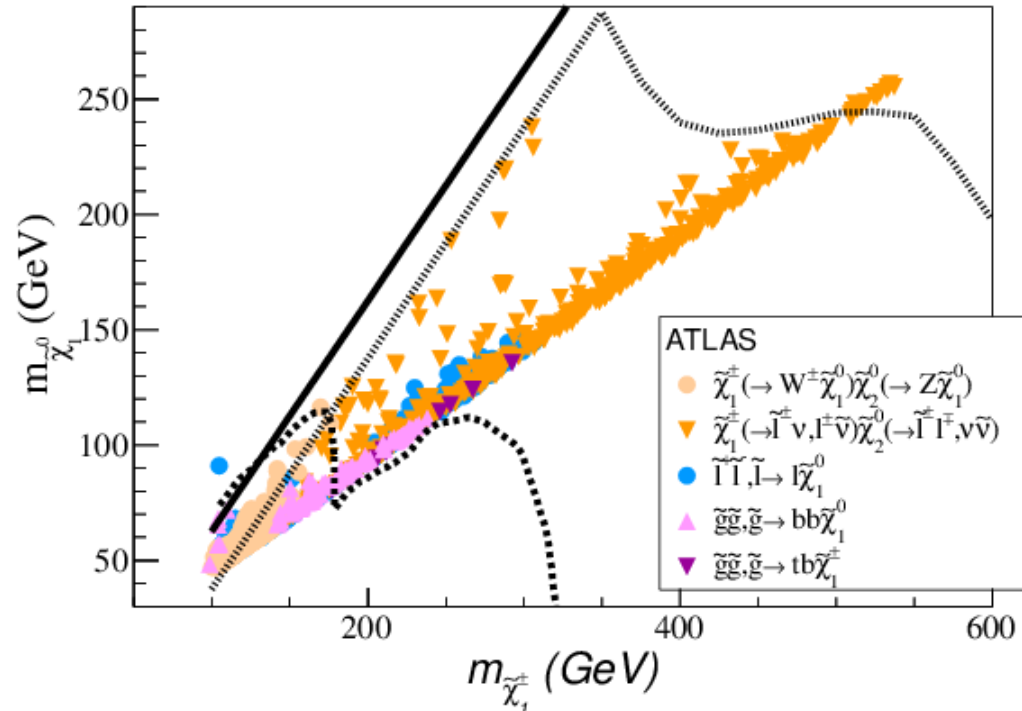


SUSY scan - weak sector



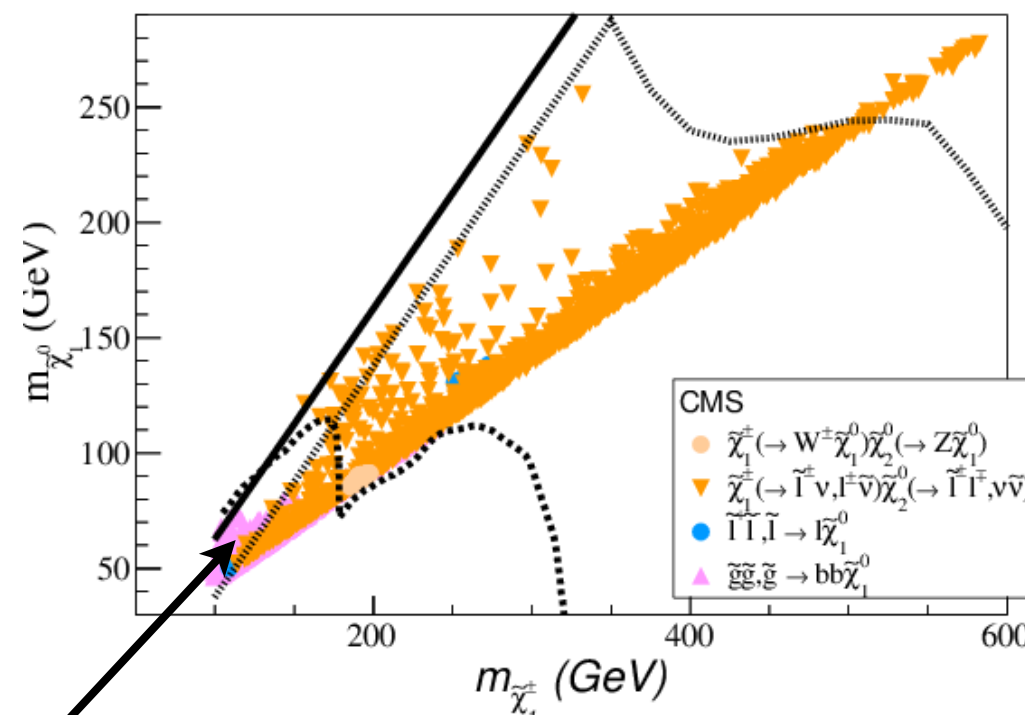
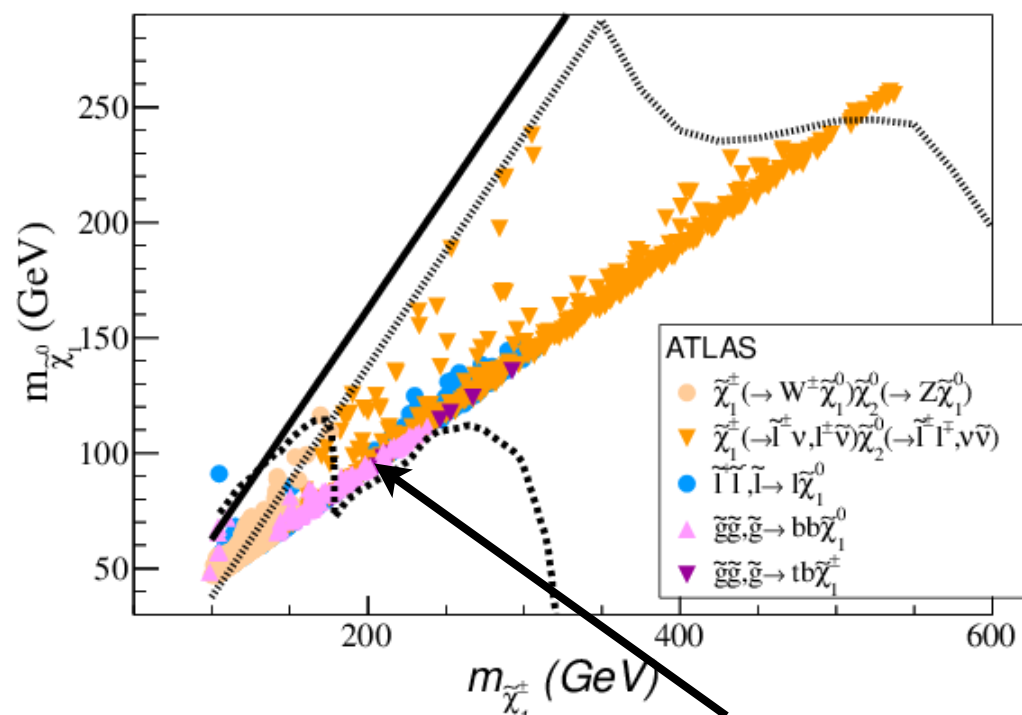
SUSY scan - weak sector

- Breakdown of the excluded parameter space by analysis

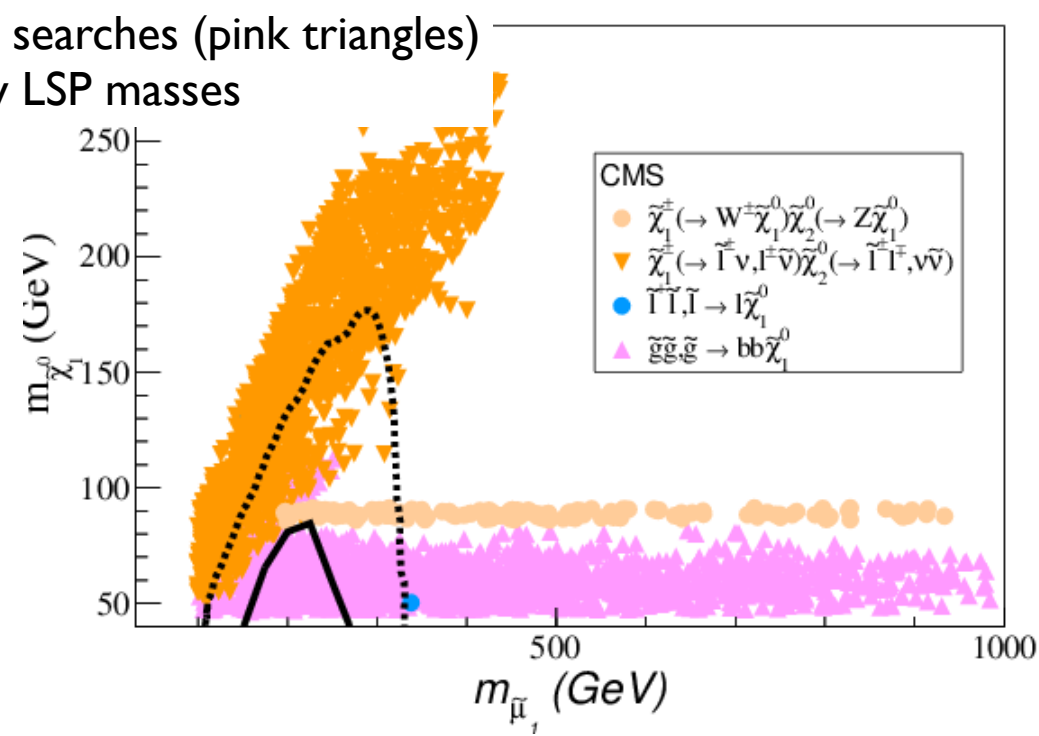
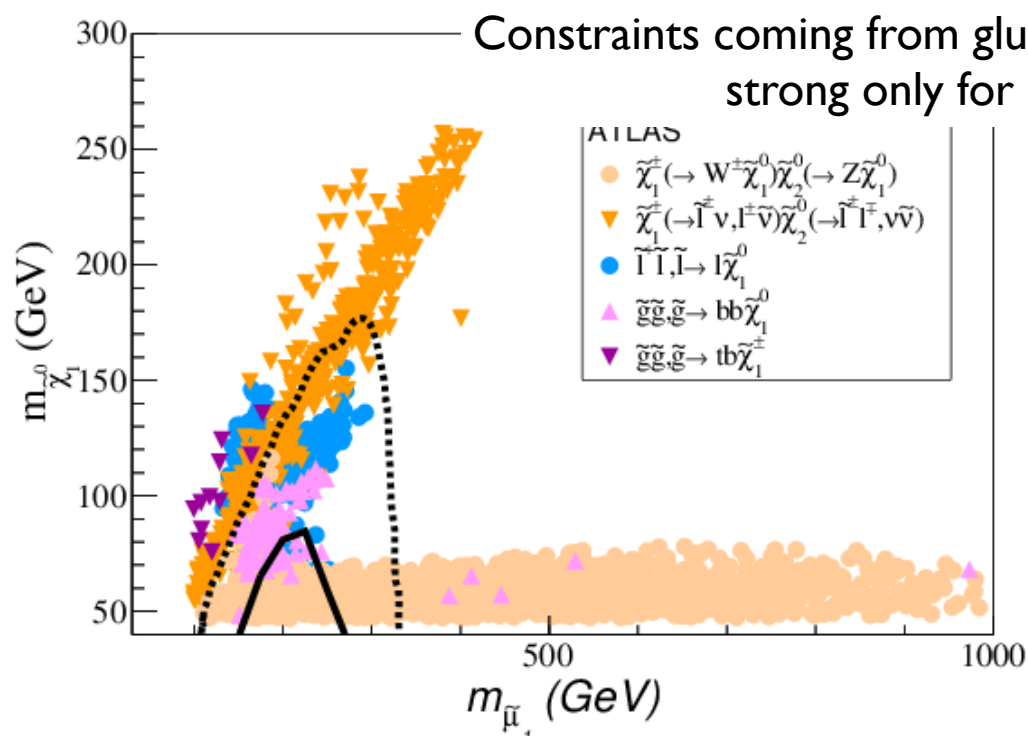


SUSY scan - weak sector

- Breakdown of the excluded parameter space by analysis

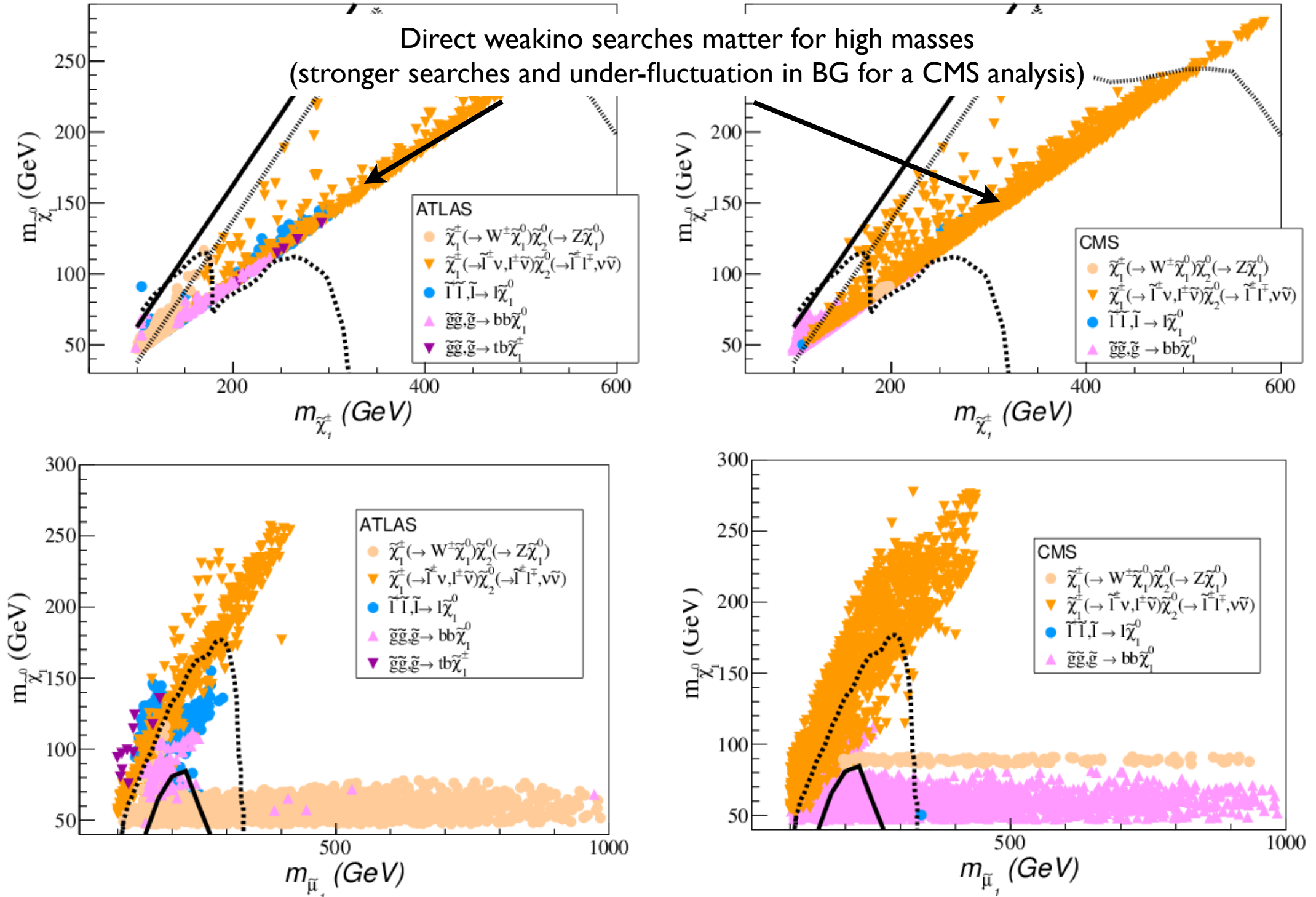


Constraints coming from gluino searches (pink triangles)
strong only for low LSP masses



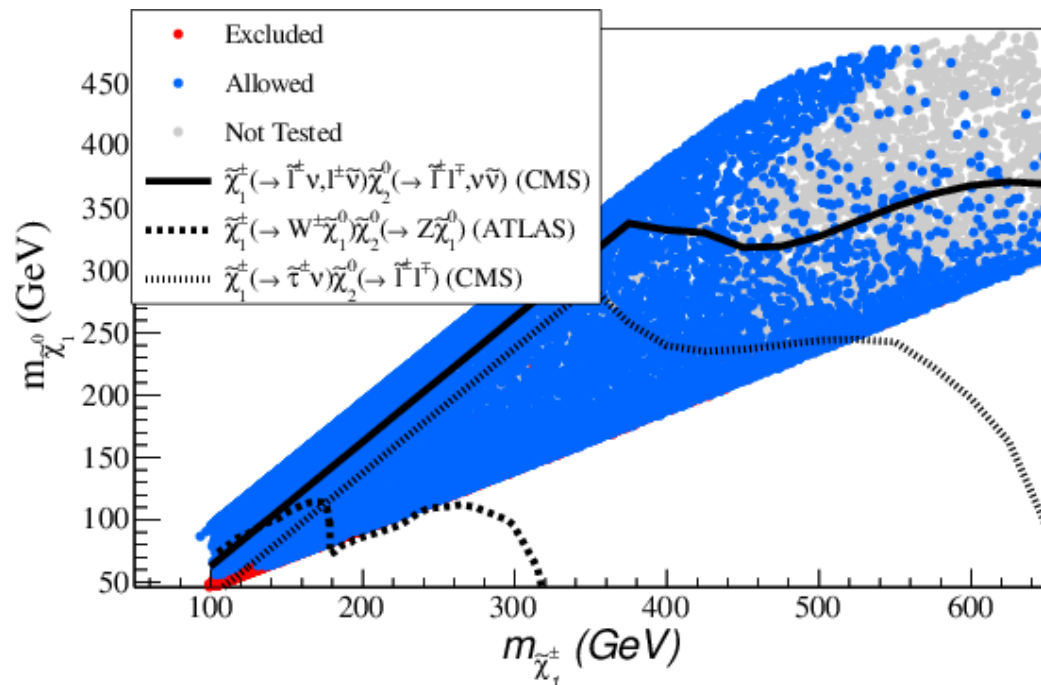
SUSY scan - weak sector

- Breakdown of the excluded parameter space by analysis



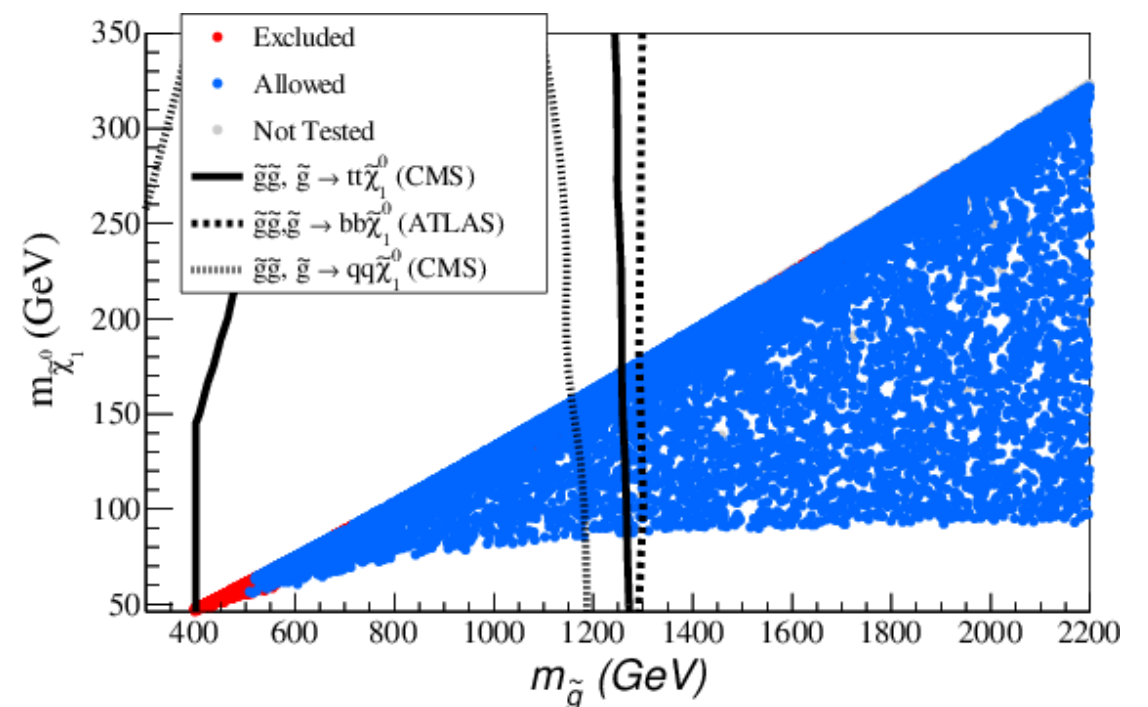
SUSY scan - weak sector

- Some allowed points may lie below excluded points



- Gluino decays via on-shell squarks are kinematically forbidden for small masses
- Uncovered gluino decay topologies e.g. $BR(\tilde{g} \rightarrow \tilde{\chi}^\pm + tb)$

- Chargino - LSP nature, higgsino have smaller production cross-section
- Right handed sleptons have smaller production CS



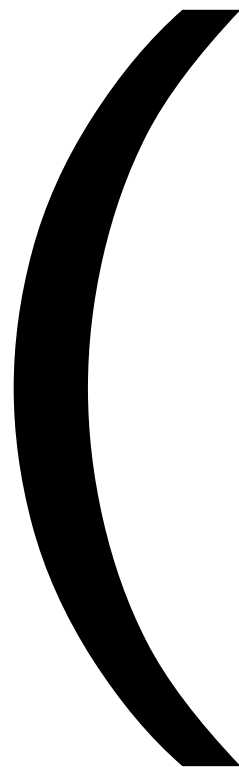
Can it be used to test parameter space for some interesting scenario?

A real life application

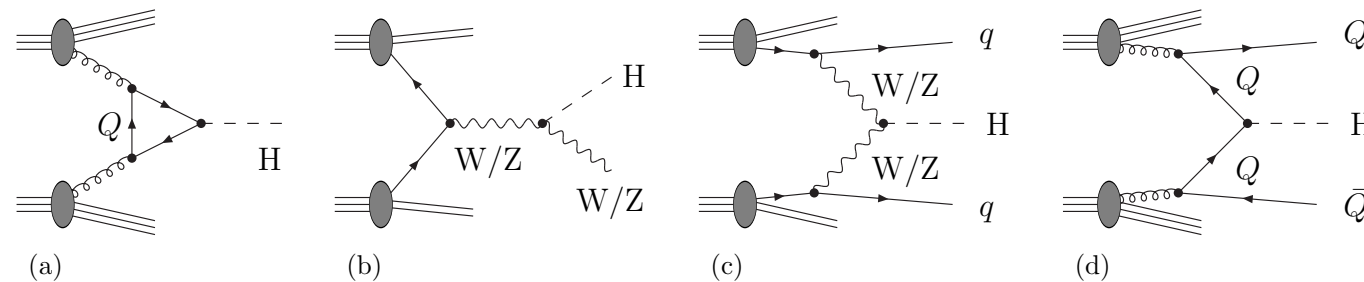
Do LHC results on the SUSY particles, Higgs signal strengths and constraints on DM from direct and indirect detection experiments rule out light neutralino DM?

Already many studies exist in literature, I'll not list them here

based on:
arxiv:1308.3735 [hep-ph] (published PLB)



126 GeV Higgs at the LHC

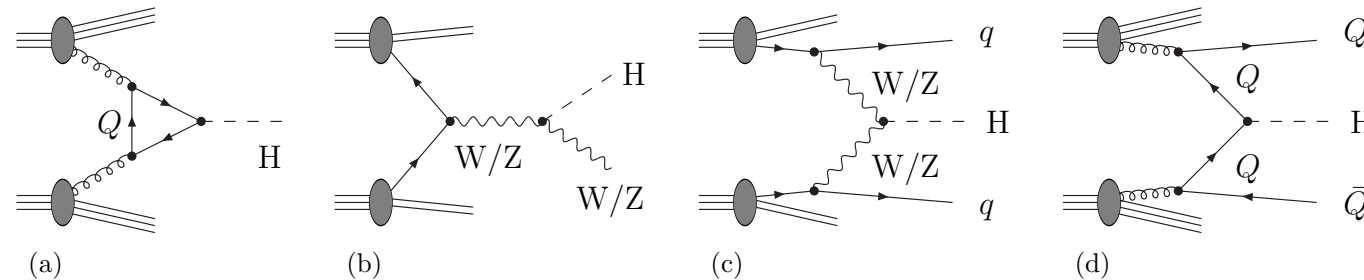


- Four production modes ggF, VH, VBF, and ttH
- Five final states: $\gamma\gamma$, $ZZ^{(*)}$, $WW^{(*)}$, $b\bar{b}$ and $\tau\tau$. only four are independent - ZZ and WW related by custodial symmetry
- Loop induced ggF production and $\gamma\gamma$ final state are susceptible to BSM contributions - in case of SUSY light staus and neutralino contribute
- Experimentally we get information on the signal strengths

$$\mu_i \equiv (\sigma \times \text{BR})_i / (\sigma \times \text{BR})_i^{\text{SM}},$$

for each final state

126 GeV Higgs at the LHC



- Higgs effective Lagrangian:

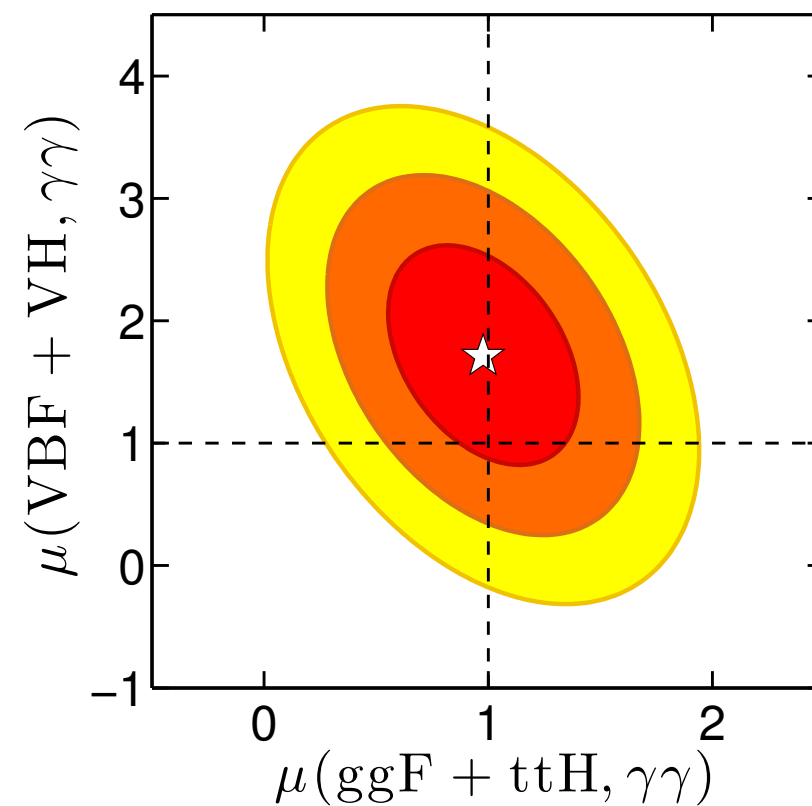
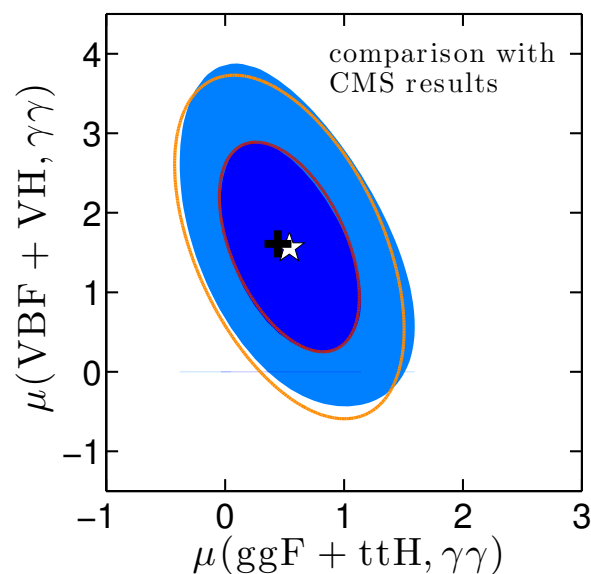
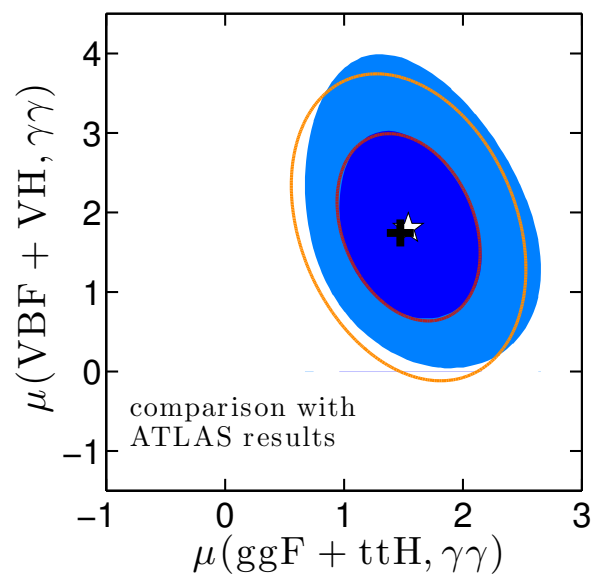
$$\mathcal{L} = g \left[C_V \left(M_W W_\mu W^\mu + \frac{M_Z}{\cos \theta_W} Z_\mu Z^\mu \right) - C_U \frac{m_t}{2M_W} \bar{t}t - C_D \frac{m_b}{2M_W} \bar{b}b - C_D \frac{m_\tau}{2M_W} \bar{\tau}\tau \right] H .$$

C's scale couplings relative to SM ones; $C_U=C_D=C_V=1$ is SM.

- Additional loop contribution modify the couplings to gluons and photons

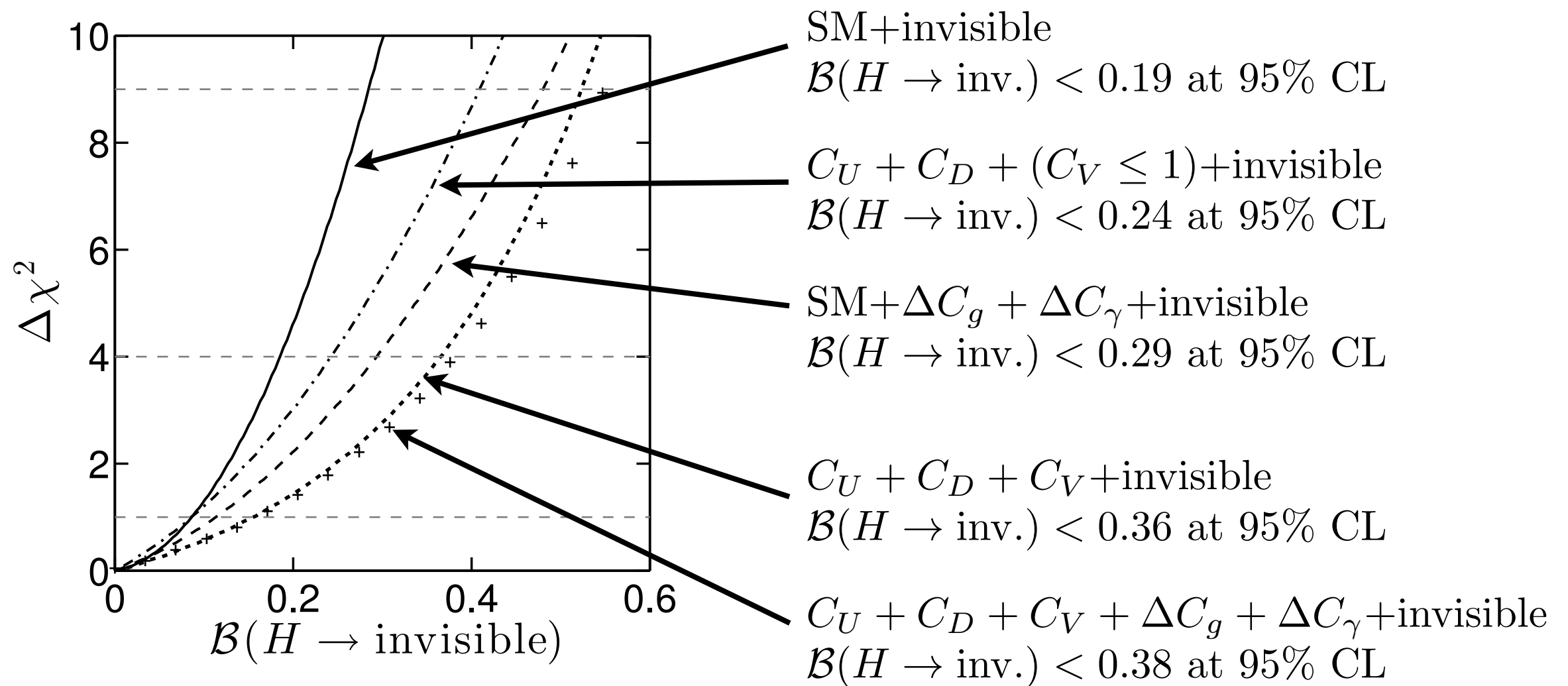
126 GeV Higgs at the LHC

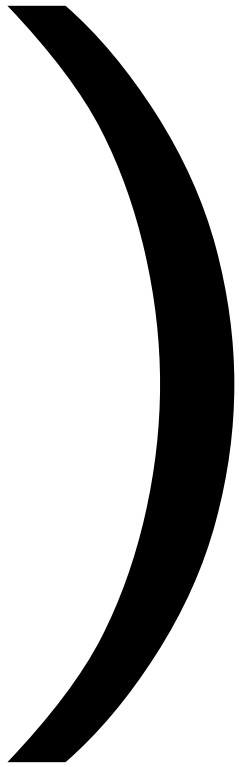
- A combined likelihood in (ggF+ttH) and (VBF+VH) planes was derived using ATLAS, CMS and Tevatron results



126 GeV Higgs at the LHC

- How much invisible Higgs decay is allowed?





A real life application

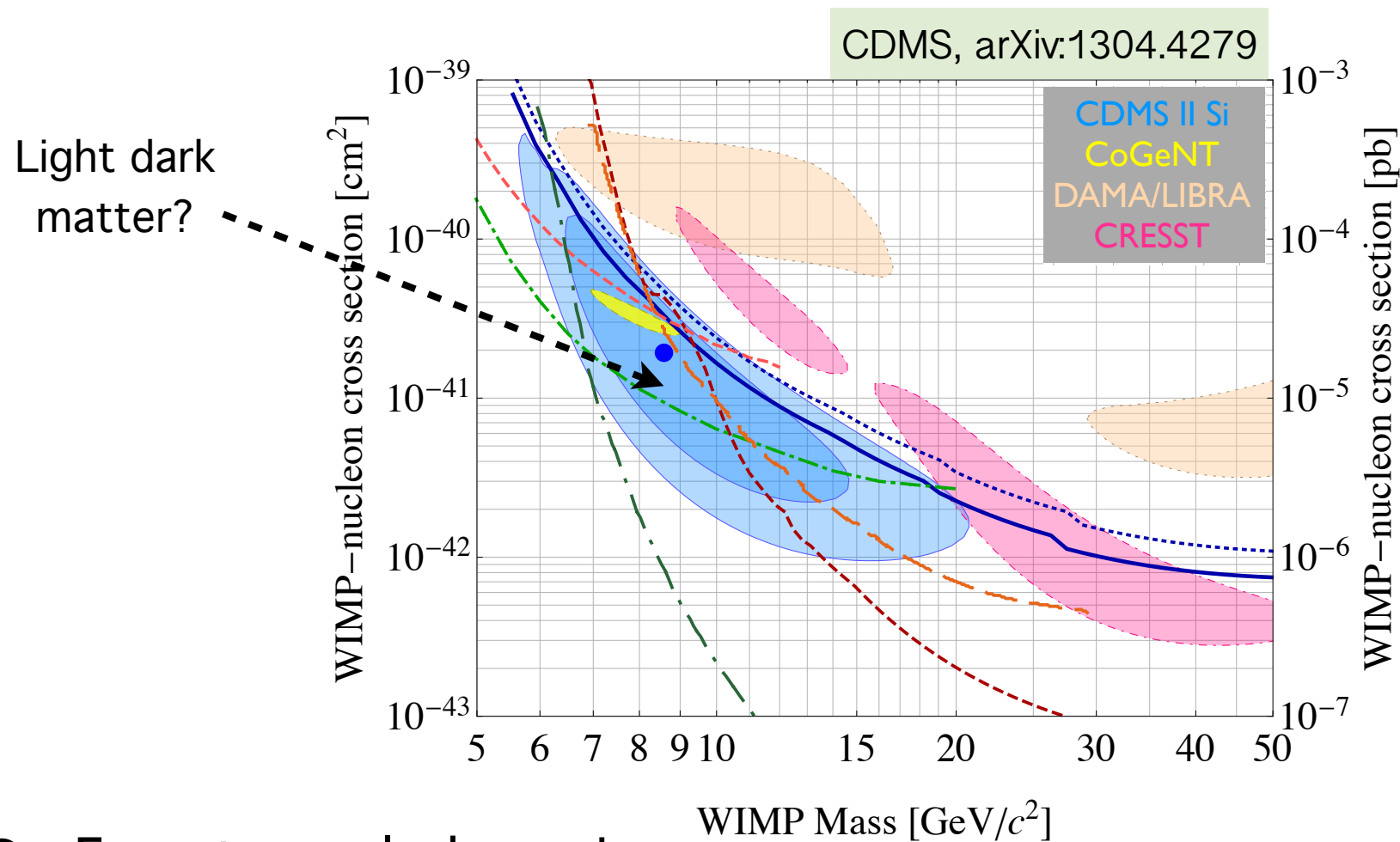
Do LHC results on the SUSY particles, Higgs signal strengths and constraints on DM from direct and indirect detection experiments rule out light neutralino DM?

Already many studies exist in literature, I'll not list them here

based on:
arXiv:1308.3735 [hep-ph] (published PLB)

Why light neutralino?

- Light SUSY spectrum
- Hints from direct (and may be indirect detection) ~ 10 GeV

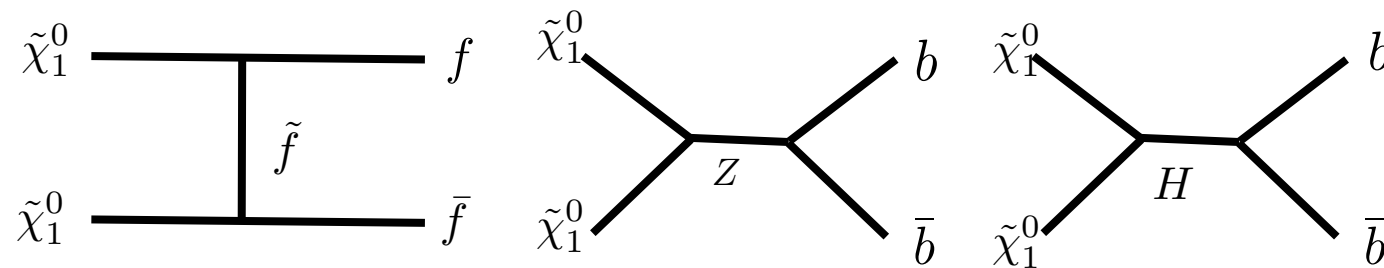


- Easy to exclude region
 - No resonance below 45 GeV ($M_Z/2$)
 - No co-annihilation under 100 GeV (LEP limits) (counter example light sbottoms) [arXiv:1308.2153](#)

How light is light?

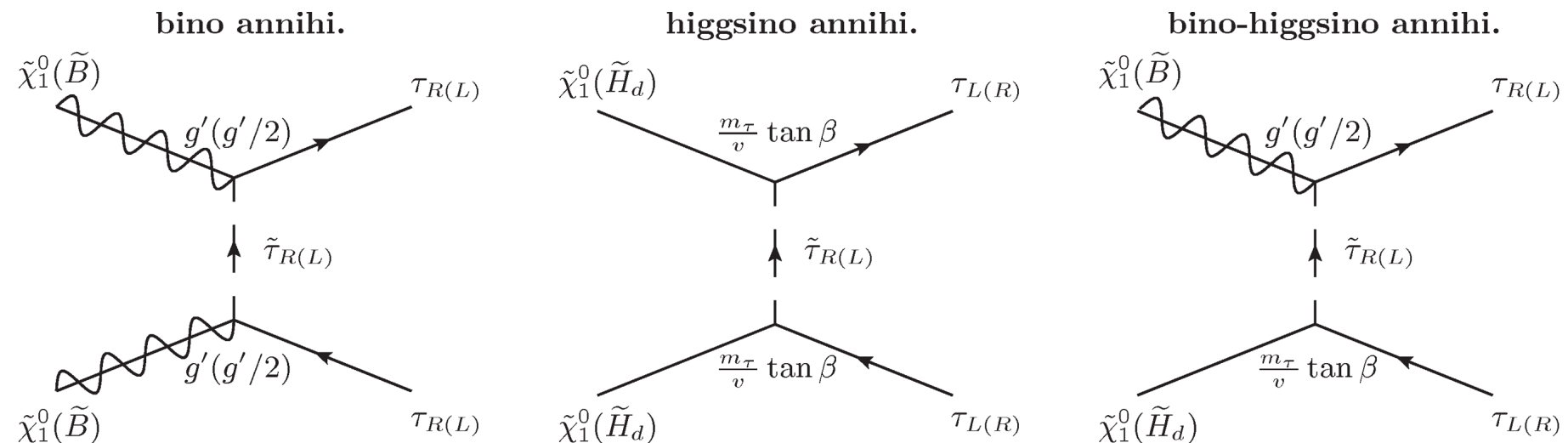
- Relaxing gaugino universality: few collider constraints
 - Z width, LEP bounds, invisible Higgs decays

- Most important annihilation channels:



- Region of interest: $m_{\tilde{\chi}_1^0} < m_h/2$ - Z and H exchange not effective
- Light slepton exchange of interest to us here

How light is light?



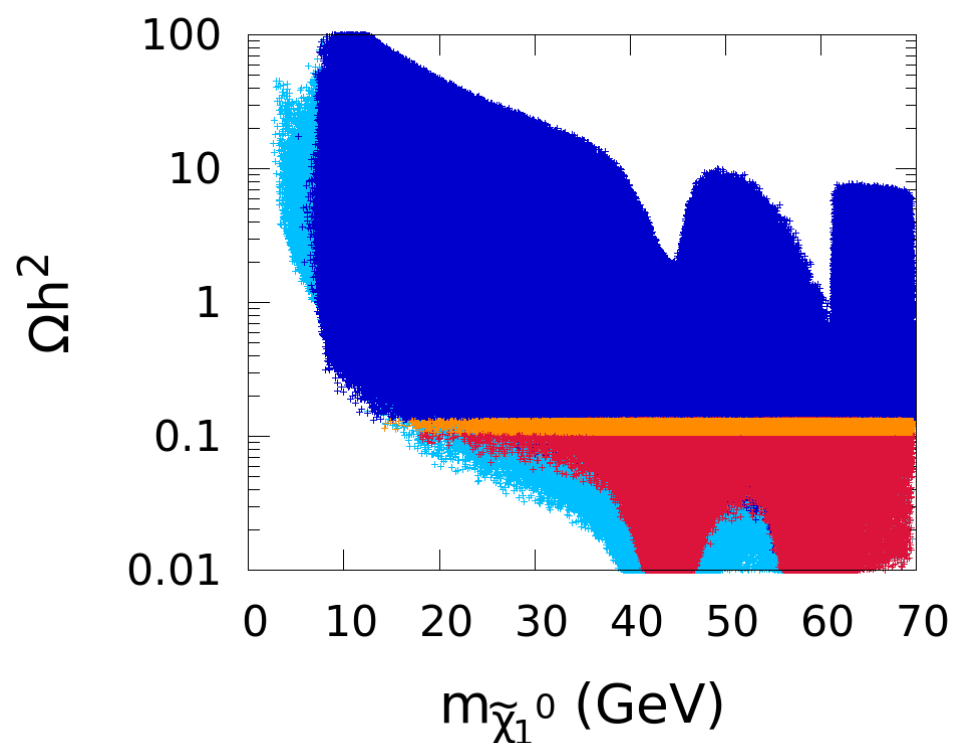
- RH stau annihilation is more efficient, also get enhancements for high $\tan(\beta)$ and higgsino LSP

Light chargino	LEP and LHC	✓
Invisible Z, Higgs decays	LEP and LHC	✓
Light neutralino 2	LEP $\sigma(e^+ e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_2^0)$	✓
Slepton and stau	LEP and LHC	✓

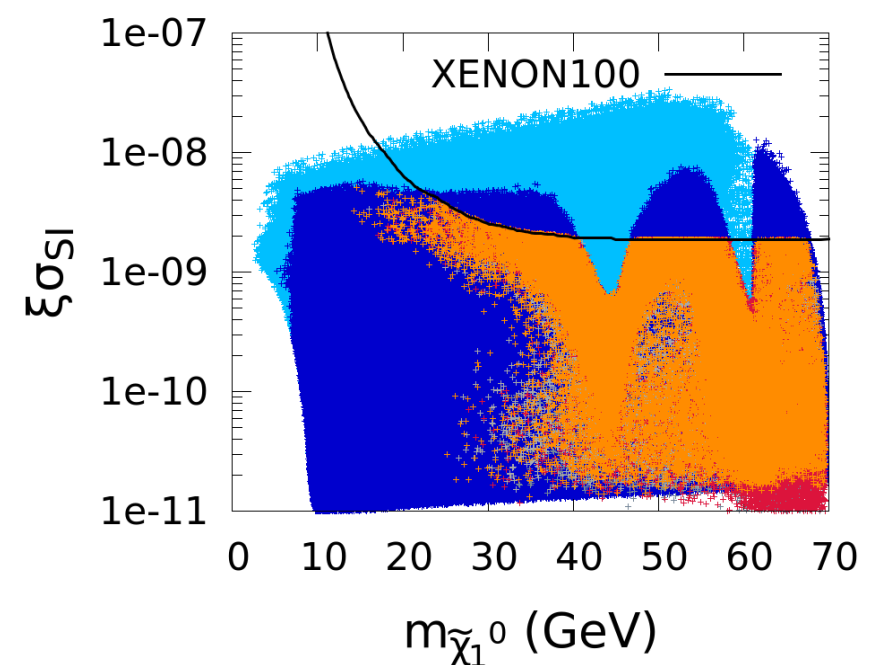
How light is light?

- pMSSM scan over 11 parameters

- $M_1, M_2, \mu, \tan\beta, M_A, A_t, M_{1L}, M_{1R}, M_{3L}, M_{3R}, A_\tau$
- LEP limits, Z width, flavor physics, heavy Higgs searches @LHC, Higgs mass, Higgs couplings, Xenon100



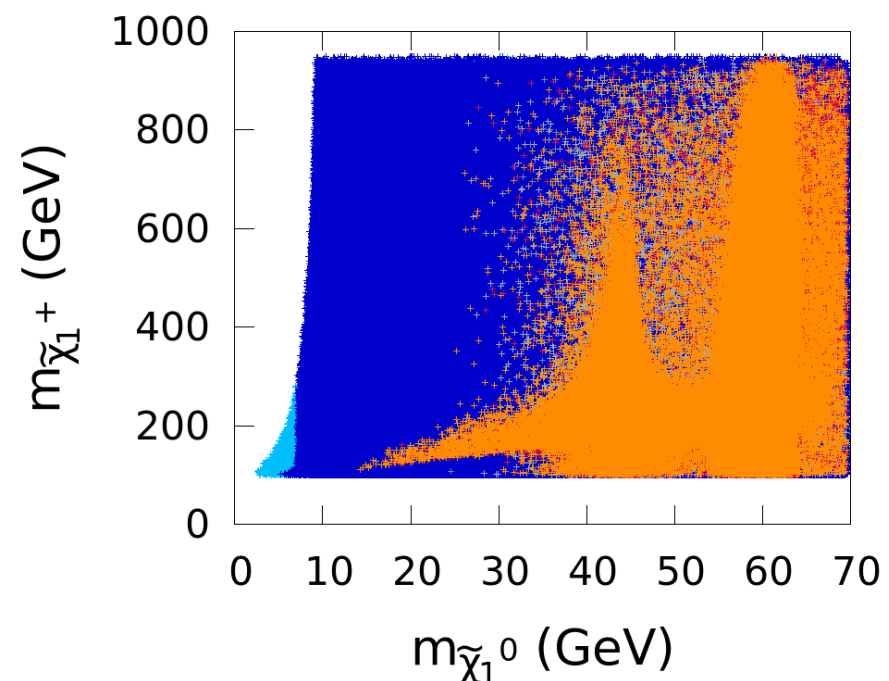
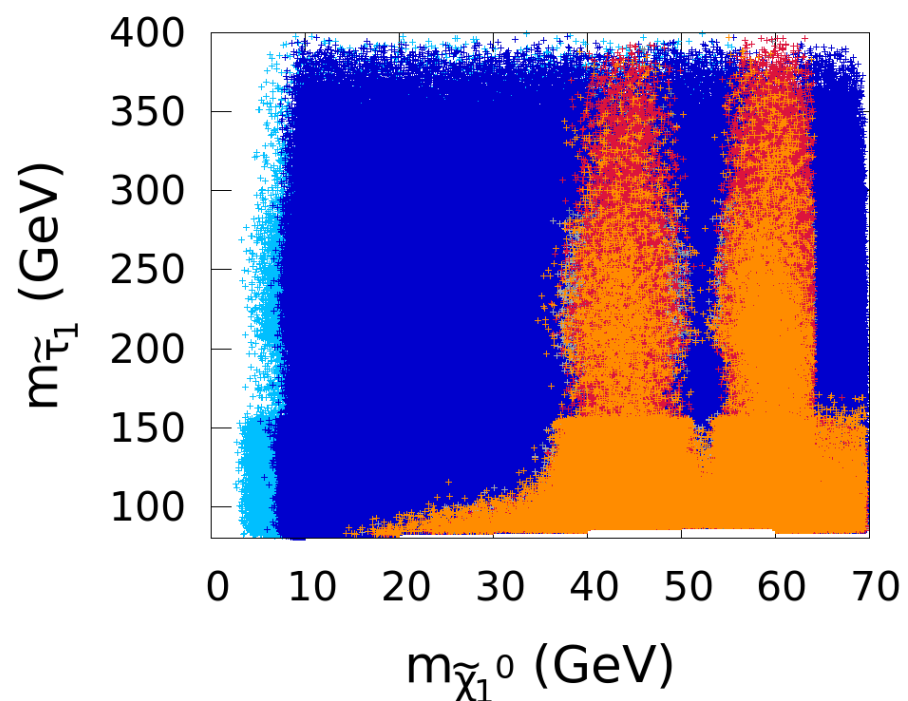
Basic constraints
Higgs couplings fits



LHC results + upper limit of relic
LHC results + exact relic

How light is light?

- DM < 35 GeV associated with light stau + light chargino



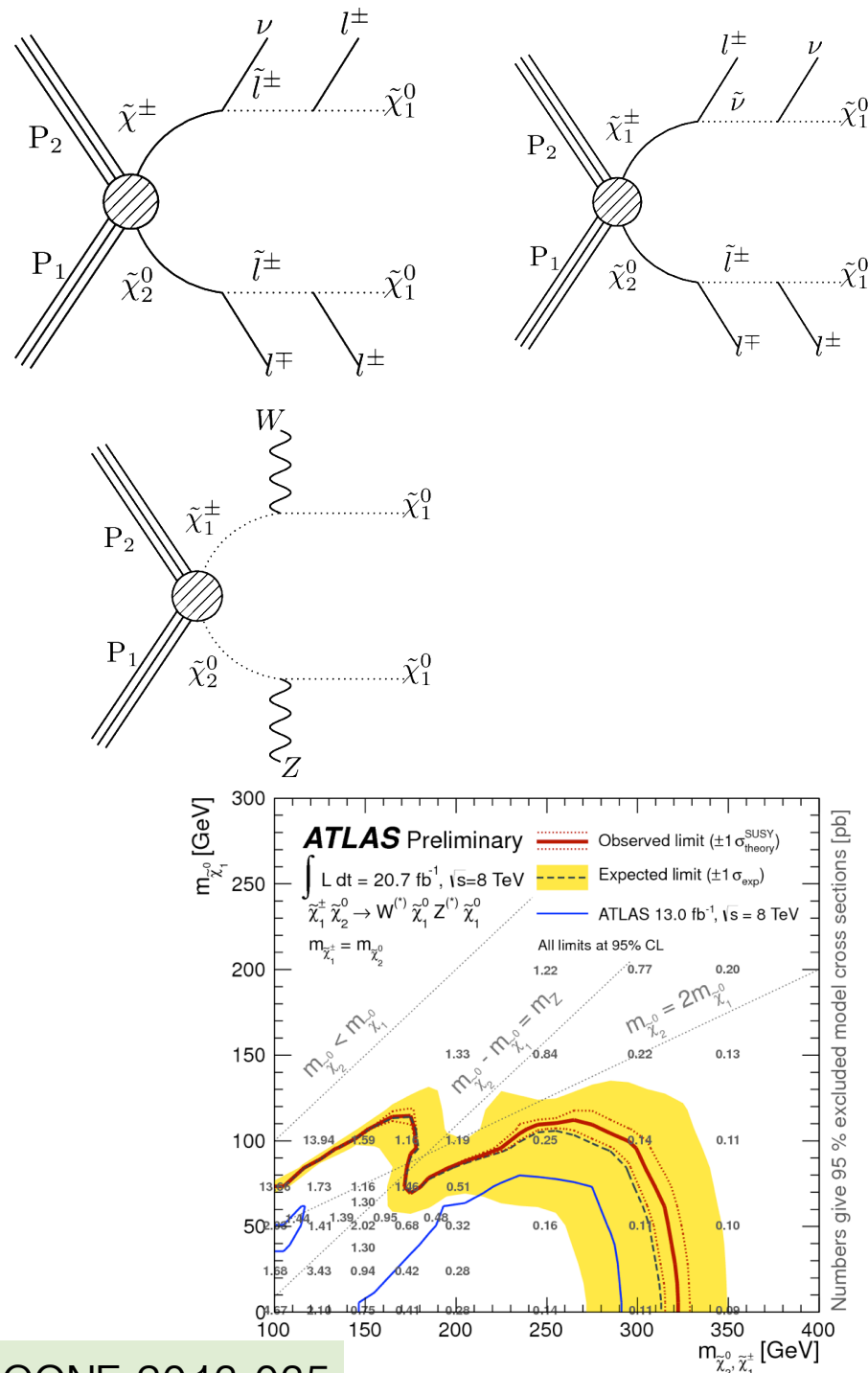
- LHC searches put constraints on light electroweak -ino and slepton production

Basic constraints
Higgs couplings fits

LHC results + upper limit of relic
LHC results + exact relic

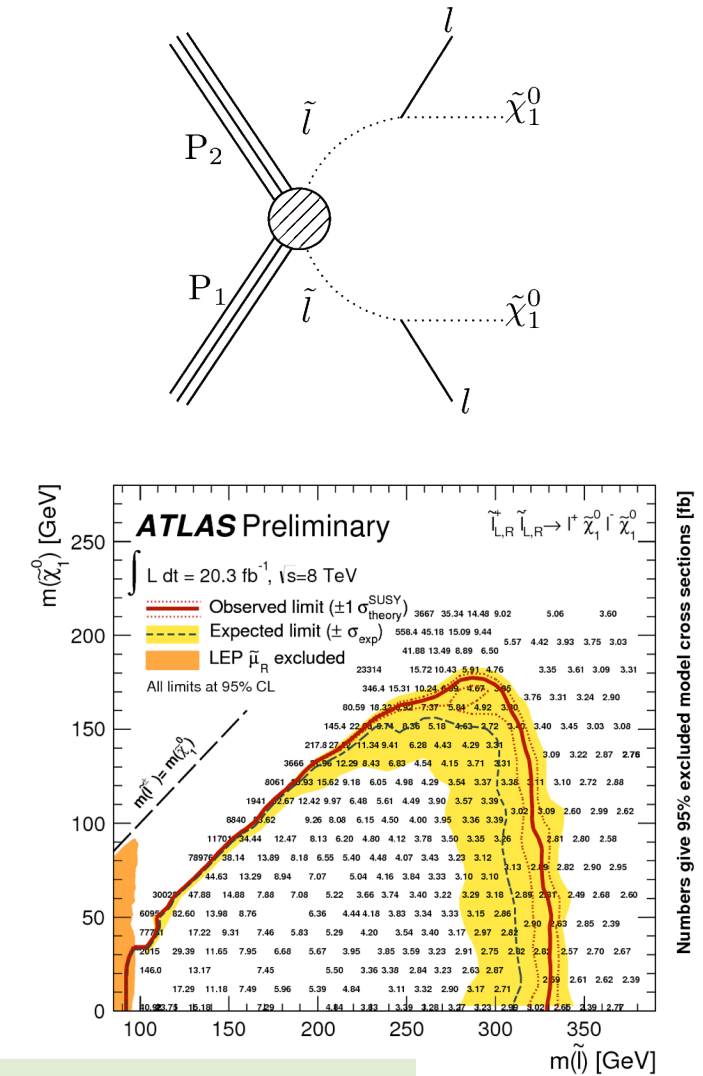
LHC searches

● Direct electroweak -ino production



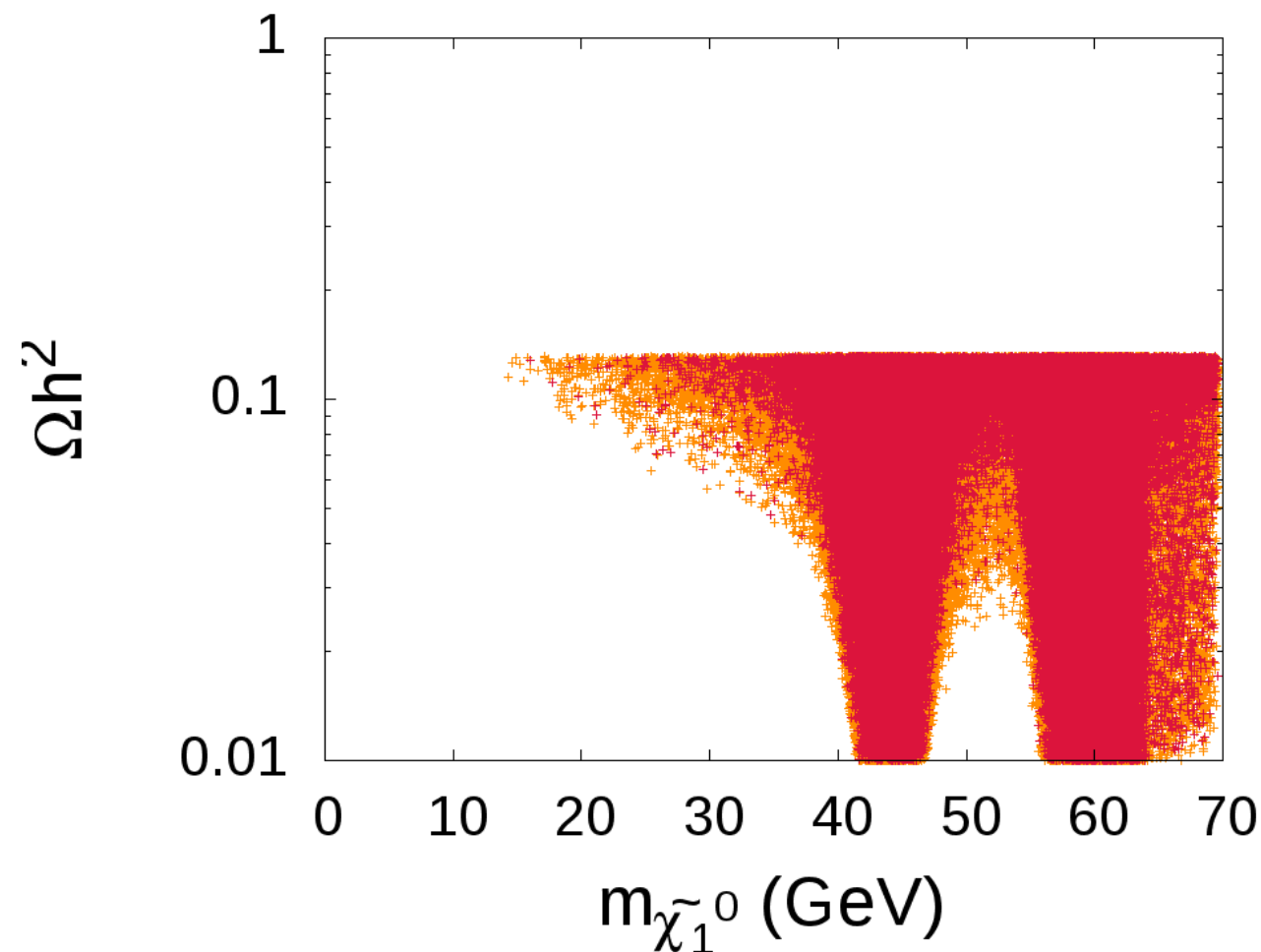
ATLAS-CONF-2013-035

● Direct slepton production



ATLAS-CONF-2013-049

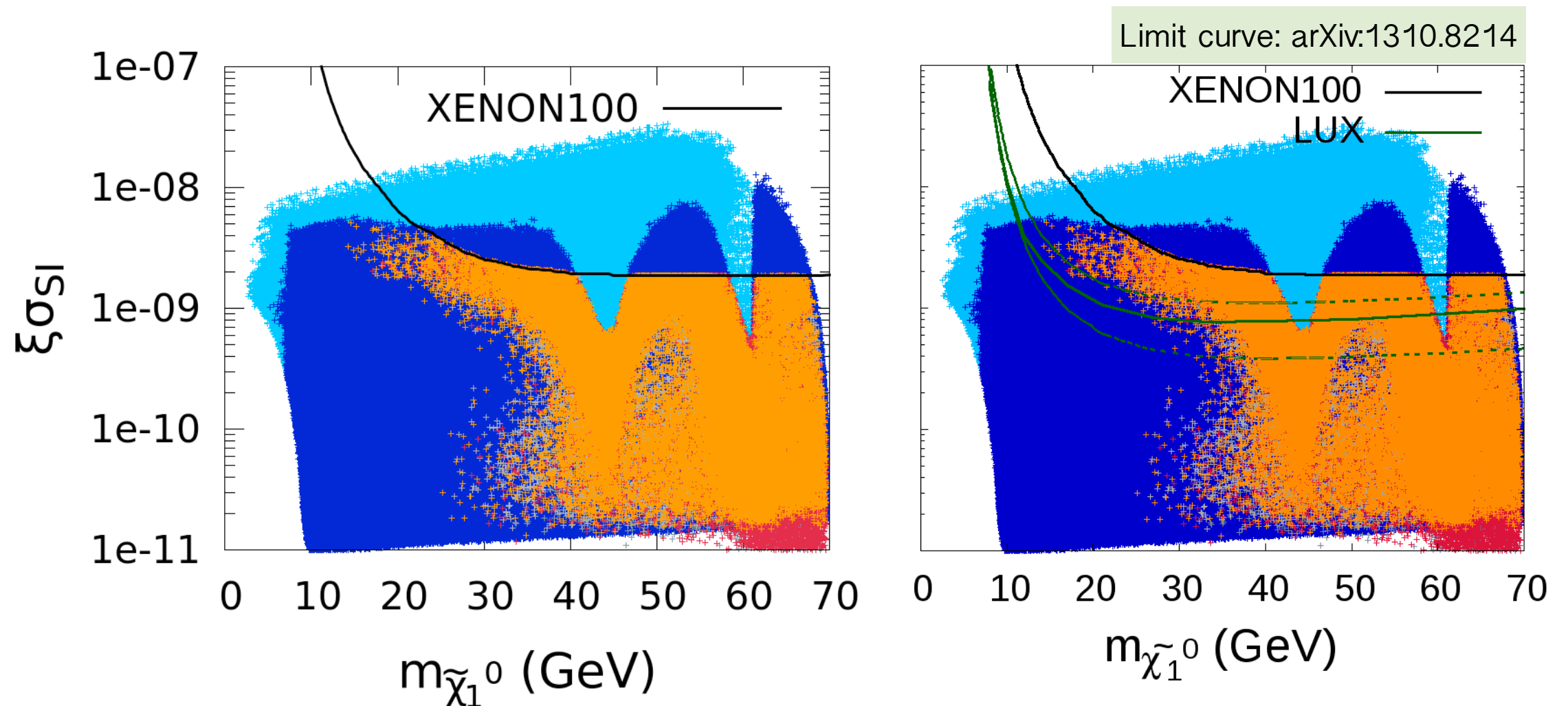
Applying SModelS



All points passing relic
density upper limits
Points excluded by the
LHC limits

- SMS results used from ATLAS-CONF-2013-049, CMS-PAS-SUS-12-022, ATLAS-CONF-2013-035
- Density of points reduced - LHC SMS results do rule out some scenarios
- In general light neutralino still possible

LUX limits



- Neutralino DD CS is driven by higgsino component, suppressed when LSP has small higgsino component
- LUX disfavors the light neutralino DM region we had identified to be viable

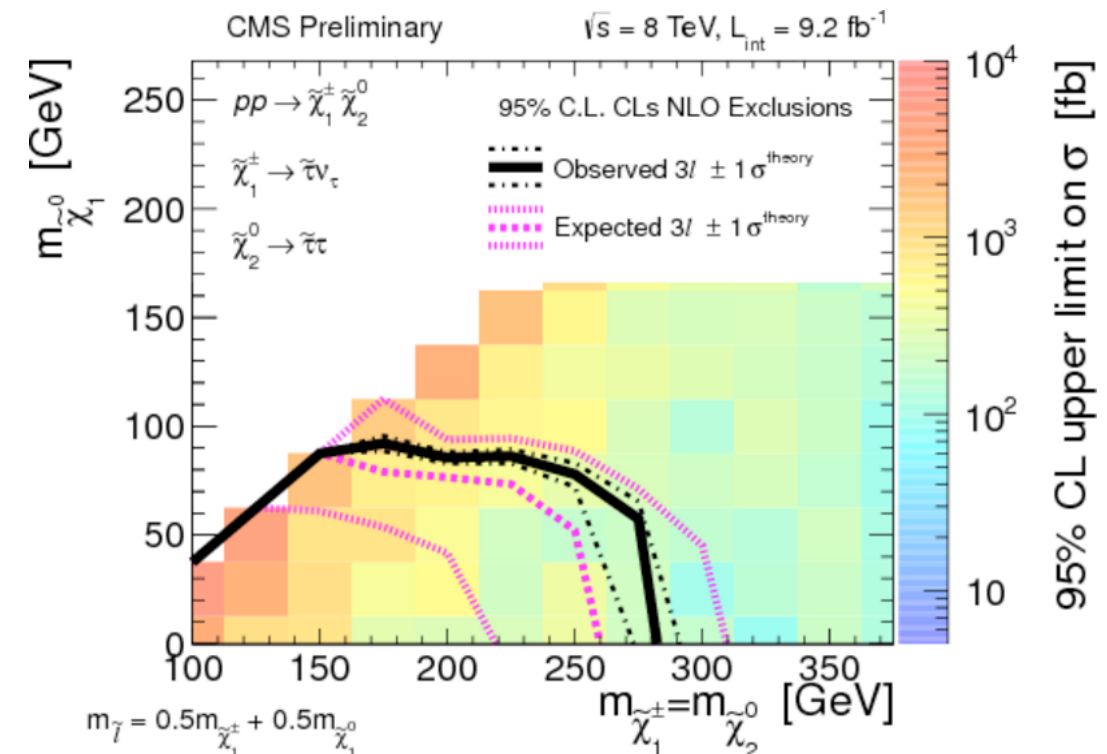
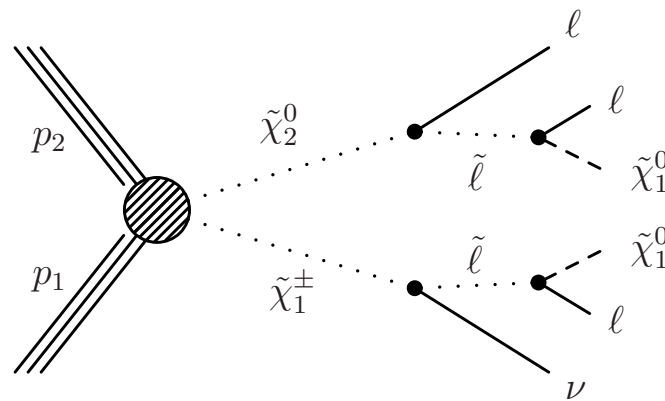
Basic constraints
Higgs couplings fits

LHC results + upper limit of relic
LHC results + exact relic

SMS approach - what's next?

- SMS approach is not perfect yet
- Not all SMS topologies are present
- It is not always possible to use experimental SMS results, sometimes the results have a coarse grid or in case of a one step decay, only one mass slice is given

CMS - SUS -12-022

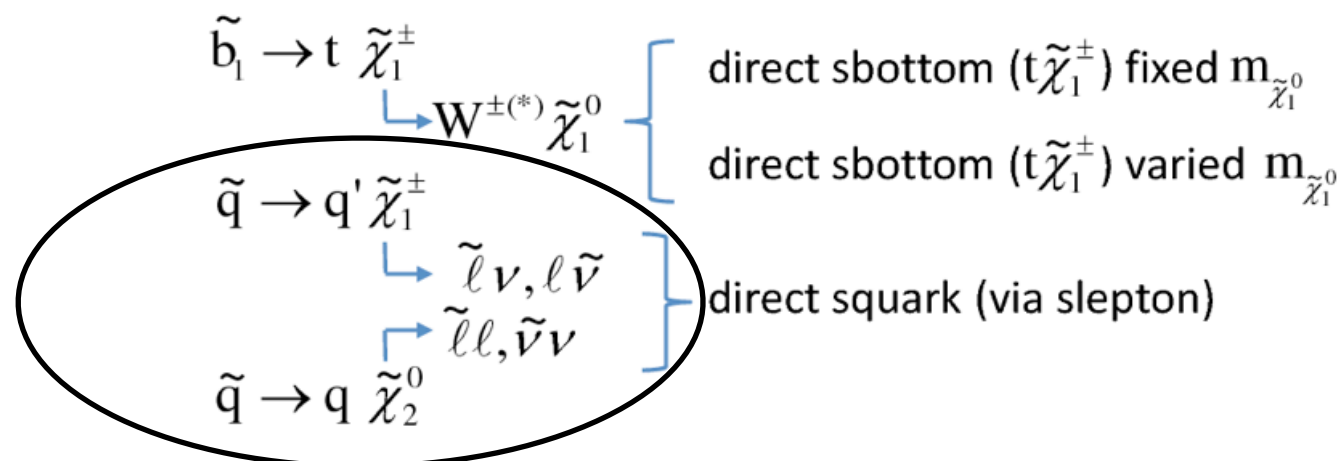
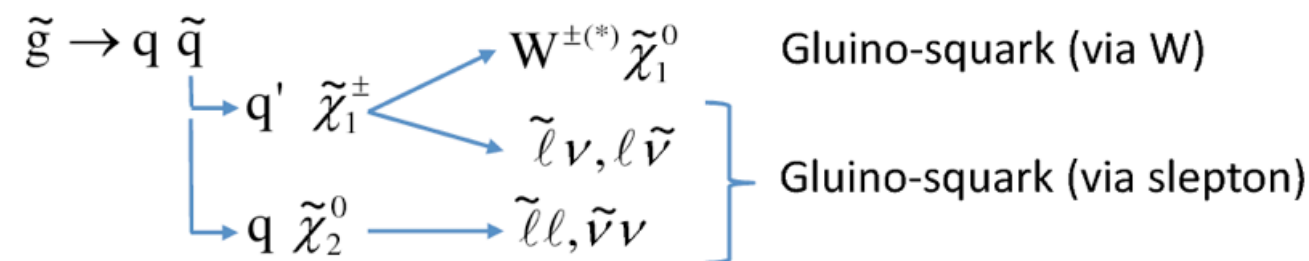
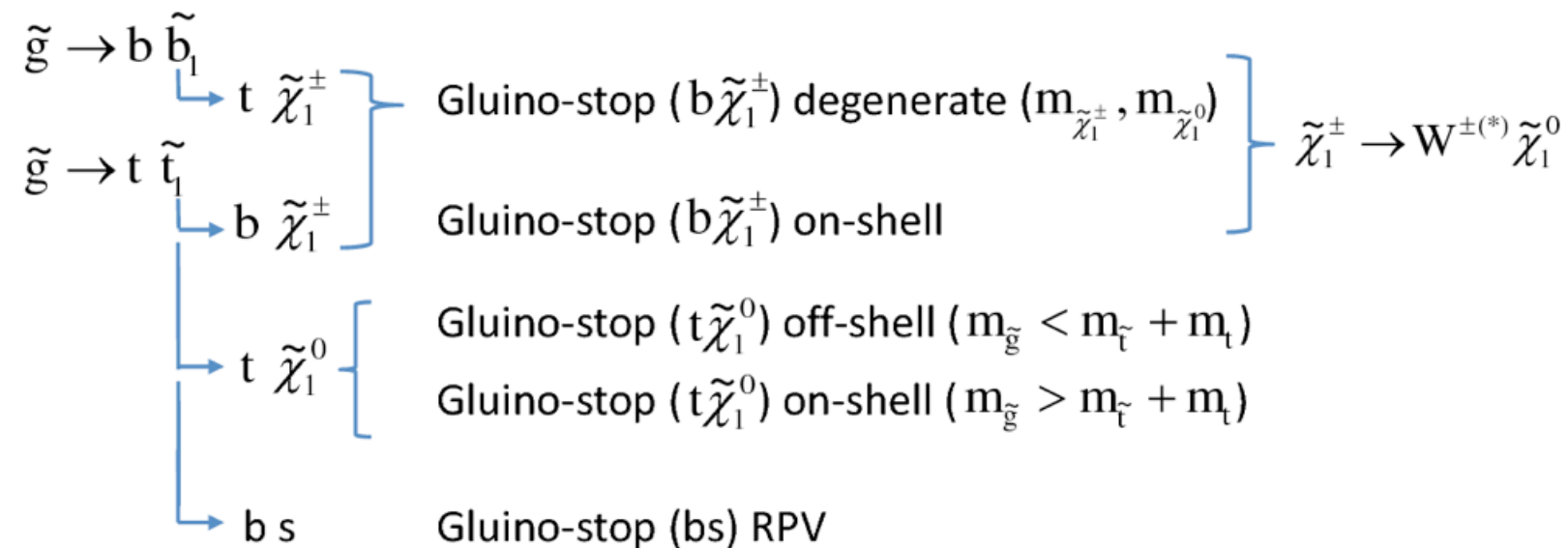


SMS approach - what's next?

- Results presented are not always usable
- My nightmare SMS analysis: ATLAS-CONF-2013-007
 - Involves topologies with more than four SUSY particles
 - Plots often include strong assumptions on the masses involved
 - Binning is not uniform

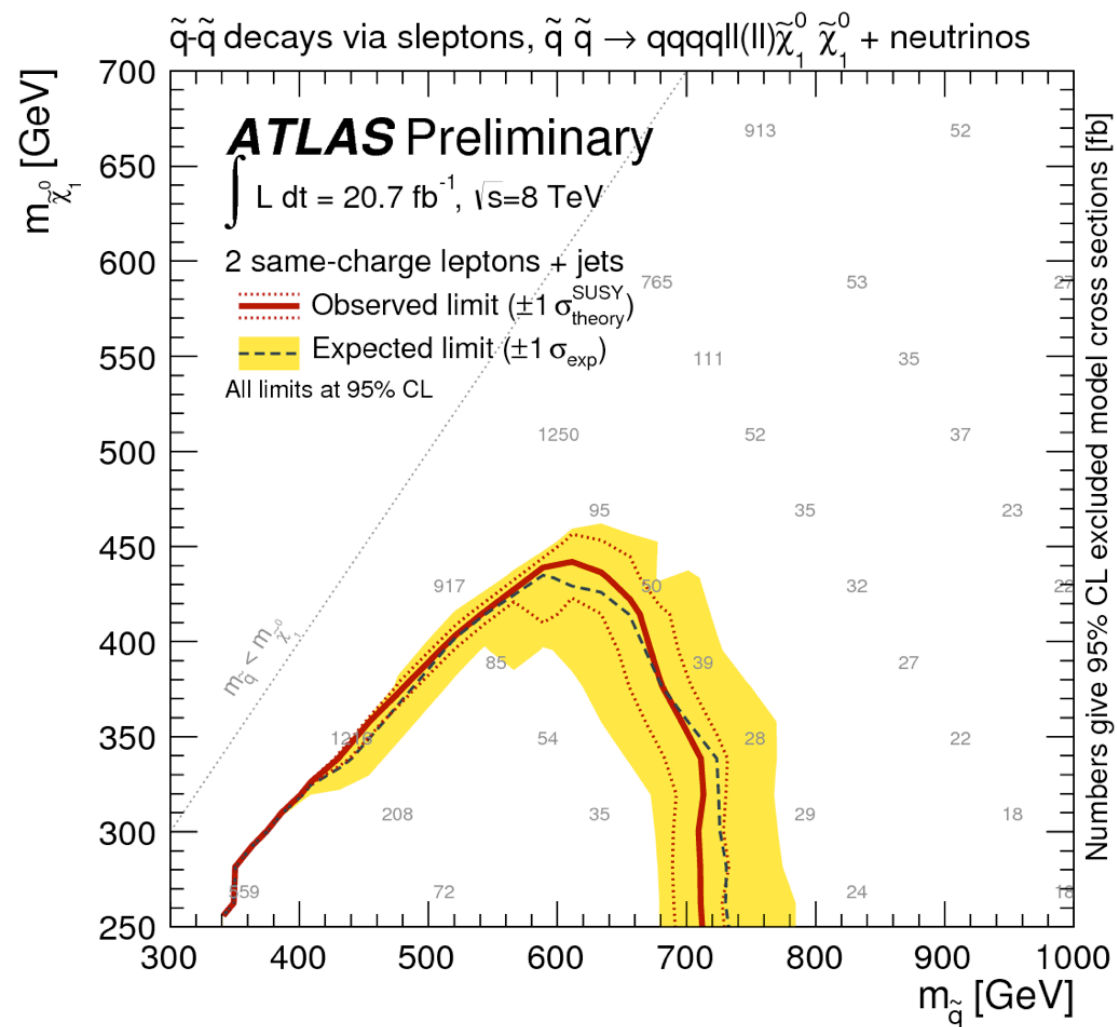
SMS approach - what's next?

ATLAS-CONF-2013-007



SMS approach - what's next?

ATLAS-CONF-2013-007

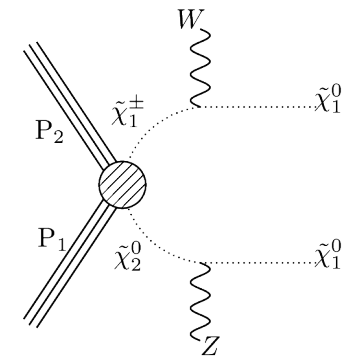
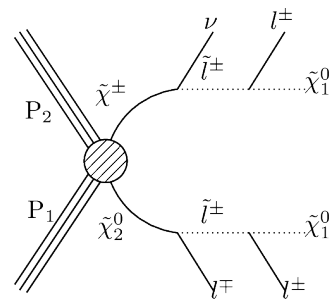
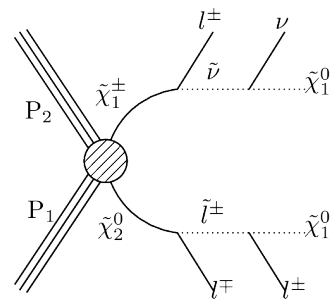


- Two of the four masses fixed
- Only democratic slepton decays
- Irregular binning (less severe)

SMS approach - what's next?

- In principle several topologies can contribute to the same final state with different efficiencies

Tri-lepton final state: ATLAS-CONF-2013-035

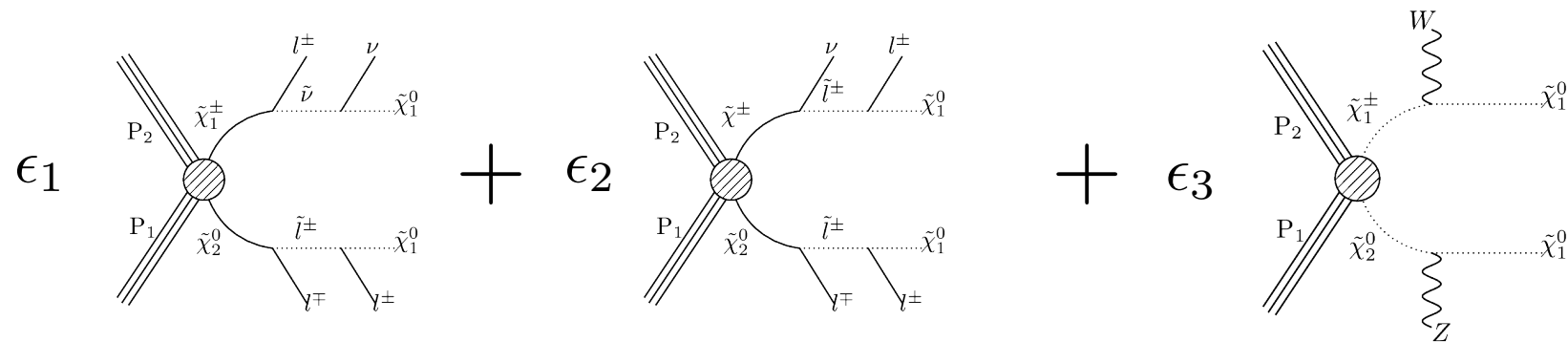


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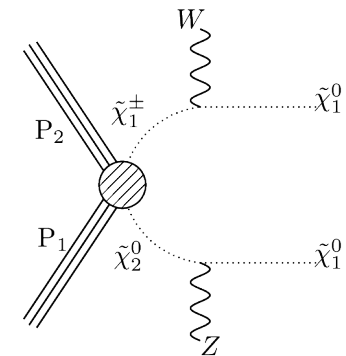
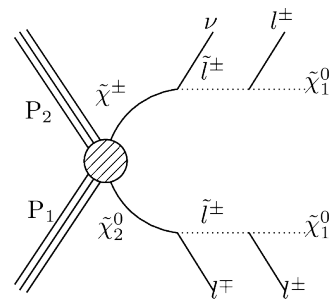
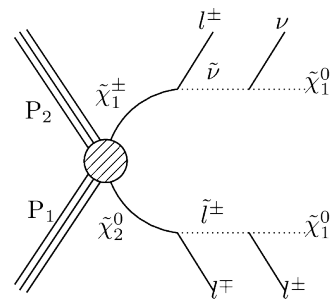
Real life



SMS approach - what's next?

- In principle several topologies can contribute to the same final state with different efficiencies

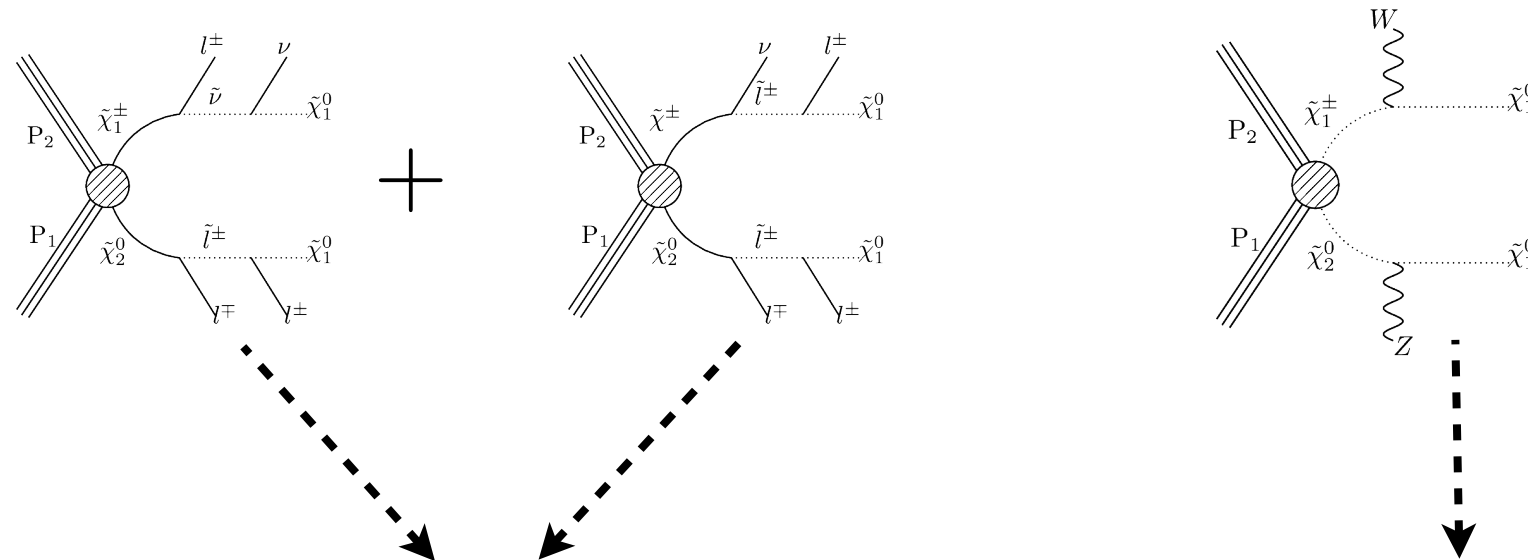
Tri-lepton final state: ATLAS-CONF-2013-035



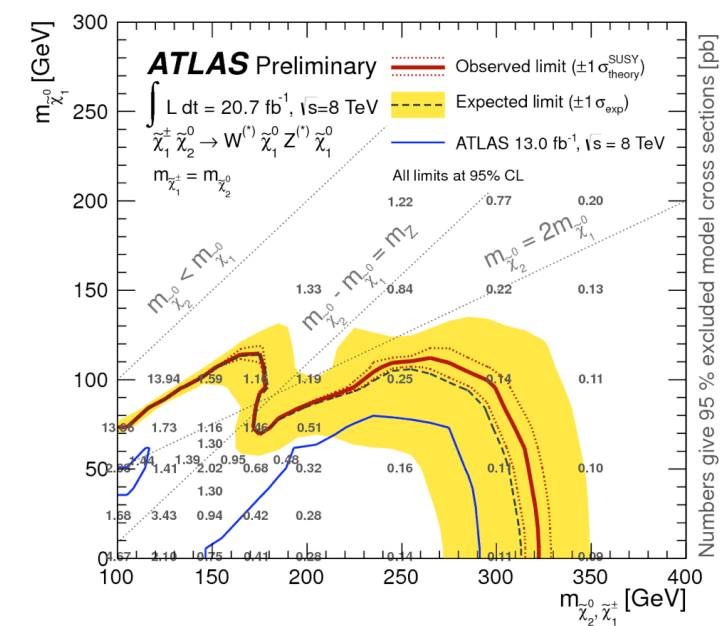
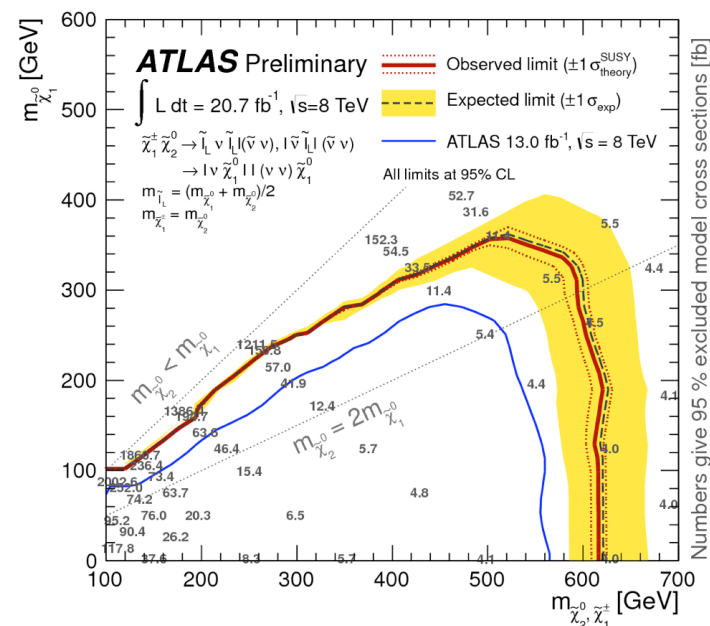
SMS approach - what's next?

- In principle several topologies can contribute to the same final state with different efficiencies

Tri-lepton final state: ATLAS-CONF-2013-035



Current SMS results



SMS approach - what's next?

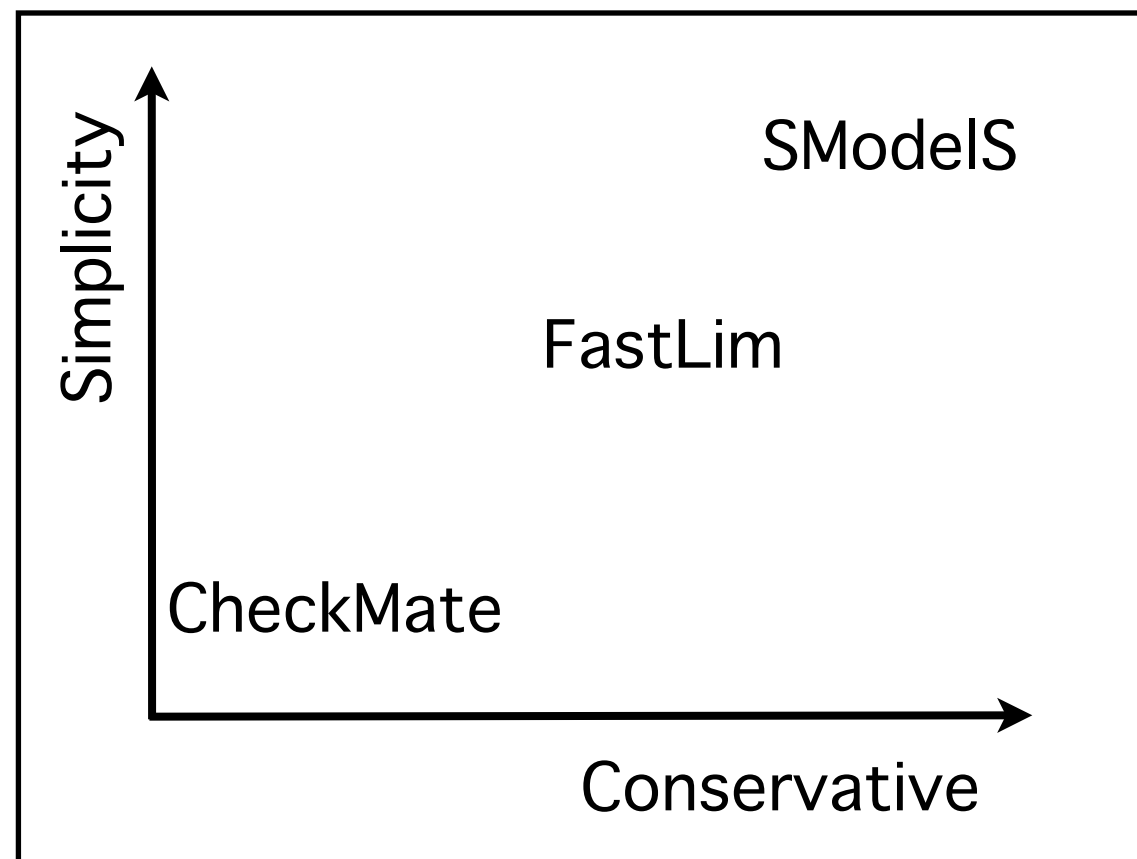
- To utilize this approach one needs to develop efficiency maps for each analysis and for each topology that can potentially contribute to the final state
- Need to re-implement the analysis - requires manpower and availability of information from experimental collaborations in a systematic manner
- FastLim is developing the efficiency maps, aim is to reconstruct the number of events for each signal region
- SModelS is also capable of supporting efficiency maps approach, and in future we might consider exploiting this feature

SMS approach - what's next?

- Generally the development will be slow - a community effort to contribute to the efficiency maps is underway, this effort should also make the re-implemented analysis publicly available
- Also need to develop a reasonable likelihood in order to be able to combine several signal regions - e.g. CMS-SUS-PAS-13-011 has 16 signal regions

SMS approach - what's next?

- It will be difficult to tackle long cascade decays with SMS approach
- A completely different approach is being taken by checkMATE
- The tool identifies the most sensitive topology and then tests it via Monte Carlo simulation

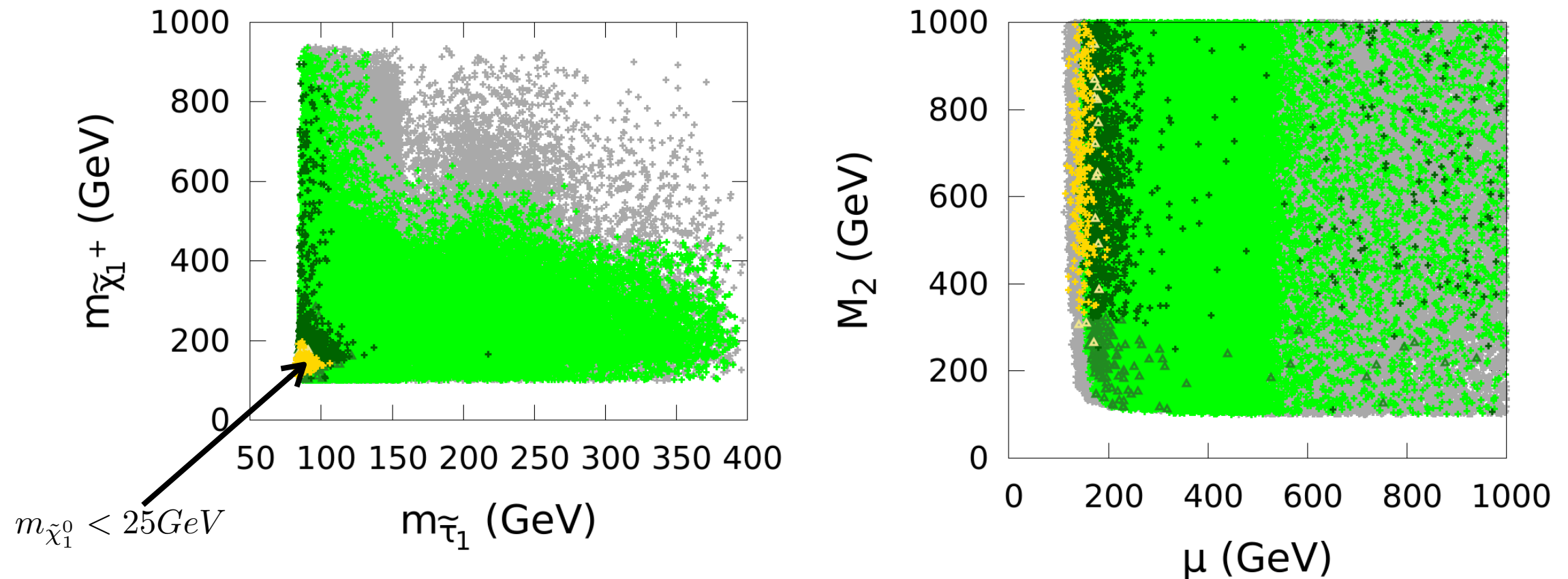


Conclusions

- SMS results are a good way to test BSM theories and can have a good constraining power
- SModelS is designed to utilize this power and constrain BSM scenarios
- The formalism of the code is generic and can be applied to any BSM spectra for which SMS results are applicable
- Currently, the code can handle scenarios with Z_2 parity
- It contains the most comprehensive database of SMS SUSY results
- It can be used in order to understand the features of parameter space under consideration, it can also be used to study viability of an interesting BSM scenario
- Stay tuned applying LHC searches to your favorite BSM model is being made easy!

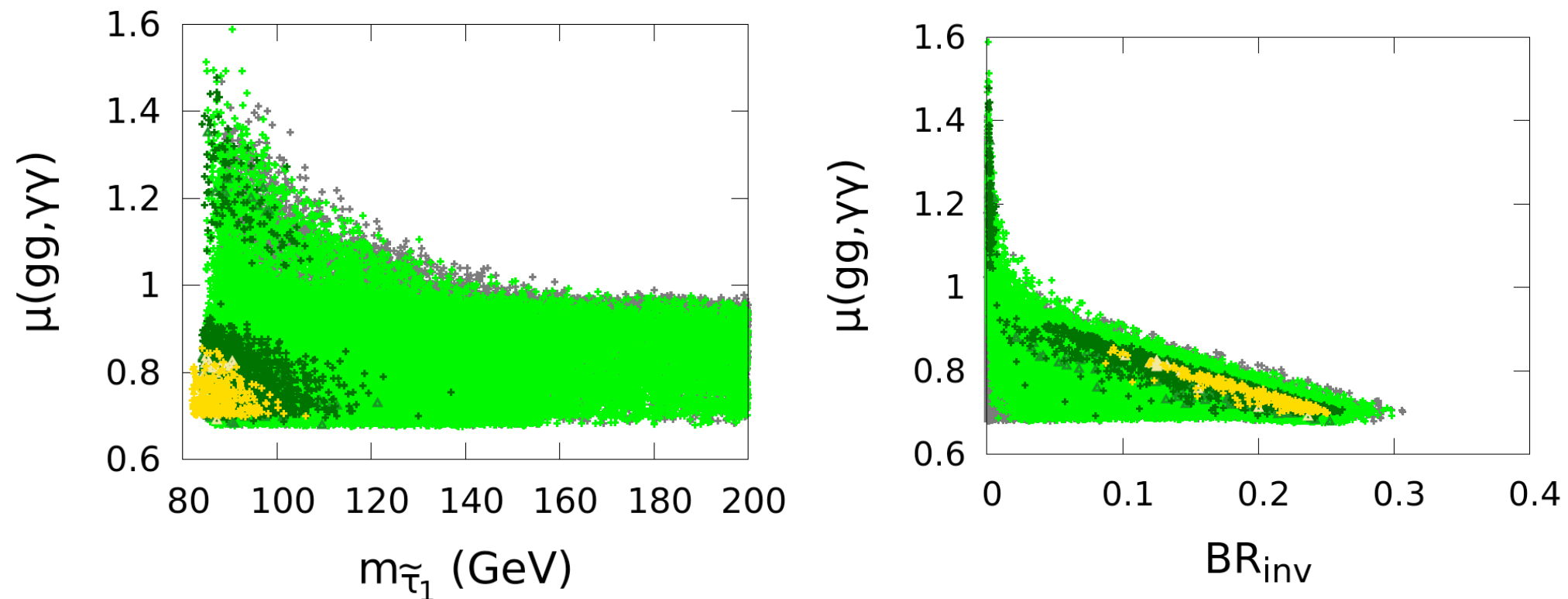
Back-up

Applying SModelS



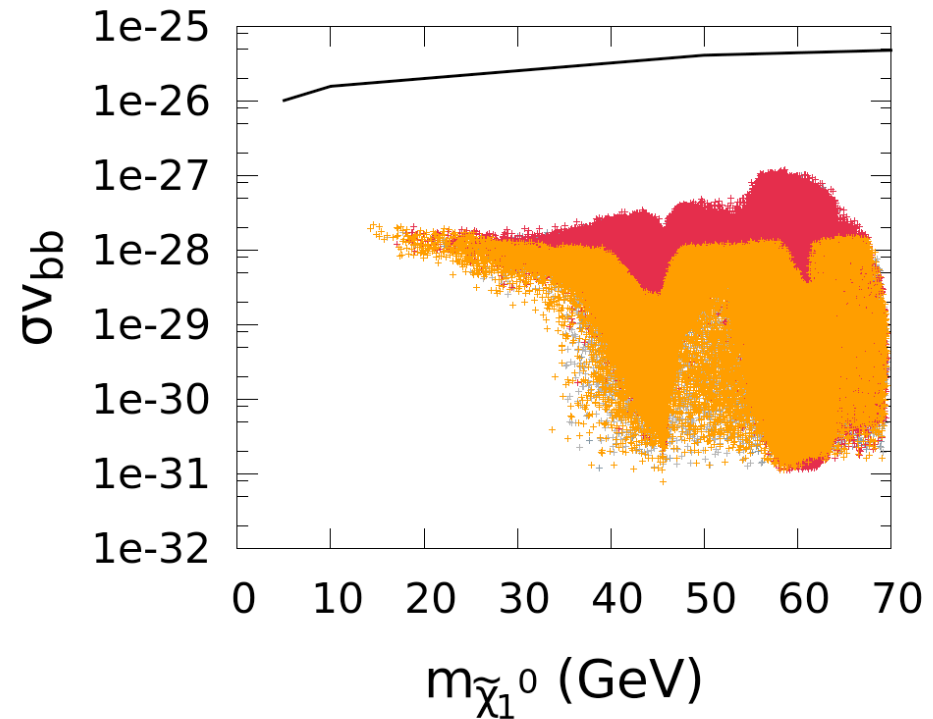
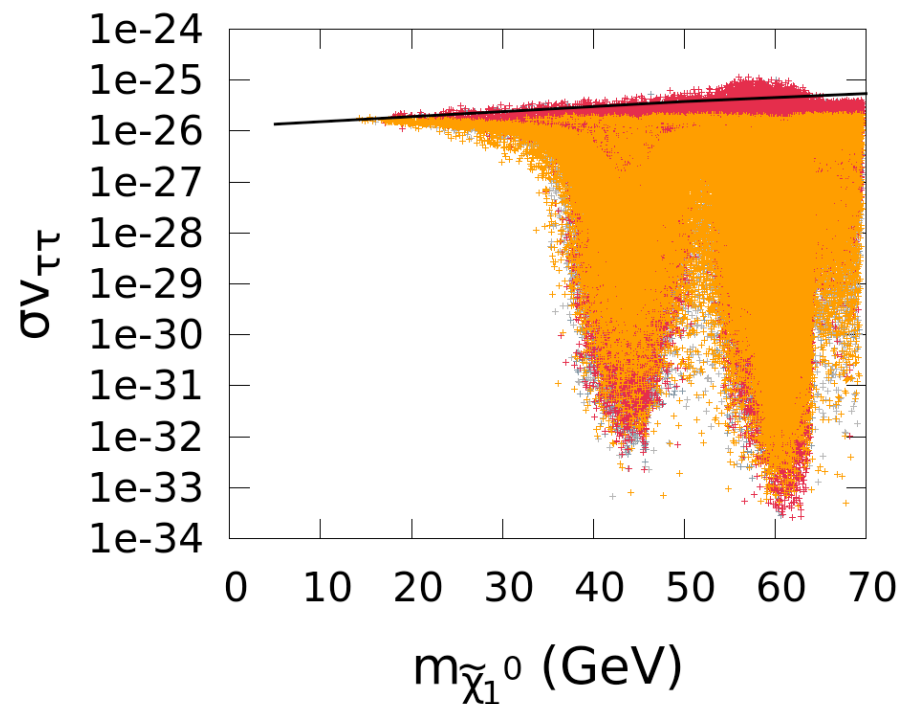
- Mass slices 10-15; 25-35; 35-50; 50-60 GeV
- Density of points reduced - LHC does rule out some scenarios
- In general light neutralino still possible

Higgs signal strengths



- Mass slices 10-15; 25-35; 35-50; 50-60 GeV
- Light neutralino and light stau lead to modifications
- Heavily mixed stau increases diphoton rate
- Lightest neutralino associated with some invisible Higgs decays
- At 14 TeV with $ZH \rightarrow$ invisible, better sensitivity expected

Indirect detection limits



- We test for FERMI-LAT limits - photons produced from DM annihilation in dwarf spheroidal galaxies in $b\bar{b}$ or $\tau\tau$ channel
- Large fraction of LSP < 30 GeV points are several orders of magnitude below the limit

Why not use monojet channel?

- Direct LSP production probed via monojet signature at the LHC
- Limits given on the spin-independent interactions of DM
- Limits applicable for models involving heavy mediators
- Not applicable for MSSM since the mediators are not heavy enough

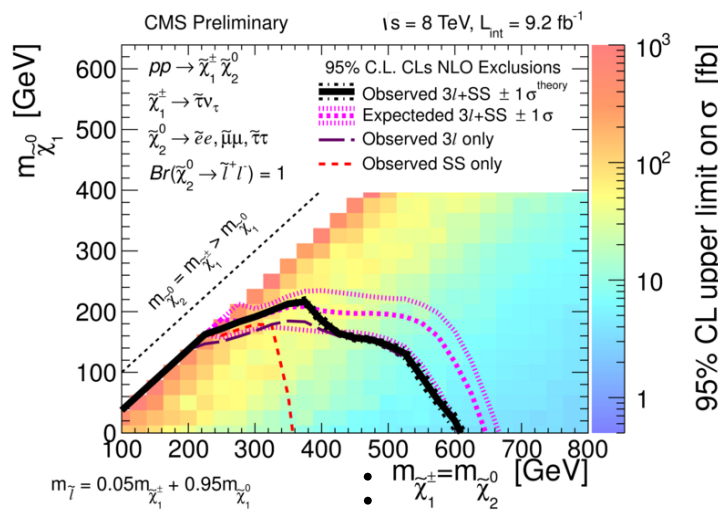
Light neutralino DM - scan details

$\tan \beta$	$[5, 50]$	M_{L_3}	$[70, 500]$
M_{A^0}	$[100, 1000]$	M_{R_3}	$[70, 500]$
M_1	$[10, 70]$	A_τ	$[-1000, 1000]$
M_2	$[100, 1000]$	M_{L_1}	$[100, 500]$
μ	$[100, 1000]$	M_{R_1}	$[100, 500]$

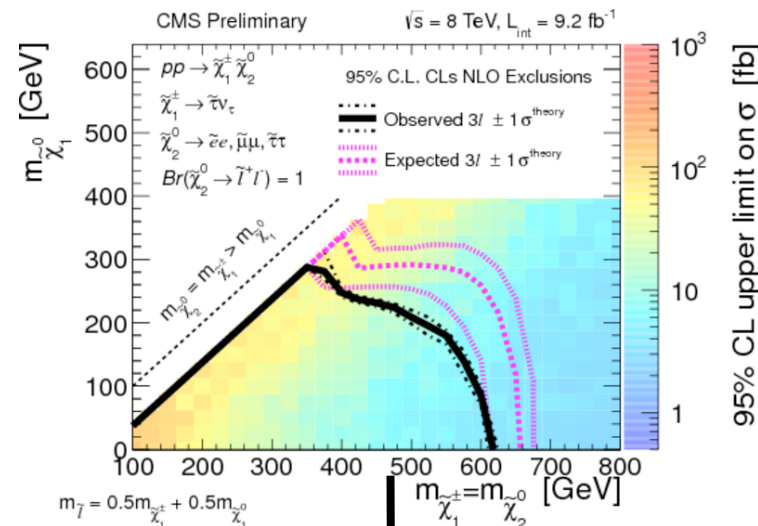
LEP limits	$m_{\tilde{\chi}_1^\pm} > 100 \text{ GeV}$ $m_{\tilde{\tau}_1} > 84 - 88 \text{ GeV (depending on } m_{\tilde{\chi}_1^0})$ $\sigma(e^+e^- \rightarrow \tilde{\chi}_{2,3}^0 \tilde{\chi}_1^0 \rightarrow Z^{(*)}(\rightarrow q\bar{q})\tilde{\chi}_1^0) \lesssim 0.05 \text{ pb}$
invisible Z decay	$\Gamma_{Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0} < 3 \text{ MeV}$
μ magnetic moment	$\Delta a_\mu < 4.5 \times 10^{-9}$
flavor constraints	$\text{BR}(b \rightarrow s\gamma) \in [3.03, 4.07] \times 10^{-4}$ $\text{BR}(B_s \rightarrow \mu^+\mu^-) \in [1.5, 4.3] \times 10^{-9}$
Higgs mass	$m_{h^0} \in [122.5, 128.5] \text{ GeV}$
$A^0, H^0 \rightarrow \tau^+\tau^-$	CMS results for $\mathcal{L} = 17 \text{ fb}^{-1}$, m_h^{max} scenario
Higgs couplings	ATLAS, CMS and Tevatron global fit, see text
relic density	$\Omega h^2 < 0.131$ or $\Omega h^2 \in [0.107, 0.131]$
direct detection	XENON100 upper limit
indirect detection	Fermi-LAT bound on gamma rays from dSphs
$pp \rightarrow \tilde{\chi}_2^0 \tilde{\chi}_1^\pm$ $pp \rightarrow \tilde{\ell}^+ \tilde{\ell}^-$	Simplified Models Spectra approach, see text

Tau dominated scenario

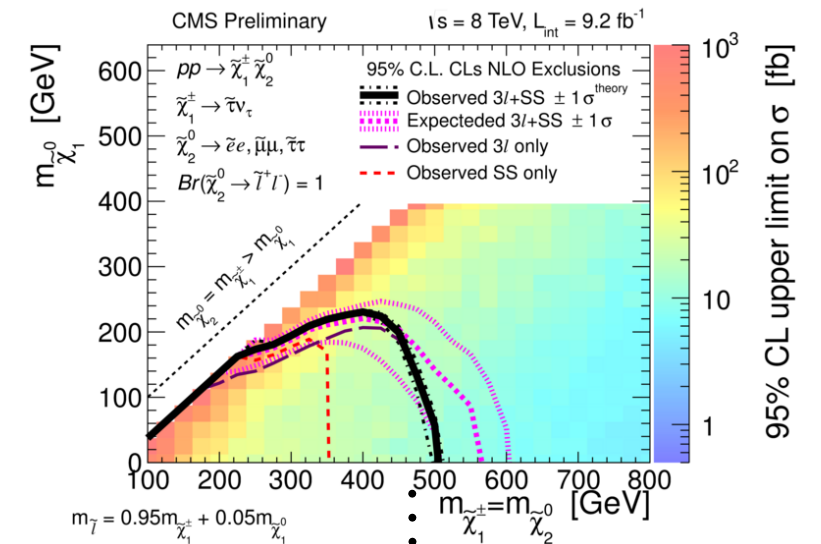
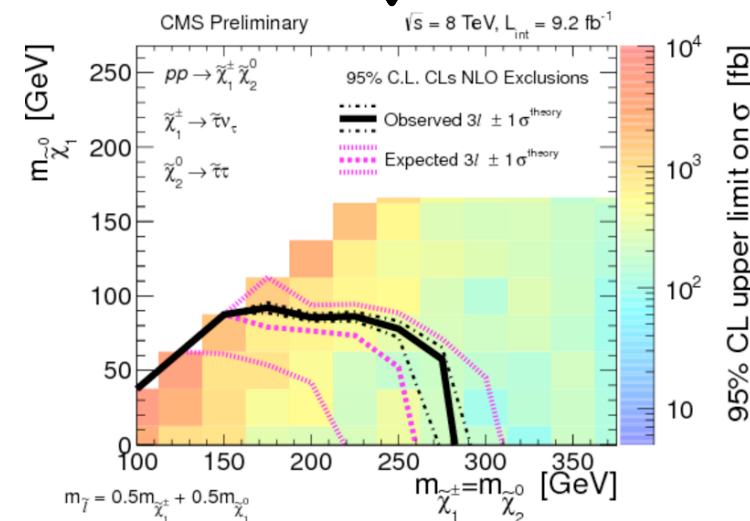
- For topologies involving an intermediate particles, three mass slices are given. We interpolate over these slices



interpolate using
k-factors



interpolate and
derive k-factors



interpolate using
k-factors

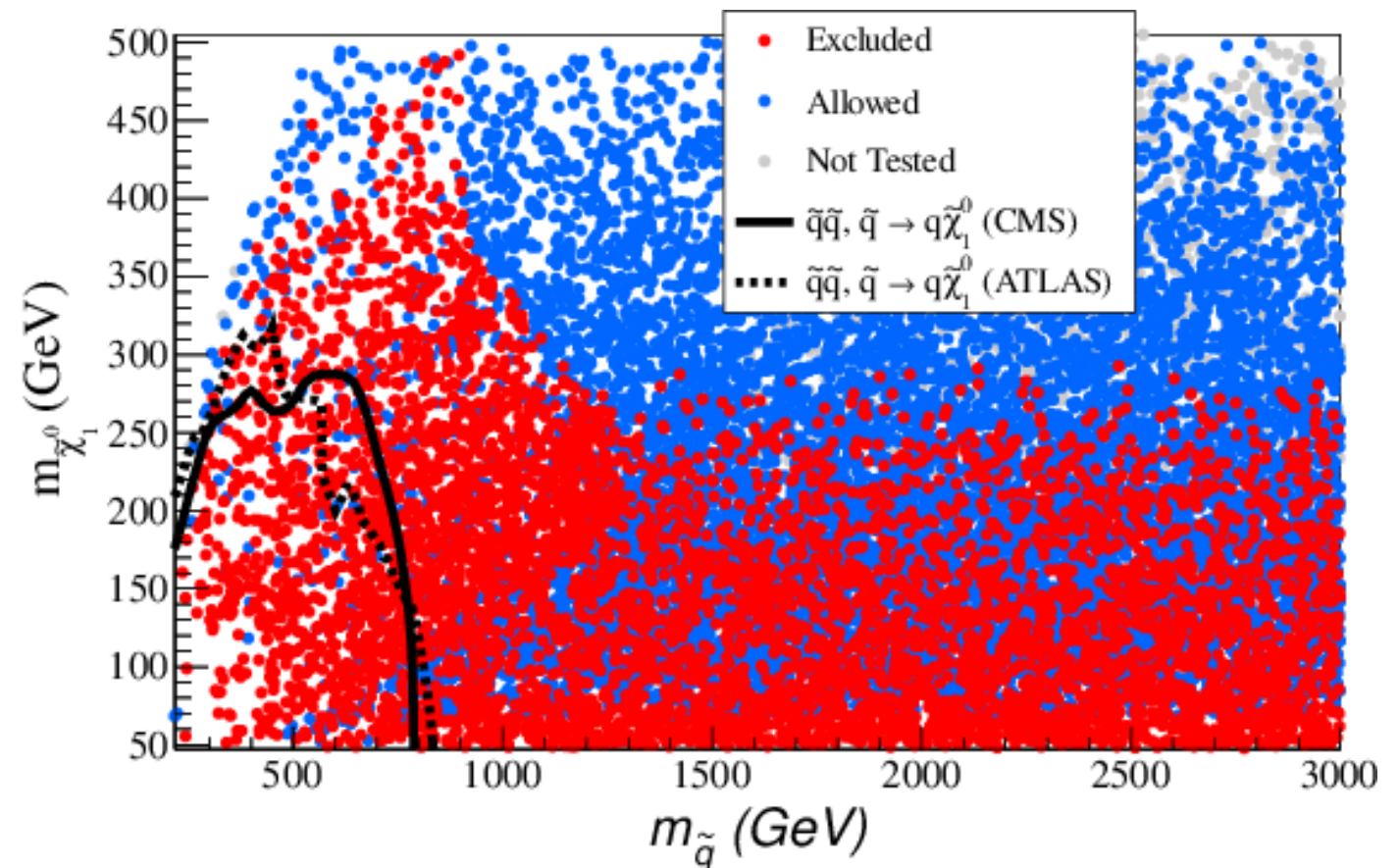
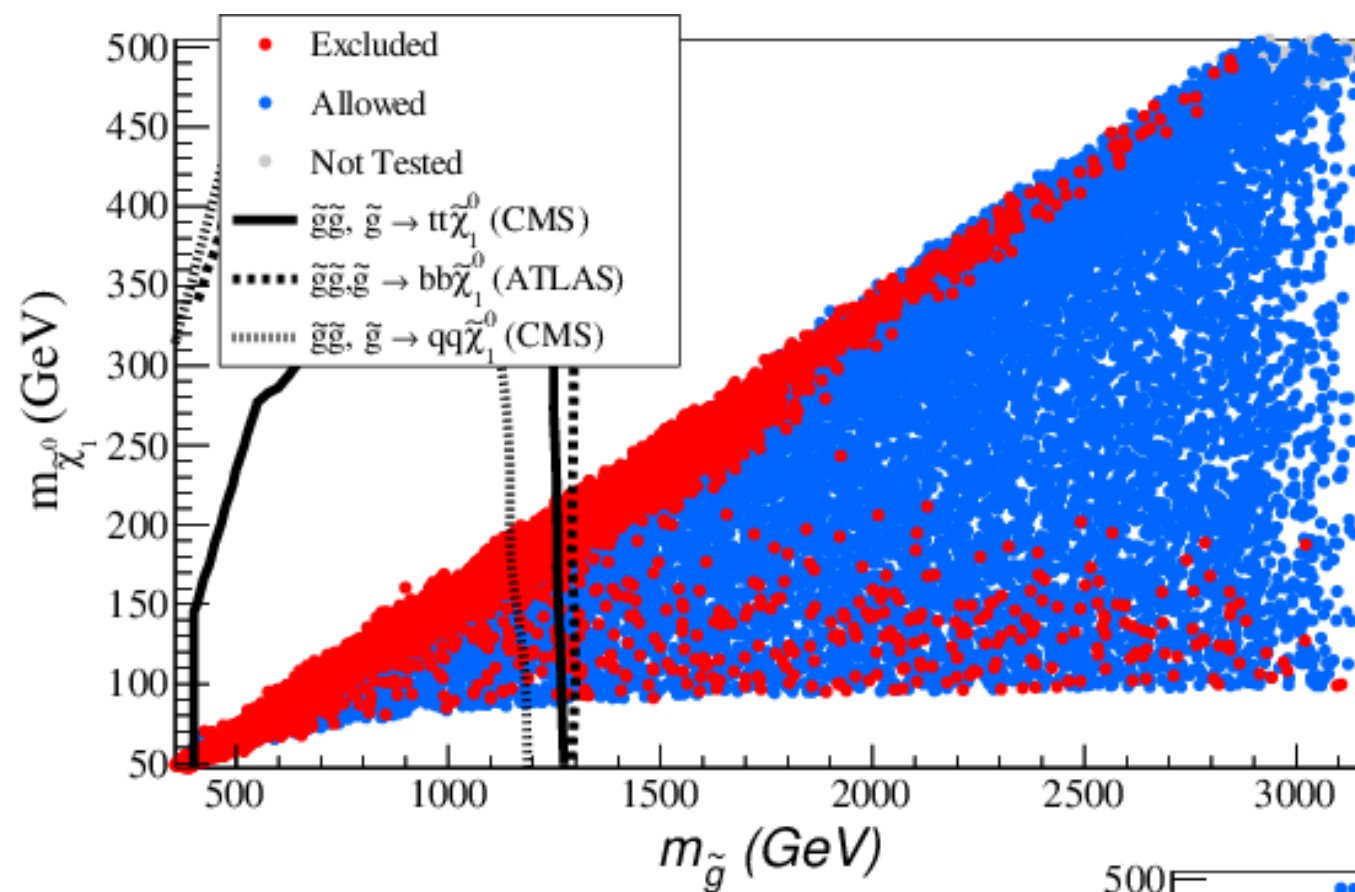


SModelS scan - strong focus

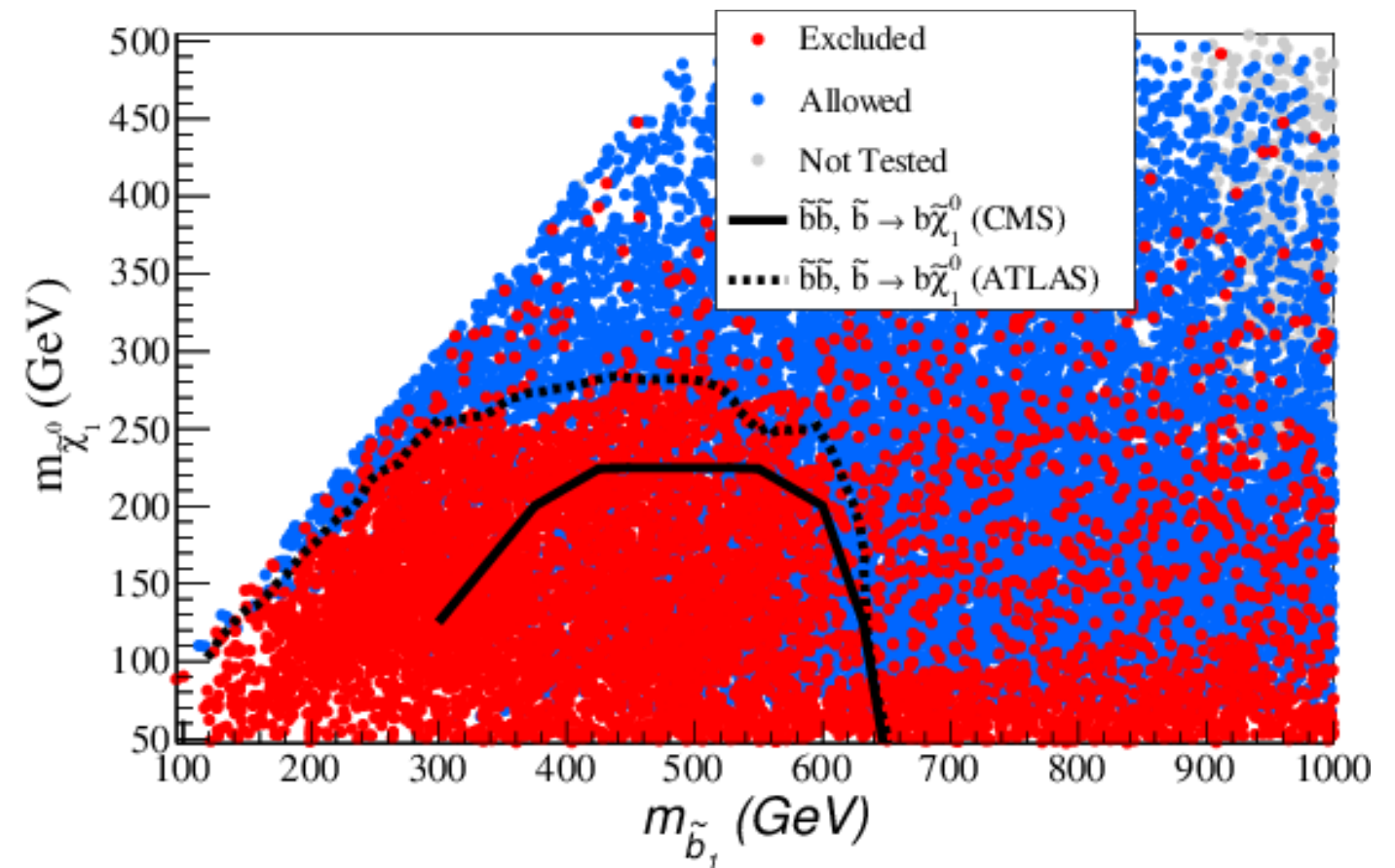
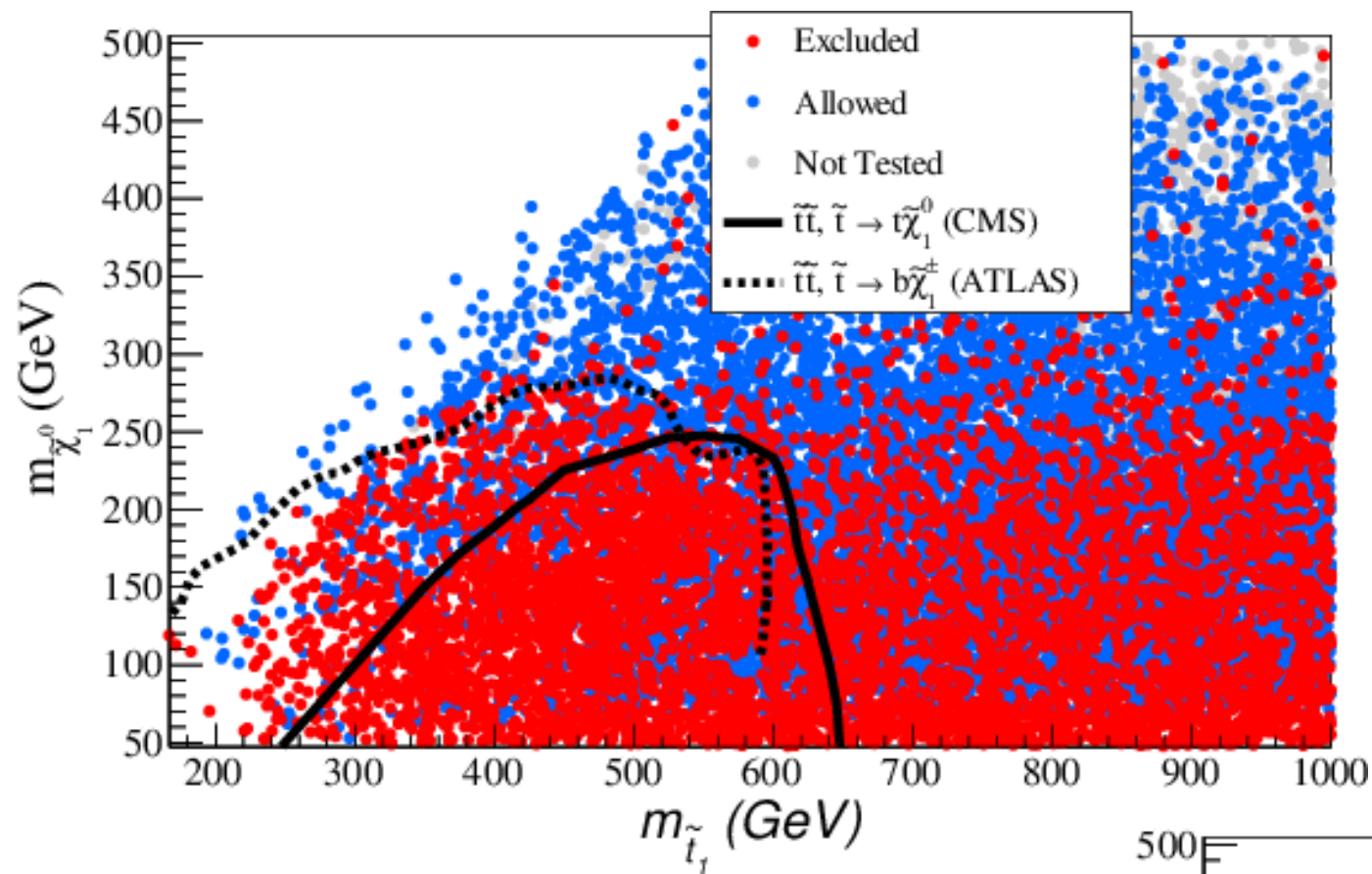
- pMSSM scan over 6 parameters
 - Gaugino masses obey GUT relation
 - Flavor constraints, invisible Z width, Higgs mass, LEP limits imposed
 - Scan ranges

M_2	μ	$\tan \beta$	$M_{\tilde{q}}$	$M_{\tilde{Q}_3}$	$M_{\tilde{D}_3}$	$M_{\tilde{U}_3}$	A_t	A_b
0.1–1	0.1–1	3–60	0.1–5	0–2	0–2	0–2	$[1,3]\max(M_{\tilde{Q}_3}, M_{\tilde{Q}_3})$	± 1

SModelS scan - strong focus



SModelS scan - strong focus



SModelS scan - strong focus

